Machine learning and anomaly detection

at the ATLAS experiment

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Overview

A (very brief!) look at ATLAS machine learning activities for non-collider physicists:

- A quick tour of the experiment (using Higgs-related examples)
- Machine learning can improve our measurements of the Higgs boson!
- Unsupervised searches and anomaly detection (aka is there something beyond the Higgs?)

CERN and the LHC

CERN: the European Organisation for Nuclear Research

The acronym doesn't make sense.

It's now also an international organisation beyond Europe.

And we mostly do high energy particle physics, rather than nuclear.



A brief history of CERN

- 1949: a concrete idea for a renewal of nuclear research and scientific excellence in Europe
- 1952: picking the right location (Geneva, Switzerland)
- 1954: start of construction and official birth of CERN
- 1957: the 600 MeV Synchro-Cyclotron starts up
- 1959: the 28 GeV Proton Synchrotron starts up
- 1971: first proton-proton collisions!
- 1976: the 400-450 GeV Super Proton Synchroton starts up
- 1983: discovery of the W^\pm and Z^0 bosons!
- 1989: the 100-200 GeV Large Electron Positron collider starts up
- 1990: the first website is up at CERN
- 2008: the 7-8-13 TeV Large Hadron Collider starts up
- 2012: discovery of the Higgs boson!

The CERN accelerator complex Complexe des accélérateurs du CERN



▶ H⁻ (hydrogen anions) ▶ p (protons) ▶ ions ▶ RIBs (Radioactive Ion Beams) ▶ n (neutrons) ▶ \overline{p} (antiprotons) ▶ e⁻ (electrons)

LHC - Large Hadron Collider // SPS - Super Proton Synchrotron // PS - Proton Synchrotron // AD - Antiproton Decelerator // CLEAR - CERN Linear Electron Accelerator for Research // AWAKE - Advanced WAKefield Experiment // ISOLDE - Isotope Separator OnLine // REX/HIE - Radioactive EXperiment/High Intensity and Energy ISOLDE // LEIR - Low Energy Ion Ring // LINAC - LINear ACcelerator // n_TOF - Neutrons Time Of Flight //

The LHC in numbers

Property	Value
Circumference	27 km
Depth	100 m
Magnet operating temperature	1.9 K (-271.3°C)
Number of magnets	9,593
Beam pressure	1.013×10 ⁻¹⁰ mbar
Nominal energy (protons)	6.5 TeV
Number of bunches per proton beam	2,808
Number of protons per bunch	1.2x10 ¹¹
Number of turns per second	11,245
Number of collisions per second	1 billion
Cost	CHF 4.3 billions
Energy consumption (CERN)	1.3 TWh/year
Energy production (Geneva)	3 TWh/year

Both colder and emptier than interstellar space!

The timeline



Of course, Covid-19 complicates things... Now expecting Run 3 to start in Q1 of 2022.

But there's still a lot to do with Run 2 data!

The biggest news so far

Observation of a new particle in the search for the Standard Model Higgs boson with the ATLAS detector at the LHC $^{\rm th}$

ATLAS Collaboration*

This paper is dedicated to the memory of our ATLAS colleagues who did not live to see the full impact and significance of their contributions to the experiment.

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ABSTRACT

A search for the Standard Model Higgs boson in proton–proton collisions with the ATLAS detector at the LHC is presented. The datasets used correspond to integrated luminosities of approximately 4.8 fb⁻¹ collected at $\sqrt{s} = 7$ TeV in 2011 and 5.8 fb⁻¹ at $\sqrt{s} = 8$ TeV in 2012. Individual searches in the channels $H \rightarrow ZZ^{(*)} \rightarrow 4\ell$, $H \rightarrow \gamma\gamma$ and $H \rightarrow WW^{(*)} \rightarrow e\nu\mu\nu$ in the 8 TeV data are combined with previously published results of searches for $H \rightarrow ZZ^{(*)} \rightarrow 4\ell$ and $H \rightarrow \gamma\gamma$ channels in the 7 TeV data and results from improved analyses of the $H \rightarrow ZZ^{(*)} \rightarrow 4\ell$ and $H \rightarrow \gamma\gamma$ channels in the 7 TeV data. Clear evidence for the production of a neutral boson with a measured mass of 126.0 ± 0.4 (stat) ± 0.4 (sys) GeV is presented. This observation, which has a significance of 5.9 standard deviations, corresponding to a background fluctuation probability of 1.7×10^{-9} , is compatible with the production and decay of the Standard Model Higgs boson.

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Clear evidence for the production of a neutral boson with a measured mass of 126 GeV [...] significance of 5.9 σ

Branching ratios of the Higgs boson



A closer look at these discoveries



A closer look at these discoveries



The ATLAS experiment

The collaboration



3,000 scientists!

The detector



The eight toroid magnets and the calorimeter.

The detector



100 m underground / 7,000 tonnes / 100 million electronic channels / 3,000 km of cables

Detector technologies at a glance

Trackers measure the momentum of charged particles:

- gaseous detectors rely on the ionisation of gas (xenon, CO₂) and the ensuing transition radiation
- solid-state detectors enable the creation of electron-hole pairs in the dense material (Silicon), arranged in strips or pixels

Inner detector (tracker)



Detector technologies at a glance

Calorimeters measure the energy deposited by incoming particles as they travel through it:

- alternate layers of dense, absorbing material (iron or lead) and active medium (liquid argon, LAr)
- the master equation is $E=E_0e^{-x/X_0}$ where X_0 is the radiation length \propto material
- electromagnetic cascade decays give precise measurements of electrons and photons, but hadronic cascade are much more complex
- in general, the energy resolution increases with energy

Calorimeters



Detector technologies at a glance

Muon spectrometers are essentially trackers too:

- use gas-based instrumentation (drift tubes, thin-gap and resistive-plate chambers) to measure the momentum of muons
- assuming that whatever made it through the ECAL+HCAL is a muon!

Muon spectrometer



Particle identification in ATLAS



Quarks hadronise and form jets: messy! → many jet constituents

Particle identification in ATLAS

Jet clustering





Particle identification in ATLAS

Jet tagging



Quiz time: what's that Higgs?





 $H
ightarrow ZZ^*
ightarrow e^+ e^- \mu^+ \mu^-$





 $H o \gamma\gamma$

Machine learning at ATLAS

Early $H ightarrow au^+ au^-$ at ATLAS



BDT (with ROOT TMVA) trained on ~10 variables, crucial for background discrimination and eventual observation of the Higgs coupling to taus! (4.5σ)

With and without Machine Learning...

Table 1 | Effect of machine learning on the discovery and study of the Higgs boson

Analysis	Years of data collection	Sensitivity without machine learning	Sensitivity with machine learning	Ratio of <i>P</i> values	Additional data required
CMS^{24} $H \rightarrow \gamma\gamma$	2011–2012	2.2 <i>σ</i> , <i>P</i> = 0.014	2.7 <i>σ</i> , <i>P</i> = 0.0035	4.0	51%
$ATLAS^{43}$ $H \rightarrow \tau^+ \tau^-$	2011–2012	2.5 <i>σ</i> , <i>P</i> = 0.0062	3.4 <i>σ</i> , <i>P</i> = 0.00034	18	85%
ATLAS ⁹⁹ VH → bb	2011–2012	1.9 σ , P = 0.029	2.5 <i>σ</i> , <i>P</i> = 0.0062	4.7	73%
$ATLAS^{41}$ VH \rightarrow bb	2015–2016	2.8 <i>σ</i> , <i>P</i> = 0.0026	3.0 <i>σ</i> , <i>P</i> = 0.00135	1.9	15%
${ m CMS^{100}}$ VH $ ightarrow$ bb	2011–2012	1.4 σ , P = 0.081	2.1 <i>σ</i> , <i>P</i> = 0.018	4.5	125%

Do more with less!

A. Radovic et al., Machine learning at the energy and intensity frontiers of particle physics

Machine learning instead of or beyond physics?

The $H \to \tau^+ \tau^-$ paper we saw previously used high-level variables: complex observables meant to represent physical quantities of interest (invariant masses, opening angles, sphericity, centrality...). These are very close to our understanding as (human!) physicists – it's how we usually approach the problem.

But is it the best way to go for a machine?



A booming field

The number of applications of machine learning to high energy physics has exploded in recent years:

- (supervised) separation tasks: backgrounds vs signal
- sampling and optimisation of calibration algorithms
- reconstruction of particles (tracking, clustering)
- fast detector simulation
- adversarial networks to remove dependency on limiting systematic uncertainties
- efficient data compression and AI for triggers
- transfer learning
- graphs and sets to deal with many-particle systems
- likelihood-free inference
- anomaly detection

Tackling variability of inputs with **Recurrent Neural Networks**
Recurrent units

Introduce the concept of states to deal with sequences: time series, chains of molecules, syntactic elements, etc.



The state h_t is a function of the previous state h_{t-1} and the input feature x_t . In a sense, the network "replicates" itself with each pass (unfolding/unrolling) — it has a notion of "memory".

Recurrent units

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One common problem with long chains of RNN cells: vanishing or exploding gradients. The latter can be fixed with e.g. gradient clipping. To tackle the former, more complex cell structures can be used: gated recurrent units (GRUs) and long short-term memory networks (LTSM) are popular examples. They allow for "forget-gates" to regularise the learning (a bit like how dropout layers are used in DNNs)

Recurrent NNs for b-tagging



Standard neural networks and BDTs are ill-suited to the problem of dealing with a variable number of jet constituents and tracks.

Instead, take a list of tracks inside a jet and feed it to an RNN. The ordering is physics-inspired: significance of the impact parameter.

Without expliciting vertexing, excellent performance is still achieved!

ATL-PHYS-PUB-2017-003

Recurrent NNs for b-tagging



Staying **on track**(s): a quick interlude

The tracking crisis



Crowd-source it!



Competition open to public on Kaggle \rightarrow can amateurs do better than the pros?

Crowd-source it!



Competition open to public on Kaggle \rightarrow can amateurs do better than the pros?

Sometimes, yes! New approaches (including deep learning) being folded in new generation of tracking algorithms at ATLAS :)





Spotify-annoy!



Introducing bucketing



Drastic reduction in complexity! Allows for supervised learning on "buckets", or unsupervised clustering approaches (we'll come back to those in a moment)...

Jet images and **Convolutional Neural Networks**

Jet images



Jet images



Convolutional layers



Convolutional layers in action







Image

Convolved Feature

Jet images





Max-pooling is a sample-based discretisation process: reduce the dimensionality of the current layer by down-sampling. This allows to focus on specific features.

Note: like the convolution, this operation can have overlaps!

Imagine you have built a CNN to identify animals. You may start with a picture of a cat:



A neat trick to make your dataset richer and less sensitive to variations is data augmentation:



By translating, rotating and smearing this (poor) cat, we generate "new" data and tell the CNN "not to cheat" by assuming all cats are perfectly centered in the <u>frame</u>, look to the right, are the right way up etc.

More on Medium!

By passing our original image through each convolutional filter, it's easy to visualise what information the CNN is actually extracting:



- The first layer retains most of the information, and starts detecting edges (through lighting perhaps?)
- The second layer makes this edge detection more explicit.
- The third layer has identified the relative position of the eyes and the nose
- After that, the CNN starts encoding features deeper and deeper, in a lowlevel representation that will be useful for classification!

We can picture the filters themselves, by looking at what image they respond most maximally to (e.g. gradient ascent from a blank picture):



More on Medium!

(back to) Jet images



Average quarks and gluons

Pre-process the data: center, rotate and normalise \rightarrow alternative to generating new data!



Average quarks and gluons

Can already observe some physics: colour flow and octet radiation! (more separation in W decays, more diffuse radiation in gluons)



A CNN quark/gluon tagger



- Input is a 16x16 pixel image of different types of constituents: truth particles, charged tracks, calorimeter clusters/towers
- The CNN learns non-linear representations of the image with the goal of discriminating between quark and gluon jet images

Performance



- Using calo+track information, the CNN performs maximally (as well as with truth information).
- And it also beats the log-likelihood model using a combination of substructure observables!
- However there are still some Monte Carlo generator biases...

Back to physics: is there something beyond the Higgs?

Supersymmetry?

ATLAS SUSY Searches* - 95% CL Lower Limits

ATLAS Preliminary $\sqrt{s} = 13 \text{ TeV}$

July 2020

	Model	Signatur	e $\int \mathcal{L} dt$ [fb	- ¹] Ma	iss limit		Reference		
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q \tilde{\chi}_1^0$	0 e, µ 2-6 jets mono-jet 1-3 jets	E_T^{miss} 139 E_T^{miss} 36.1	 <i>q</i> [10× Degen.] <i>q</i> [1×, 8× Degen.] 	0.43 0.71	1.9 m(ℓ ₁ ⁰)<400 GeV m(q̄)-m(ℓ ₁ ⁰)=5 GeV	ATLAS-CONF-2019-040 1711.03301		
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q \bar{q} \tilde{\chi}_1^0$	0 <i>e</i> , <i>µ</i> 2-6 jets	E_T^{miss} 139	ğ ĩg	Forbidden	2.35 m(k˜1)=0 GeV 1.15-1.95 m(k˜1)=1000 GeV	ATLAS-CONF-2019-040 ATLAS-CONF-2019-040		
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}W\tilde{\chi}_{1}^{0}$ $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}(\ell\ell)\tilde{\chi}_{1}^{0}$	1 e, μ 2-6 jets $ee, \mu\mu$ 2 jets 0 e, μ 7-11 jets	139 <i>E</i> ^{miss} _T 36.1 <i>E</i> ^{miss} _T 120	ğ ğ		2.2 m(k ⁰ ₁)<600 GeV 1.2 m(g)-m(x ⁰ ₁)=50 GeV 1.97 m ⁽⁰⁾ ₂ = 50 GeV	ATLAS-CONF-2020-047 1805.11381 ATLAS_CONE_2020.002		
	$gg, g \rightarrow qq w Z \Lambda_1$ $\tilde{\sigma}\tilde{\sigma} = \tilde{\sigma} \rightarrow t \tilde{T} \tilde{Y}^0,$	$SS e, \mu$ 6 jets 0-1 e, μ 3 b	$E_T = 139$ 139 $E_T^{miss} = 79.8$	S PS Sz		1.15 m(𝔅) <000 GeV 2.25 m(𝔅) <200 GeV	1909.08457 ATLAS-CONF-2018-041		
	88,8 - 10 - 1	SS e, µ 6 jets	139	° Ř		1.25 m(g)-m(t ₁ ⁰)=300 GeV	1909.08457		
3 rd gen. squarks direct production	$\tilde{b}_1 \tilde{b}_1, \tilde{b}_1 \rightarrow b \tilde{\chi}_1^0 / t \tilde{\chi}_1^x$	Multiple Multiple	36.1 139	b ₁ Forbidden	0.9 Forbidden 0.74	$m(\tilde{\chi}_{1}^{0})=300 \text{ GeV}, BR(b\tilde{\chi}_{1}^{0})=1$ $m(\tilde{\chi}_{1}^{0})=200 \text{ GeV}, m(\tilde{\chi}_{1}^{*})=300 \text{ GeV}, BR(b\tilde{\chi}_{1}^{*})=1$	1708.09266, 1711.03301 1909.08457		
	$b_1b_1, b_1 \rightarrow b\chi_2^\circ \rightarrow bh\chi_1^\circ$	$0 e, \mu$ $6 b$ 2τ $2 b$	E_T^{mins} 139 E_T^{miss} 139	b1 Forbidden b1 -	0.13-0.85	0.23-1.35 $\Delta m(\tilde{x}_{2}^{\circ}, \tilde{x}_{1}^{\circ}) = 130 \text{ GeV}, m(\tilde{x}_{1}^{\circ}) = 100 \text{ GeV}$ $\Delta m(\tilde{x}_{2}^{\circ}, \tilde{x}_{1}^{\circ}) = 130 \text{ GeV}, m(\tilde{x}_{1}^{\circ}) = 0 \text{ GeV}$	1908.03122 ATLAS-CONF-2020-031		
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow t \tilde{\mathcal{X}}_1^\circ$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow W b \tilde{\tilde{\mathcal{X}}}_1^0$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow W b \tilde{\tilde{\mathcal{X}}}_1^\circ$	$1 e, \mu = 2 \text{ interval}$ $1 e, \mu = 3 \text{ jets/} 1 b$	E_T^{miss} 139 E_T^{miss} 139 E_T^{miss} 26.1	\tilde{t}_1 \tilde{t}_1 $\tilde{\tau}$	0.44-0.59	1.25 $m(\tilde{\chi}_1^0)=1 \text{ GeV}$ $m(\tilde{\chi}_1^0)=400 \text{ GeV}$ 1.16 $m(\tilde{\chi}_1^0)=400 \text{ GeV}$	ATLAS-CONF-2020-003, 2004.14060 ATLAS-CONF-2019-017 1803-10178		
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{t}_1 \tilde{b} v, \tilde{t}_1 \rightarrow \tilde{t} \tilde{t}_0$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{c} \tilde{\chi}_1^0 / \tilde{c} \tilde{c}, \tilde{c} \rightarrow \tilde{c} \tilde{\chi}_1^0$	$0 e, \mu$ $2 c$ $0 e, \mu$ mono-jet	E_T 36.1 E_T^{miss} 36.1 E_T^{miss} 36.1	\tilde{r}_1 \tilde{c} \tilde{t}_1 \tilde{t}_1	0.85 0.46 0.43	m(t)/>-00000 m(t)/=0 GeV m(t)/=50 GeV m(t)/=50 GeV	1805.01649 1805.01649 1711.03301		
	$ \begin{split} \tilde{t}_1 \tilde{t}_1, \tilde{t}_1 {\rightarrow} t \tilde{\chi}_2^0, \tilde{\chi}_2^0 {\rightarrow} Z/h \tilde{\chi}_1^0 \\ \tilde{t}_2 \tilde{t}_2, \tilde{t}_2 {\rightarrow} \tilde{t}_1 + Z \end{split} $	1-2 e, μ 1-4 b 3 e, μ 1 b	E_T^{miss} 139 E_T^{miss} 139		0.067- Forbidden 0.86	-1.18 $m(\tilde{t}_2^0)$ =500 GeV $m(\tilde{t}_1^0)$ =360 GeV, $m(\tilde{t}_1)$ = 40 GeV	SUSY-2018-09 SUSY-2018-09		
EW direct	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0$ via WZ	$\begin{array}{ll} 3 \ e, \mu \\ e e, \mu \mu \end{array} \ge 1 \ { m jet}$	$\begin{array}{ccc} E_T^{\mathrm{miss}} & 139 \\ E_T^{\mathrm{miss}} & 139 \end{array}$	$ \begin{array}{ccc} \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} & \ \tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} & \ 0.205 \end{array} $	0.64	$m(\tilde{\chi}_1^0)=0$ $m(\tilde{\chi}_1^\pm)\cdot m(\tilde{\chi}_1^0)=5$ GeV	ATLAS-CONF-2020-015 1911.12606		
	$ ilde{\chi}_1^{\pm} ilde{\chi}_1^{\mp}$ via WW $ ilde{\chi}_1^{\pm} ilde{\chi}_2^0$ via Wh	2 e, μ 0-1 e, μ 2 b/2 γ	$\begin{array}{ll} E_T^{\rm miss} & {\rm 139} \\ E_T^{\rm miss} & {\rm 139} \end{array}$	$ ilde{x}_1^{\pm}$ $ ilde{x}_1^{\pm}/ ilde{x}_2^{0}$ Forbidden	0.42	$m(\tilde{\chi}_{1}^{0})=0$ $m(\tilde{\chi}_{1}^{0})=70~{ m GeV}$	1908.08215 2004.10894, 1909.09226		
	$\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp}$ via $\tilde{\ell}_{L}/\tilde{\nu}$ $\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau \tilde{\chi}_{1}^{0}$	2 e, μ 2 τ	E_T^{miss} 139 E_T^{miss} 139	$\tilde{\chi}_{1}^{\pm}$ $\tilde{\tau}$ [$\tilde{\tau}_{L}, \tilde{\tau}_{R,L}$] 0.16-0.3	0.12-0.39	$\begin{split} \mathbf{m}(\tilde{\ell},\tilde{\nu}) = 0.5(\mathbf{m}(\tilde{\chi}_{1}^{\pm}) + \mathbf{m}(\tilde{\chi}_{1}^{0})) \\ \mathbf{m}(\tilde{\chi}_{1}^{0}) = 0 \end{split}$	1908.08215 1911.06660		
	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_1^{\prime\prime}$	$2 e, \mu$ 0 jets $ee, \mu\mu \ge 1$ jet	E_T^{miss} 139 E_T^{miss} 139	<i>ι̃</i> <i>ℓ̃</i> 0.256	0.7	$m(\tilde{\ell})^{-1}=0$ $m(\tilde{\ell})-m(\tilde{\ell}_{1}^{0})=10 \text{ GeV}$	1908.08215 1911.12606		
	$HH, H \rightarrow hG/ZG$	$\begin{array}{ccc} 0 \ e, \mu & \geq 3 \ b \\ 4 \ e, \mu & 0 \ \text{jets} \end{array}$	E_T^{miss} 36.1 E_T^{miss} 139	<i>H</i> 0.13-0.23 <i>H</i>	0.29-0.88 0.55	$\begin{array}{c} BR(\widetilde{k}_1^\circ) \to h\widetilde{G})=1\\ BR(\widetilde{k}_1^\circ) \to Z\widetilde{G})=1 \end{array}$	1806.04030 ATLAS-CONF-2020-040		
Long-lived particles	$\text{Direct}\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk 1 jet	E_T^{miss} 36.1	$ \tilde{\chi}_{1}^{\pm} = 0.15 $	0.46	Pure Wino Pure higgsino	1712.02118 ATL-PHYS-PUB-2017-019		
	Stable \tilde{g} R-hadron Metastable \tilde{g} R-hadron, $\tilde{g} \rightarrow qq \tilde{\chi}_1^0$	Multiple Multiple	36.1 36.1	\tilde{g} \tilde{g} [$\tau(\tilde{g})$ =10 ns, 0.2 ns]		2.0 2.05 2.4 m(𝑋10)=100 GeV	1902.01636,1808.04095 1710.04901,1808.04095		
	$ \begin{aligned} \tilde{\chi}_{1}^{\pm} \tilde{\chi}_{1}^{\mp} / \tilde{\chi}_{1}^{0} , \tilde{\chi}_{1}^{\pm} \rightarrow Z\ell \rightarrow \ell\ell\ell \\ LFV \ pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e\mu/e\tau/\mu\tau \end{aligned} $	3 е, µ еµ,ет,µт	139 3.2	$\tilde{X}_{1}^{\mp}/\tilde{X}_{1}^{0}$ [BR($Z\tau$)=1, BR(Ze)=1] \tilde{v}_{τ}	0.625 1.0	15 Pure Wino 1.9 λ' ₃₁₁ =0.11, λ _{132/133/233} =0.07	ATLAS-CONF-2020-009 1607.08079		
	$\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\mp}/\tilde{\chi}_{2}^{0} \rightarrow WW/Z\ell\ell\ell\ell\nu\nu$ $\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} \rightarrow aa\tilde{\chi}_{1}^{0}\tilde{\chi}_{1}^{0} \rightarrow aaa$	4 <i>e</i> , μ 0 jets 4-5 large- <i>R</i> je	E_T^{miss} 36.1 ts 36.1	$\tilde{\chi}_{1}^{\pm}/\tilde{\chi}_{2}^{0} [\lambda_{i33} \neq 0, \lambda_{12k} \neq 0]$ $\tilde{\chi} [m(\tilde{\chi}_{1}^{0})=200 \text{ GeV}, 1100 \text{ GeV}]$	0.82	1.33 m($\tilde{\chi}_1^0$)=100 GeV 1.3 1.9 Large $\lambda_{112}^{\prime\prime}$	1804.03602 1804.03568		
Vdb	T L VV V V	Multiple Multiple	36.1	$\tilde{g} = [\lambda''_{112} = 2e-4, 2e-5]$ $\tilde{\ell} = [\lambda''_{112} = 2e-4, 1e-2]$	0.55 1.0	2.0 $m(\tilde{\chi}_1^0)=200 \text{ GeV, bio-like}$	ATLAS-CONF-2018-003		
4	$\vec{t}, \vec{t} \rightarrow \vec{k}_1, \vec{x}_1 \rightarrow \vec{t} \vec{b} \vec{s}$ $\vec{t}, \vec{t} \rightarrow \vec{b} \vec{x}_1, \vec{x}_1 \rightarrow \vec{b} \vec{b} \vec{s}$	$\geq 4b$	139	\tilde{t}	Forbidden 0.95	m(X [*] ₁)=200 GeV, billo mks m(X [*] ₁)=500 GeV	ATLAS-CONF-2020-016		
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow 0 \tilde{s}$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow q \ell$	$2 e, \mu$ $2 b$ 1μ DV	36.1 136	$ \begin{array}{l} \tilde{t}_1 & [qq, vs] \\ \tilde{t}_1 & \\ \tilde{t}_1 & [1e\text{-}10 < \lambda'_{23k} < 1e\text{-}8, 3e\text{-}10 < \lambda'_{23} \end{array} $	<pre></pre>	0.4-1.45 BR($\tilde{t}_1 \rightarrow be/b\mu$)>20% 1.6 BR($\tilde{t}_1 \rightarrow q\mu$)=100%, $\cos\theta_i$ =1	1710.05544 2003.11956		
*Only a selection of the available mass limits on new states or 10^{-1} 1 Mass scale [TeV]									

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

Exotic particles?

ATLAS Exotics Searches* - 95% CL Upper Exclusion Limits

 ℓ, γ Jets; $\mathsf{E}^{\text{miss}}_{-}$ (\mathcal{L} dt[fb⁻¹]

Status: May 2020 Model ATLAS Preliminary

 $\int \mathcal{L} dt = (3.2 - 139) \text{ fb}^{-1}$

 $\sqrt{s} = 8$, 13 TeV **Beference**

		J - 1			
Extra dimensions	$\begin{array}{llllllllllllllllllllllllllllllllllll$	36.1 36.7 37.0 3.2 3.6 36.7 36.1 139 36.1 36.1	Mp 7.7 TeV Ms 8.6 TeV Mm 8.9 TeV Mm 8.2 TeV Mm 9.55 TeV Grace 2.3 TeV Grace 2.3 TeV Grace 2.3 TeV Grace 3.8 TeV KK mass 2.0 TeV Grace 3.8 TeV	$ \begin{split} n &= 2 \\ n &= 3 \text{ HLZ NLO} \\ n &= 6 \\ m &= 6, M_D = 3 \text{ TeV, rot BH} \\ n &= 6, M_D = 3 \text{ TeV, rot BH} \\ k/\overline{M}_{PI} = 0.1 \\ k/\overline{M}_{PI} = 1.0 \\ k/\overline{M}_{PI} = 1.0 \\ f/m &= 15\% \\ \hline \Gamma(m = 15\% \\ \Gamma(m : 1.5\% \\ rot (1.1), \mathscr{B}(A^{(1.1)} \to tt) = 1 \end{split} $	1711.03301 1707.04147 1703.09127 1606.02265 1512.02586 1707.04147 1808.02380 2004.14636 1804.10823 1803.09678
Gauge bosons	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	139 36.1 36.1 139 36.1 139 36.1 139 36.1 139 36.1 80	Z' mass 5.1 TeV Z' mass 2.42 TeV Z' mass 2.1 TeV Z' mass 2.1 TeV Z' mass 3.1 TeV W' mass 6.0 TeV W' mass 3.7 TeV W' mass 3.3 TeV V' mass 3.8 TeV V' mass 2.93 TeV Ww mass 3.2 TeV We mass 3.25 TeV We, mass 5.0 TeV	$\Gamma/m = 1.2\%$ $g_V = 3$ $g_V = 3$ $g_V = 3$ $g_V = 3$ $m(N_R) = 0.5 \text{ TeV}, g_L = g_R$	1903.06248 1709.07242 1805.09299 2005.05138 1906.05609 2004.14636 1906.08589 1712.06518 CERN-EP-2020-073 1807.10473 1807.10473
10	Cl $qqqq$ -2 jCl $\ell \ell qq$ 2 e, μ -Cl $\ell tttt$ $\geq 1 e, \mu$ >1 b, $\geq 1 j$	37.0 139 36.1	Λ Λ Λ 2.57 TeV	21.8 TeV η_{LL}^- 35.8 TeV η_{LL}^- $ C_{4t} = 4\pi$	1703.09127 CERN-EP-2020-066 1811.02305
MQ	$ \begin{array}{llllllllllllllllllllllllllllllllllll$	36.1 36.1 3.2 36.1	m _{med} 1.55 TeV m _{med} 1.67 TeV M. 700 GeV m _e 3.4 TeV	$\begin{array}{l} g_q{=}0.25, g_{\chi}{=}1.0, \ m(\chi) = 1 \ {\rm GeV} \\ g{=}1.0, \ m(\chi) = 1 \ {\rm GeV} \\ m(\chi) < 150 \ {\rm GeV} \\ y = 0.4, \ \lambda = 0.2, \ m(\chi) = 10 \ {\rm GeV} \end{array}$	1711.03301 1711.03301 1608.02372 1812.09743
TO	$ \begin{array}{llllllllllllllllllllllllllllllllllll$	36.1 36.1 36.1 36.1	LQ mass 1.4 TeV LQ mass 1.56 TeV LQ ^a mass 1.03 TeV LQ ^a mass 970 GeV	$\begin{split} \beta &= 1 \\ \beta &= 1 \\ \mathcal{B}(\mathrm{LQ}_3^{\prime\prime} \to b\tau) &= 1 \\ \mathcal{B}(\mathrm{LQ}_3^{\prime\prime} \to t\tau) &= 0 \end{split}$	1902.00377 1902.00377 1902.08103 1902.08103
Heavy	$ \begin{array}{c} VLQ\; TT \rightarrow Ht/Zt/Wb + X \\ multi-channel \\ multi-channel$	36.1 36.1 36.1 36.1 79.8 20.3	T mass 1.37 TeV B mass 1.34 TeV T s ₁ mass 1.64 TeV Y mass 1.85 TeV B mass 1.21 TeV	$ \begin{array}{l} & \mathrm{SU(2) \ doublet} \\ & \mathrm{SU(2) \ doublet} \\ & \mathcal{B}(T_{5/3} \rightarrow Wt) = 1, \ c(T_{5/3}Wt) = 1 \\ & \mathcal{B}(Y \rightarrow Wb) = 1, \ c_R(Wb) = 1 \\ & \kappa_B = 0.5 \end{array} $	1808.02343 1808.02343 1807.11883 1812.07343 ATLAS-CONF-2018-024 1509.04261
Excited	$ \begin{array}{c cccc} \text{Excited quark } q^* \rightarrow qg & - & 2j & - \\ \text{Excited quark } q^* \rightarrow q\gamma & 1\gamma & 1j & - \\ \text{Excited quark } b^* \rightarrow bg & - & 1b, 1j & - \\ \text{Excited lepton } \ell^* & 3e, \mu, \tau & - & - \\ \text{Excited lepton } \nu^* & 3e, \mu, \tau & - & - \\ \end{array} $	139 36.7 36.1 20.3 20.3	q* mass 6.7 TeV g* mass 5.3 TeV b* mass 2.6 TeV '' mass 3.0 TeV v* mass 1.6 TeV	only u^* and d^* , $\Lambda = m(q^*)$ only u^* and d^* , $\Lambda = m(q^*)$ $\Lambda = 3.0$ TeV $\Lambda = 1.6$ TeV	1910.08447 1709.10440 1805.09299 1411.2921 1411.2921
Other	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	79.8 36.1 36.1 20.3 36.1 34.4	N ^e mass 560 GeV 3.2 TeV Nr, mass 870 GeV 3.2 TeV H ^{±+} mass 870 GeV 1.2 TeV Imonpole mass 2.37 TeV 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.1 1.1	$\begin{split} m(W_R) &= 4.1 \text{ TeV}, g_L = g_R \\ \text{DY production} \\ DY production, \mathcal{B}(H_L^{\pm\pm} \to (\tau) = 1 \\ \text{DY production}, q &= 5e \\ \text{DY production}, g &= 1g_D, \text{ spin } 1/2 \\ \end{split}$	ATLAS-CONF-2018-020 1809.11105 1710.09748 1411.2921 1812.03673 1905.10130
	Tuli uata			Mass scale [TeV]	

Limit

*Only a selection of the available mass limits on new states or phenomena is shown.

†Small-radius (large-radius) jets are denoted by the letter j (J).

Resonances in bump hunts?



What is **anomaly detection**?

The basic idea

"Finding patterns that do not conform to expected behaviour." (C. Nellist, 2020)



Defined phase-space

The basic idea

"Finding patterns that do not conform to expected behaviour." (C. Nellist, 2020)

- Most commonly used in time series: e.g. fraud detection
- Assume ignorance of the type of anomaly / "catch-all" strategy: unsupervised learning!
- In particle physics: model-independent BSM searches

The basic idea

"Finding patterns that do not conform to expected behaviour." (C. Nellist, 2020)

- Most commonly used in time series: e.g. fraud detection
- Assume ignorance of the type of anomaly / "catch-all" strategy: unsupervised learning!
- In particle physics: model-independent BSM searches

Train your algorithm to form an internal representation of your training data (actual data or MC), apply to unseen events and use some quality criterion: how "new" is this unseen data?

Note: as in precision measurements, an anomaly (or excess/deviation) *doesn't necessarily translate* to BSM physics! It might be BSM, it might be poor detector performance, high-order QCD correction, extreme region of phase-space etc.
Bringing in **neural networks**: **CWoLa** hunting

Classifying WithOut LAbels

Beautiful (proven) theorem:

A classifier trained to optimally discriminate mixed sample 1 from mixed sample 2 is also optimal for discriminating S from B



... so long as B and S are drawn from the same distribution in samples 1 & 2, and stats are large enough! (fair)

CWoLa hunting



- 1. Scan range of interest, defining sidebands and signal regions
- 2. Train networks on background and signal, as defined in 1
- 3. Compute p-value and signal significance in signal region

arXiv.1902.02634

Keep rolling!

CWoLa hunting



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Keep rolling!

arXiv:1805.02664 arXiv:1902.02634

CWoLa hunting



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Keep rolling!

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4.

Compression and reconstruction: **AE**s and **VAE**s

A neural network with a squeeze



An autoencoder is a DNN with 3 defining features:

- 1. the size of the output is the same as the size of the input
- 2. the loss is measured with respect to the input (and not some target!)
- 3. there is a **bottleneck**

A neural network with a squeeze



An autoencoder is a DNN with 3 defining features:

- 1. the size of the output is the same as the size of the input
- 2. the loss is measured with respect to the input (and not some target!)
- 3. there is a bottleneck

These conditions necessarily lead to a latent space (an internal representation of the input data), with the means to translate to/from it.

A neural network with a squeeze



An autoencoder therefore offers 3 main functionalities:

- 1. it can be used to de-noise data (noise not essential to latent representation!)
- 2. the latent space might offer new opportunities for discrimination
- 3. a large reconstruction error signals an anomaly

The re-parameterisation trick



loss = $||x - \hat{x}||^2 + KL[N(\mu_x, \sigma_x), N(0, I)] = ||x - d(z)||^2 + KL[N(\mu_x, \sigma_x), N(0, I)]$

A variational autoencoder connects the decoder to the encoder via a sampling layer: the Kullback-Leibler divergence (KLd) term in the loss enforces structure in the latent space.

The network will eventually learn the most efficient balance between the reconstruction loss and the sampling loss.

The re-parameterisation trick



loss = $||x - \hat{x}||^2 + KL[N(\mu_x, \sigma_x), N(0, I)] = ||x - d(z)||^2 + KL[N(\mu_x, \sigma_x), N(0, I)]$

If we consider only the MSE loss, we lose the constraint on $\mathcal{N}(\mu, \sigma) \sim \mathcal{N}(0, 1)$ and the VAE is allowed to "cheat" by clustering events arbitrarily far apart.

If we consider only the KLd loss, we force the structure of the latent space to be $\mathcal{N}(0,1)$: we have generation without modelling!

The re-parameterisation trick



loss = $|| \mathbf{x} - \mathbf{x}' ||^2 + KL[N(\mu_x, \sigma_x), N(0, I)] = || \mathbf{x} - d(\mathbf{z}) ||^2 + KL[N(\mu_x, \sigma_x), N(0, I)]$

If we've found a balance, we have a VAE that:

1. has a continuous latent space, allowing generation of new events

2. has a structured latent space, allowing clustering and discrimination

3. has a large reconstruction error on exotic data, allowing anomaly detection

Spot the differences



VAE setup: compress 27 input jet variables down to 14 in the latent space.



Work ongoing to make sense of the latent space: here see strong correlation between the residuals on m and $p_{\rm T}$.



Un-labeled representation of the latent space.



Labeled representation of the latent space. How meaningful are these clusters?



Background is in orange, signal in blue – there is discrimination potential in the latent space!

Dark Machines

An international group of researchers *(not just LHC experimentalists!) using machine learning to go after Dark Matter in various forms:

- unsupervised collider searches (anomaly detection)
- high-dimensional sampling
- generative models
- strong lensing
- ... [your project here]

Expect the first community white paper on event-level anomaly detection very soon! (including many more interesting models and approaches I didn't have time to cover here)

darkmachines.org

Thank you.

Questions?

References

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