

# **The AION Project**

# A UK Atom Interferometer Observatory and Network

to explore Ultra-Light Dark Matter and Mid-Frequency Gravitational Waves. O. Buchmueller, Imperial College London on behalf of the AION Collaboration

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Project executed in national partnership with UK National Quantum Technology Hub in Sensors and Timing, Birmingham, UK, and international partnership with The MAGIS Collaboration and The Fermi National Laboratory, US

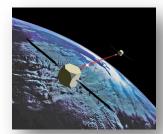
# **The AION Programme consists of 4 Stages**

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- **Stage 1:** to build and commission the 10 m detector, develop existing technology and the infrastructure for the 100 m. L~10m
- □ Stage 2: to build, commission and exploit the 100 m detector and carry out a design study for the km-scale detector.
  - > AION was selected in 2018 by STFC as a high-priority medium-scale project.
  - AION will work in equal partnership with MAGIS in the US to form a "LIGO/Virgo-style" network & collaboration, providing a pathway for UK leadership.
- Stage 1 is now funded with about £9.6M by the QTFP Programme and other sources and Stage 2 could be placed at national facility in Boulby or Daresbury (UK), possibly also at CERN (France/Switzerland).
- **Stage 3:** to build a kilometre-scale terrestrial detector.
- **Stage 4**: long-term objective a pair of satellite detectors (thousands of kilometres scale) [AEDGE proposal to ESA Voyage2050 call]
  - > AION has established science leadership in AEDGE, bringing together collaborators from European and Chinese groups (e.g. MIGA, MAGIA, ELGAR, ZAIGA).

Stage 3 and 4 will require funding on international level (ESA, EU, etc) and AION has already started to build the foundation for it.





SOURCE

ATOM SOURCE

ATOM SOURCE

L~100m

 $I \sim 1 \text{km}$ 

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### **Progressive Expansion of Quantum Detector in Length**

#### Stage Expansion of Detector Length\*\*

#### Stage 1: 2021 to 2025

10m detector protype to be build in Oxford (see backup).

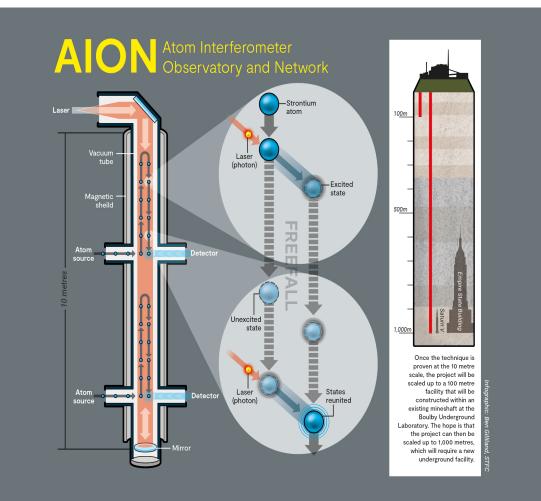
#### Stage 2: 2023 to 2030

100m detector: Site options are Boulby or Daresbury (UK), possibly also at CERN (France/Switzerland).

#### Stage 3: 2030 onwards

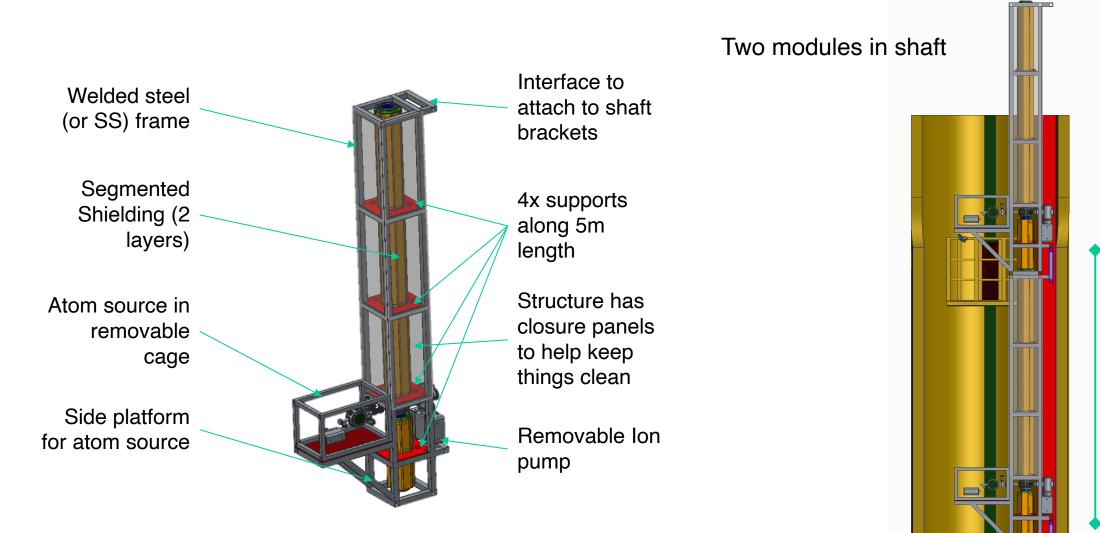
- Km-scale detector: Site options are Boulby (so far).
- > Submitted for IAG Future Projects.
- \*\* Sensitivity scales with length of detector

#### Illustration of different Stages





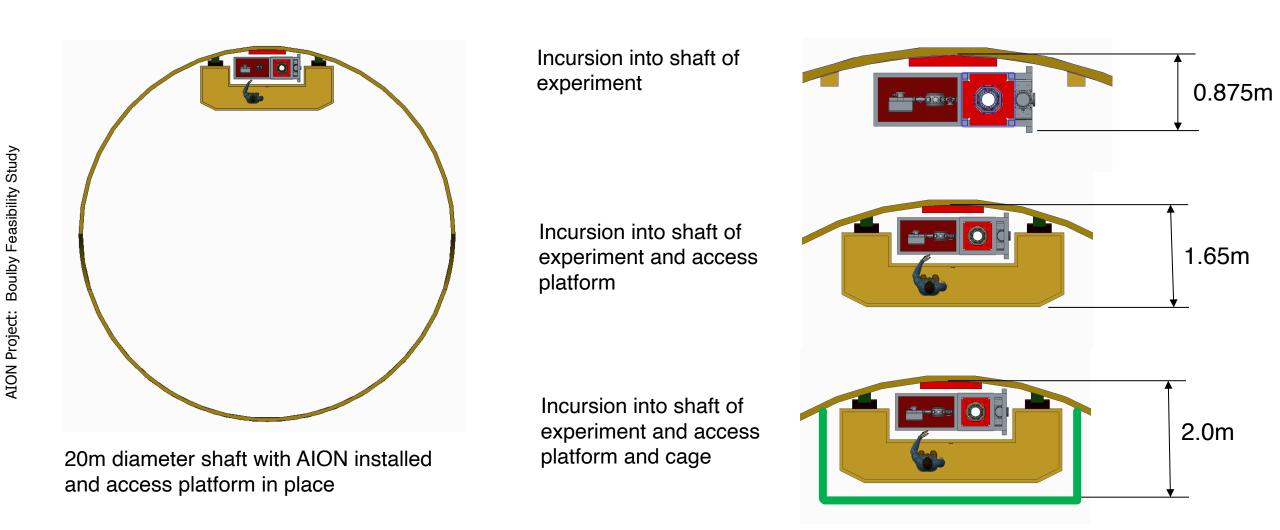




1 x module – 5m AI CN



### Larger Shafts Design (relevant for Boulby and CERN)



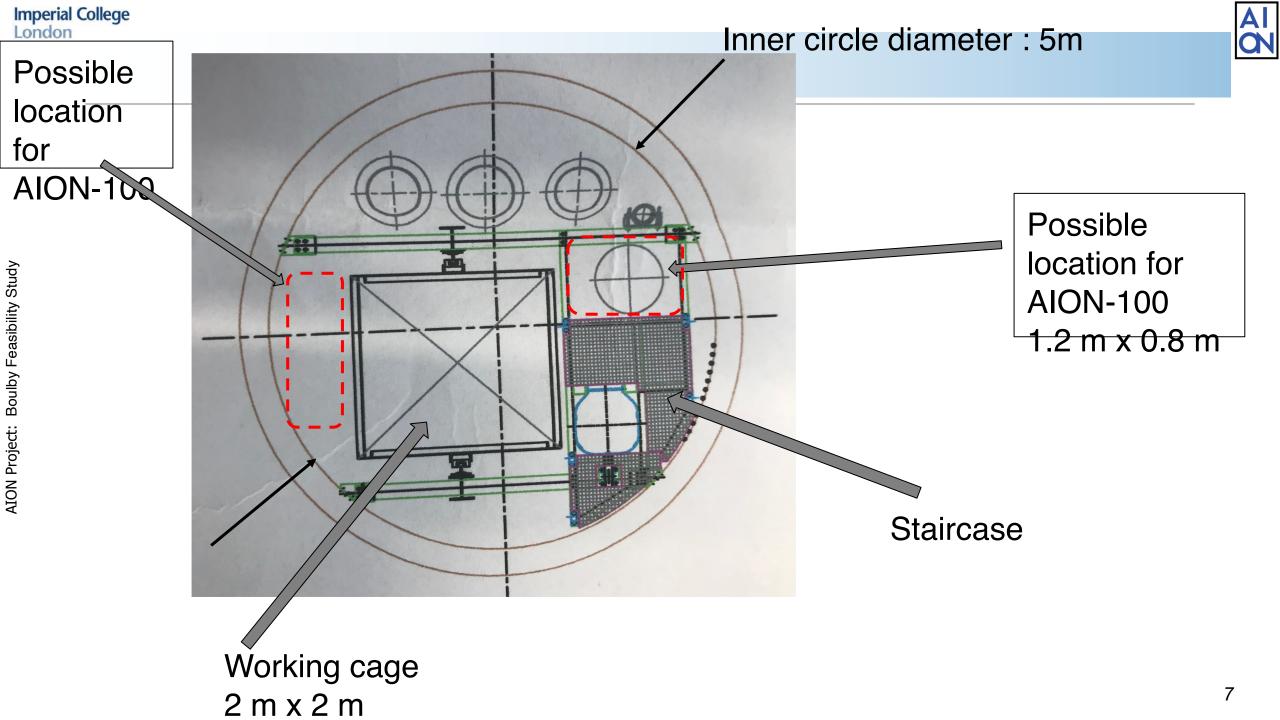
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### Location of a possible short shaft at Boulby for AION-100





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# Summary as submitted in the 2020 Infrastructure Call to IAG & UKRI

- Stage 1 (10m): which has received initial funding by an STFC / ESPRC grant (total £9.6M), will provide proof-of-principle of the basic technology and examine the scalability from lab-based to purpose-built infrastructure. A 10m interferometer will be constructed in the low-vibration Beecroft building in Oxford, which also includes a high-specification laser laboratory. The performance and scalability of this device will be fully characterized.
- Stage 2 (100m): Construction is planned to commence in late 2023, with a goal of device operation by 2025 and observation and exploitation of GW signals through to the end of the decade. Two potential UK sites for a 100m device are under study, at Boulby and Daresbury, each taking advantage of existing unique facilities within national laboratory campuses. This stage of the project will address fundamental engineering challenges for the construction of a large detector 'in the field'.
- Stage 3 (1km): Construction could start in 2025/6 and be completed in 2029/30, with a target of
  reaching the ultimate terrestrial sensitivity for GW / DM observation in the early 2030s. This stage
  would represent a new national / international infrastructure, complementary to existing UK 'big
  science' facilities. Boulby laboratory is under study as a potential UK site for the detector,
  building on already-announced investment in other science activities.
  - i. Building the km-scale detector in an existing mine shaft (e.g. in one of the two 1.1km vertical shafts at the Boulby mine);
    - ii. Building the km-scale detector in a new km-deep vertical shaft excavated at a facility like Boulby.



### Summary



Boulby could be an ideal location to host Stage 2 and Stage 3 of the AION project.

- This would make the UK a centre of future international activities in the area of mid- frequency band studies of dark matter and gravitational physics, capable of hosting world-leading experiments at UK sites, e.g., STFC's Boulby facilities The long-term objective of the programme is to give the UK international leadership of a future km-scale atom interferometry observatory.
- This km-scale terrestrial quantum detector for mid-frequency gravitational wave physics and refined sensitivity for ultra-light dark matter physics will open a new dimension of multi-signal and multimessenger physics in the mid-frequency band (and beyond), capable of networking with other km devices (located, e.g., in the US or China), as well as with experiments in the high- and lowfrequency bands.
- This stage will represent a new national / international research infrastructure, complementary to existing UK 'big science' facilities. The STFC Boulby underground laboratory is under study as a potential UK site for the detector, building on already-announced investment in other science activities, and on a proposed expansion of the laboratory.



# **BACKUP MATERIAL**



# **10M DETECTOR IN OXFORD**



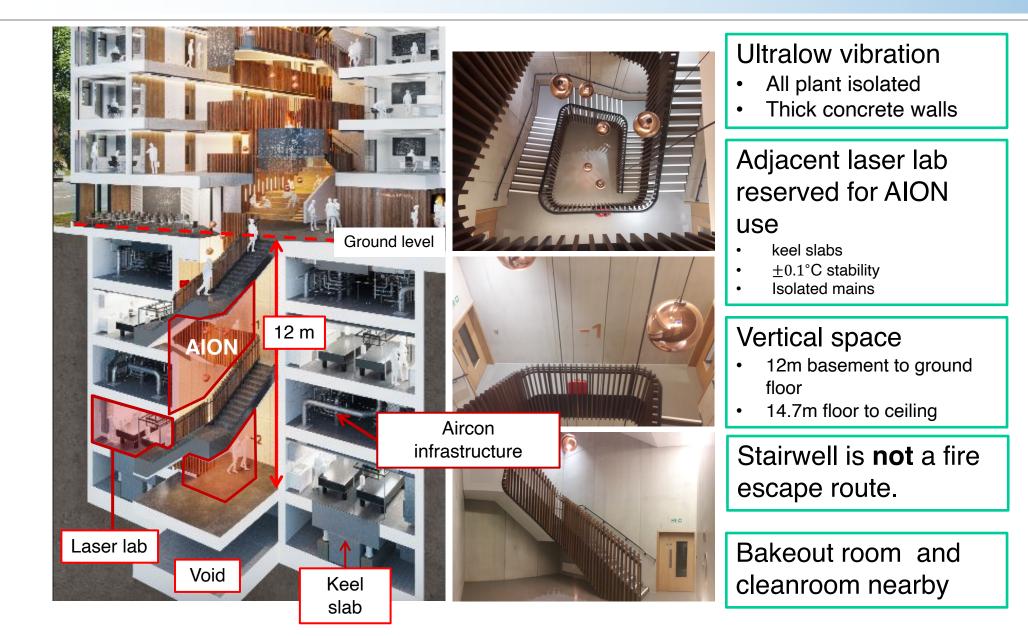
# Beecroft building, Oxford Physics

The Beecroft in Oxford is the proposed site, with a backup at RAL (MICE Hall) in case show-stoppers are encountered.



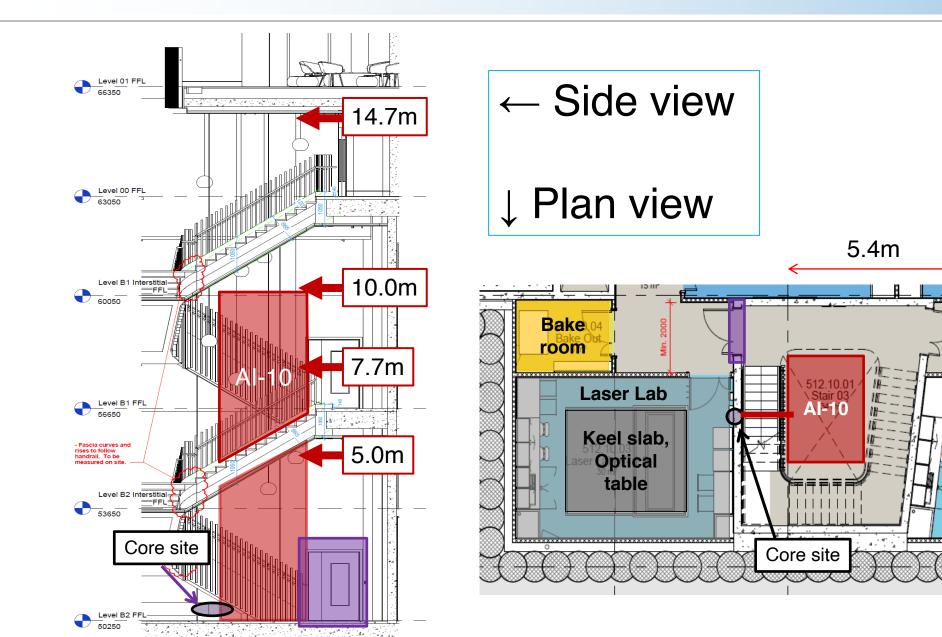


### **Beecroft building, Oxford Physics**





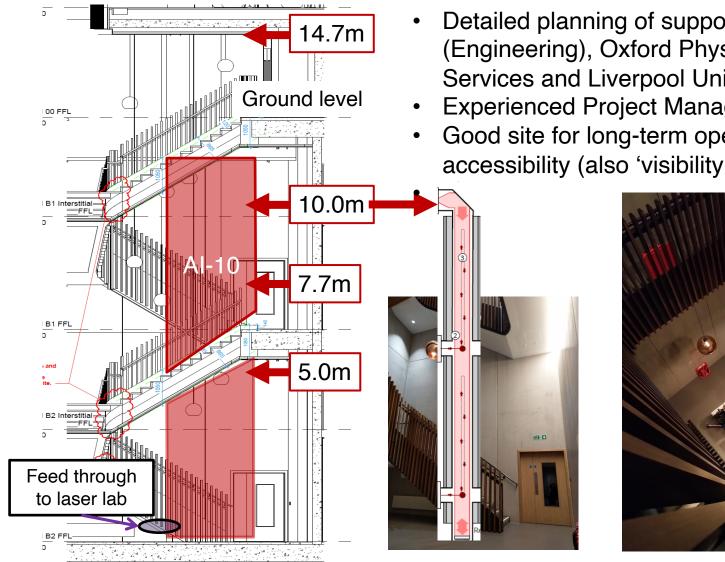
#### **Beecroft building, Oxford Physics**



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### **AION-10 site: Beecroft building, Oxford Physics**

#### **Beecroft building – brand new, low-vibration laser lab and concrete stairwell**



- Detailed planning of support structure by RAL (Engineering), Oxford Physics Technical Services and Liverpool Univ.
- **Experienced Project Manager: Roy Preece**
- Good site for long-term operation and wide accessibility (also 'visibility' and outreach).

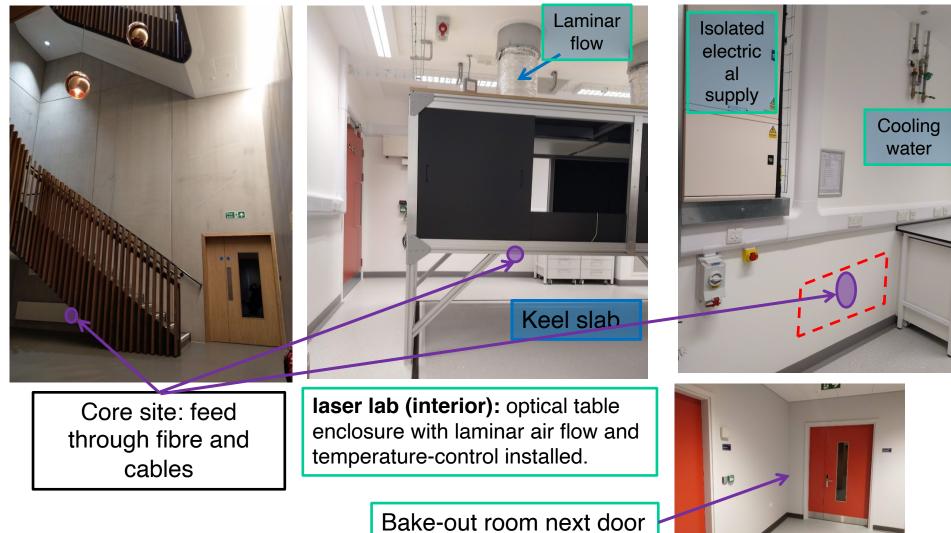




#### **Beecroft building laser lab**

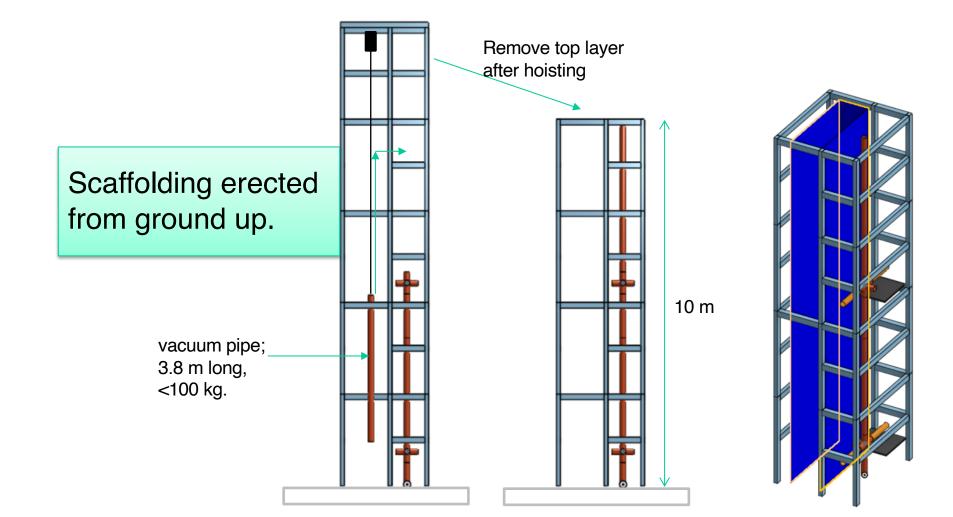


Beecroft stairwell: lowest level





# Assembly: extruded aluminium support structure



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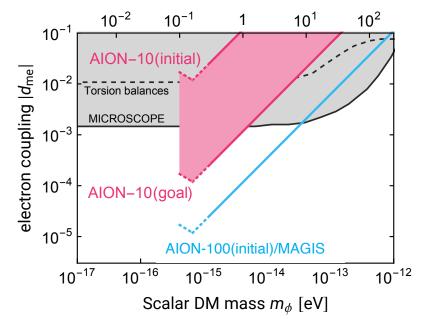


# THE SCIENCE CASE IN A NUTSHELL





# Main AION Physics Goals: Dark Matter and Gravitational Waves



#### Scientific Leadership in phenomenology already established:

#### The AION Physics Case:

AION Collaboration, AION: An Atom Interferometer Observatory and Network, arXiv:1911.11755. [accepted for publication in JCAP]

#### Working with leading theorists:

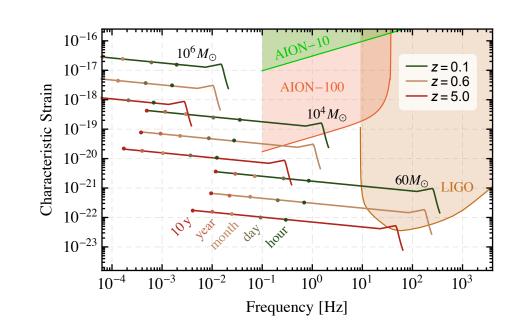
J. Ellis, M. Haehnelt, C. McCabe, J. March-Russell (AION), C. Burrage, ...

#### AEDGE

Y. El-Neaj, ..., O. Buchmueller *et al.* AEDGE: Atomic Experiment for Dark Matter And Gravity Exploration in Space, arXiv:1908.00802, *EPJ Quantum Technol.* 7, 6 (2020). [Submitted to ESA Voyage2050 call]

#### Main Physics Goals:

- Search for Ultra-Light Dark Matter
- Explore new parameter space and complement other searches.
- Focus on Scalar DM with Vector and Peudoscalar DM also under study.
- Gravitational Waves in mid-frequency band
- Explore frequencies between LISA and LIGO/VIRGO, KAGRA and Einstein Telescope
- Targets: Black hole mergers, phase transitions and cosmic string collisions





# **SEARCHES FOR ULTRA-LIGHT DARK MATTER**

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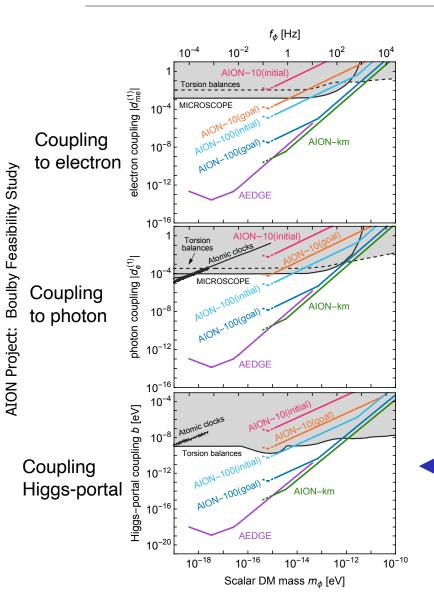


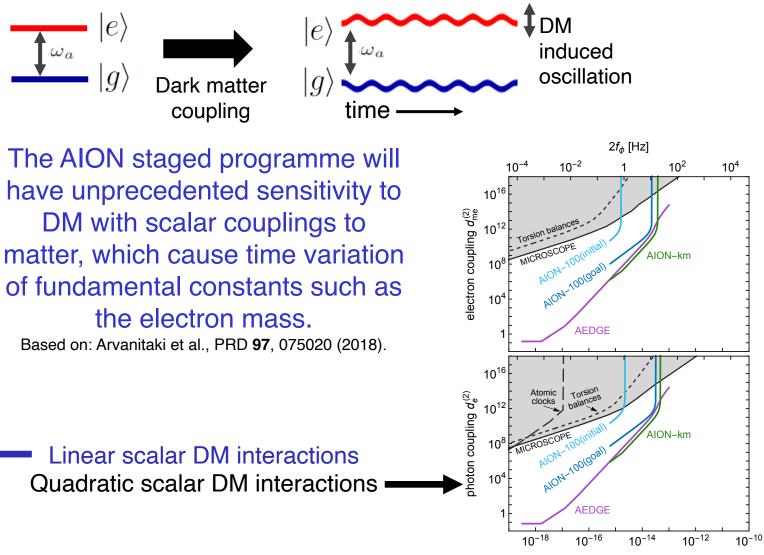
#### **Ultra-Light Scalar Dark Matter**

e

g

 $\omega_a$ 



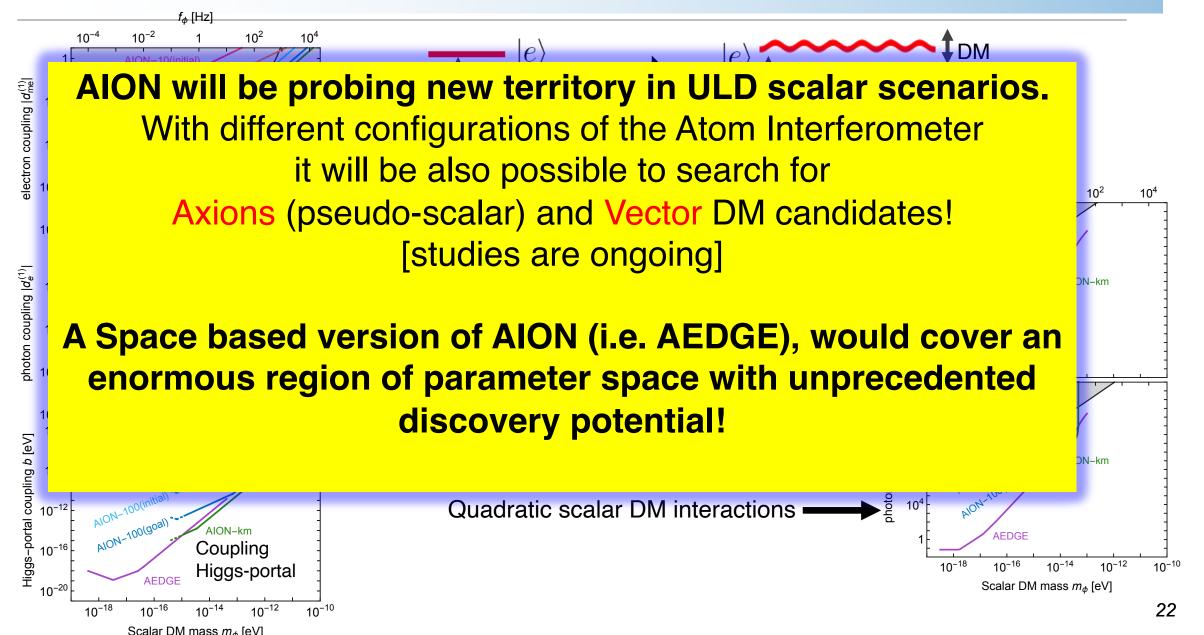


Scalar DM mass  $m_{\phi}$  [eV]





### **Ultra-Light Scalar Dark Matter**

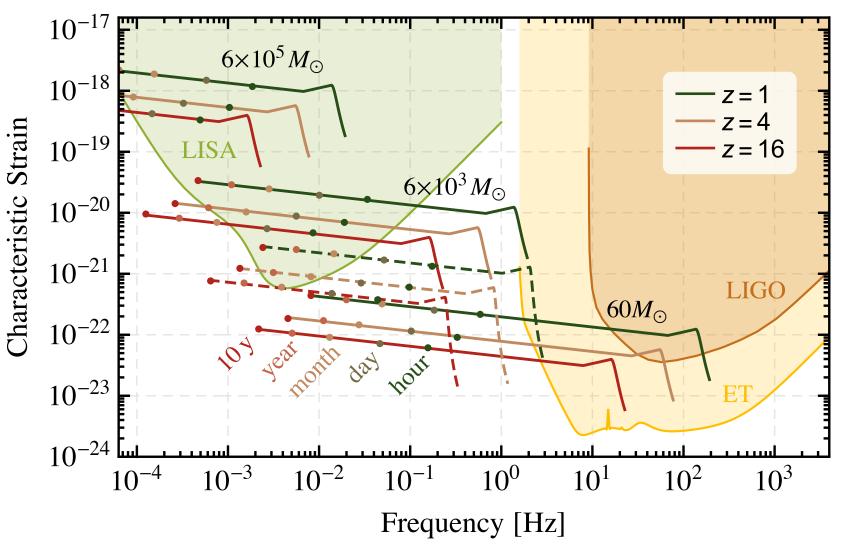




# UNEXPLORED MID-FREQUENCY GRAVITATIONAL WAVES



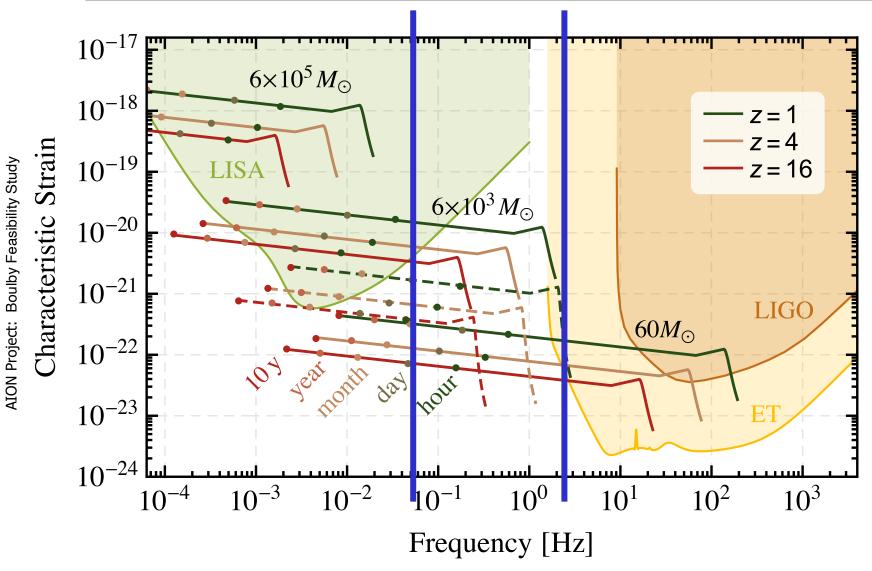
Pathway to the GW Mid-(Frequency)



AION Project: Boulby Feasibility Study

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### Pathway to the GW Mid-(Frequency)



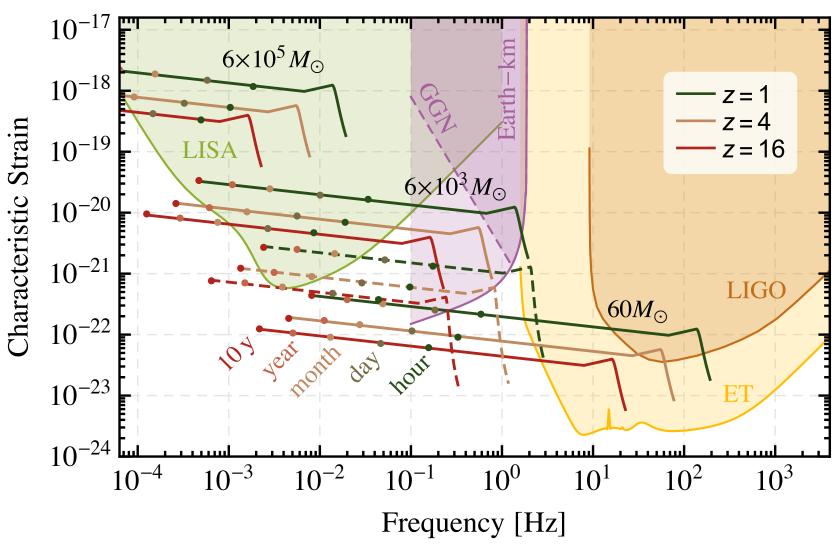
#### **Mid-band science**

- Detect sources BEFORE they reach the high frequency band [LIGO, ET]
- Optimal for sky localization: predict when and where events will occur (for multi-messenger astronomy)
- Search for Ultra-light dark matter in a similar frequency [i.e. mass] range

Mid-Band currently NOT covered A A A



#### **AION: Pathway to the GW Mid-(Frequency)**



#### **Mid-band science**

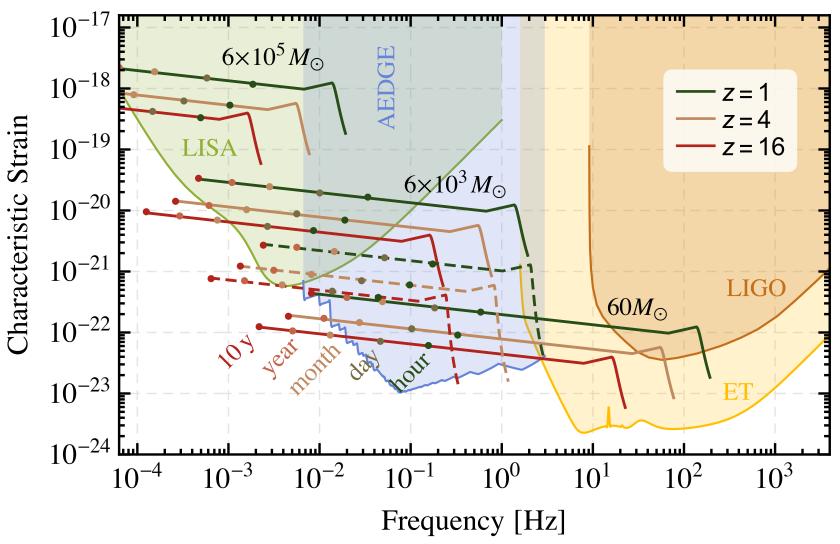
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AION: Terrestrial detectors can start filling this gap

AI ON



#### **AION: Pathway to the GW Mid-(Frequency)**



#### **Mid-band science**

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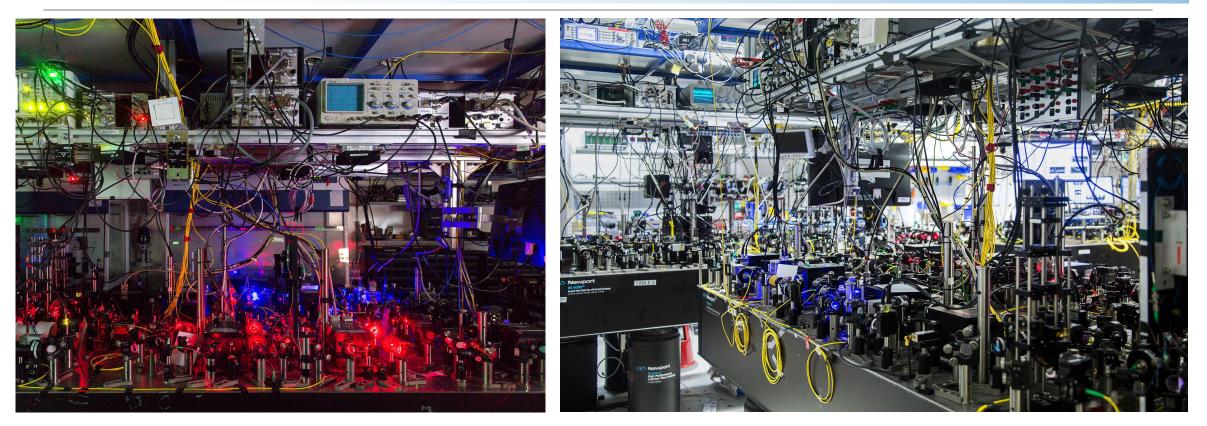
AEDGE Ultimate coverage with a space based detector AI ON



- v on gravitational
- AION & AEDGE provide a new window on gravitational physics, astrophysics & cosmology using atom interferometers, leveraging UK and international investment in quantum technologies, providing new opportunities for technology and science communities.
- The programme will enable the exploration of properties of dark matter, gravitation waves, as well as searches for new fundamental interactions with unprecedented sensitivity, opening many ground breaking discovery opportunities!







To push the state-of-the-art single photon Sr Atom Interferometry, the AION project builds dedicated Ultra-Cold Strontium Laboratories in: **Birmingham, Cambridge, Imperial College, Oxford, and RAL** 

The laboratories are expected to be fully operational in fall 2021.