

Akitaka Ariga

University of Bern / Chiba University

FASER/FASERnu

From neutrinos to new particle searches



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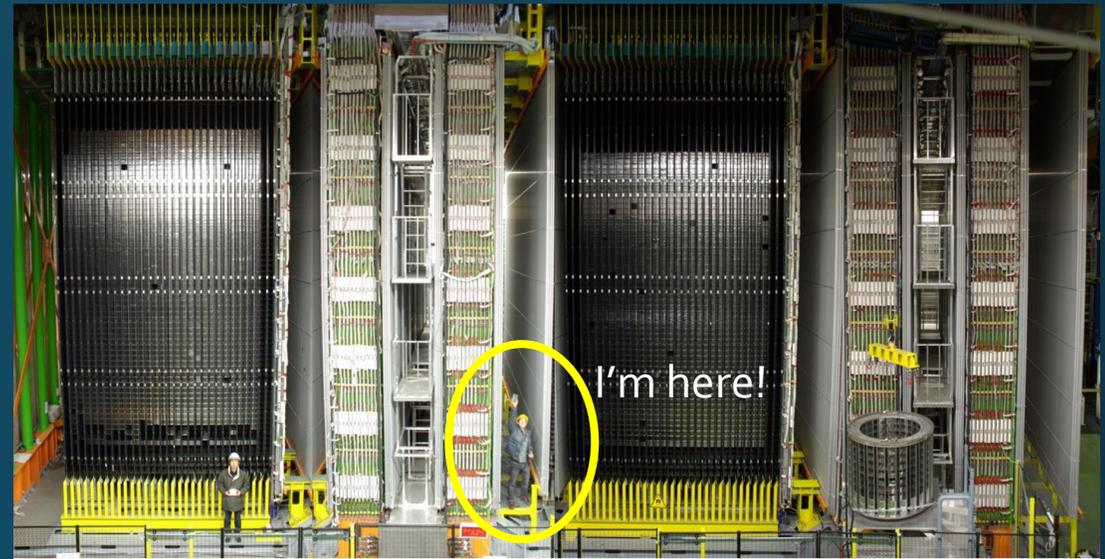


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Who am I

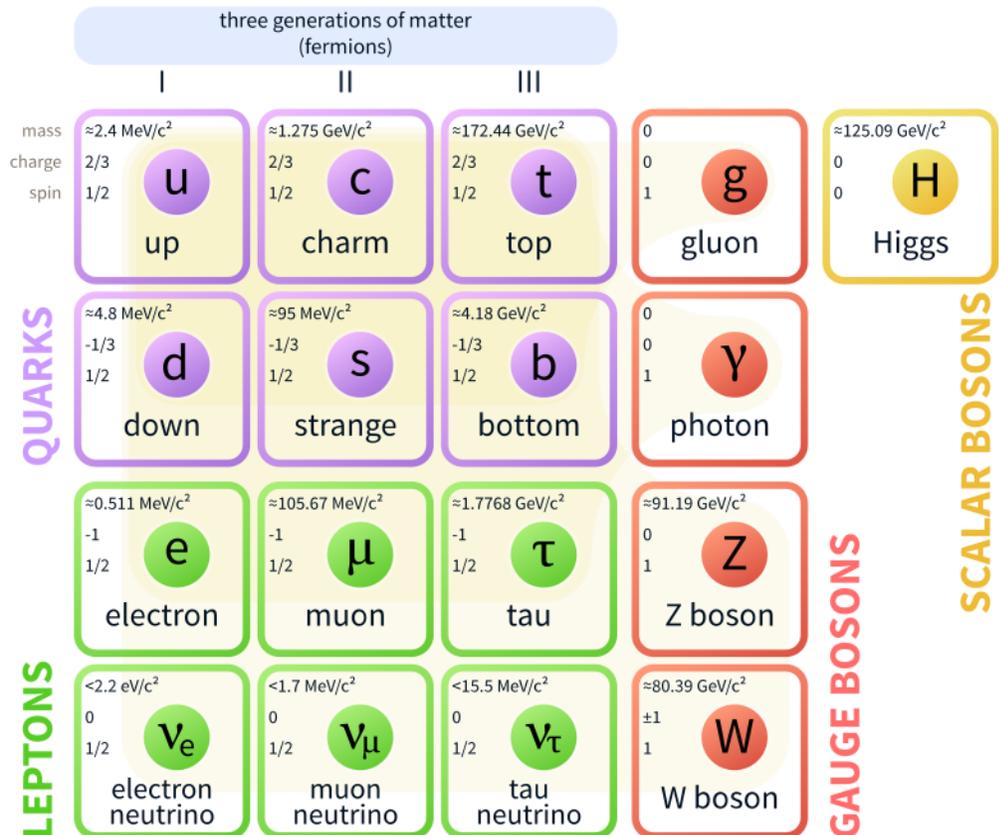
- OPERA, $\nu_{\mu} \rightarrow \nu_{\tau}$
- T2K, muon flux
- (HK, astro physics)
- AEGIS/QuPlas, antimatter exps
- Eiger-mu, Glacier muon tomography
- NA65/DsTau, tau neutrino production
- FASERnu, LHC neutrinos

- Expertise with emulsion detectors

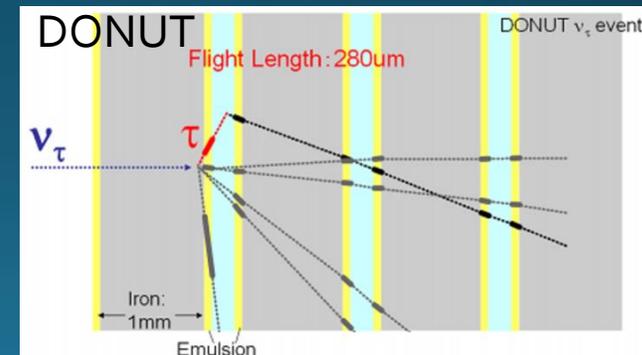


Standard model and neutrinos

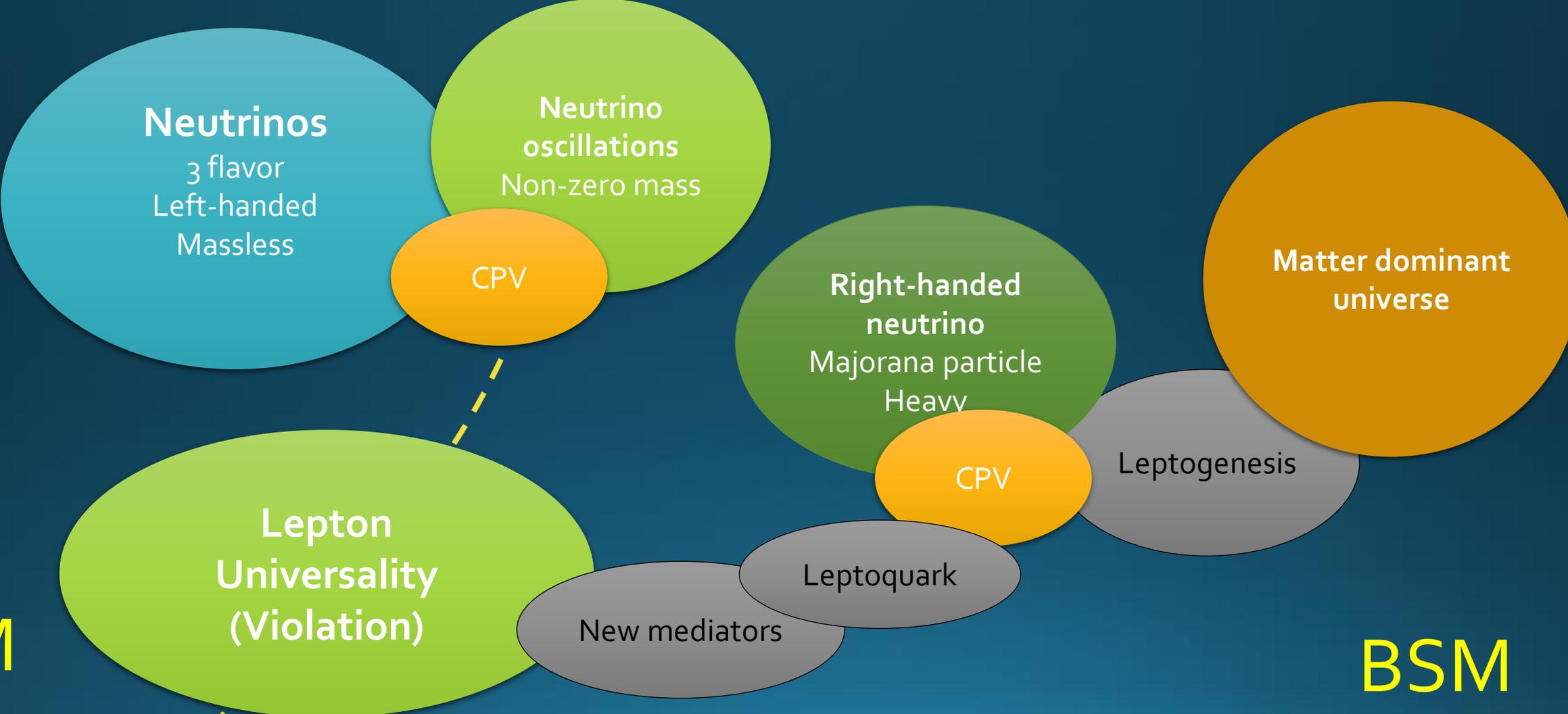
Standard Model of Elementary Particles



- In SM, neutrinos are
 - Neutral
 - Only weak interaction
 - Left-handed → massless
 - 3 flavors



Neutrino physics

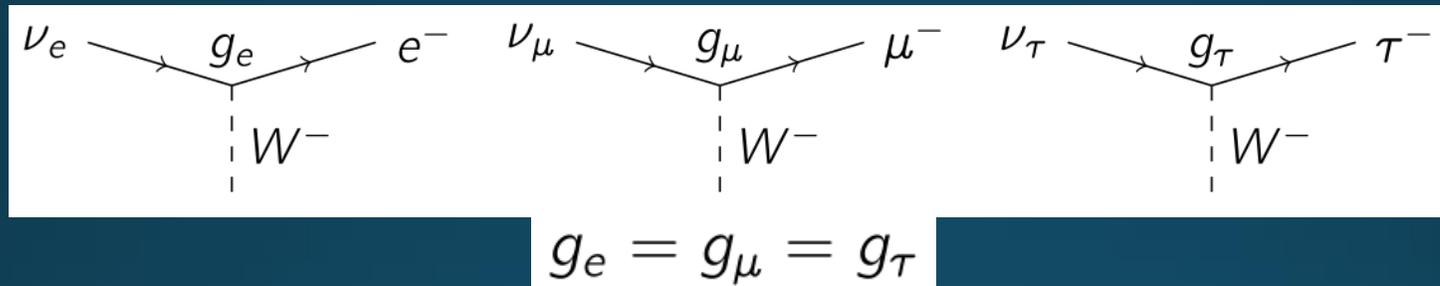


SM

BSM

Lepton Flavor Universality, “Sacred principle” of the SM

- Three lepton families equally couple to weak boson



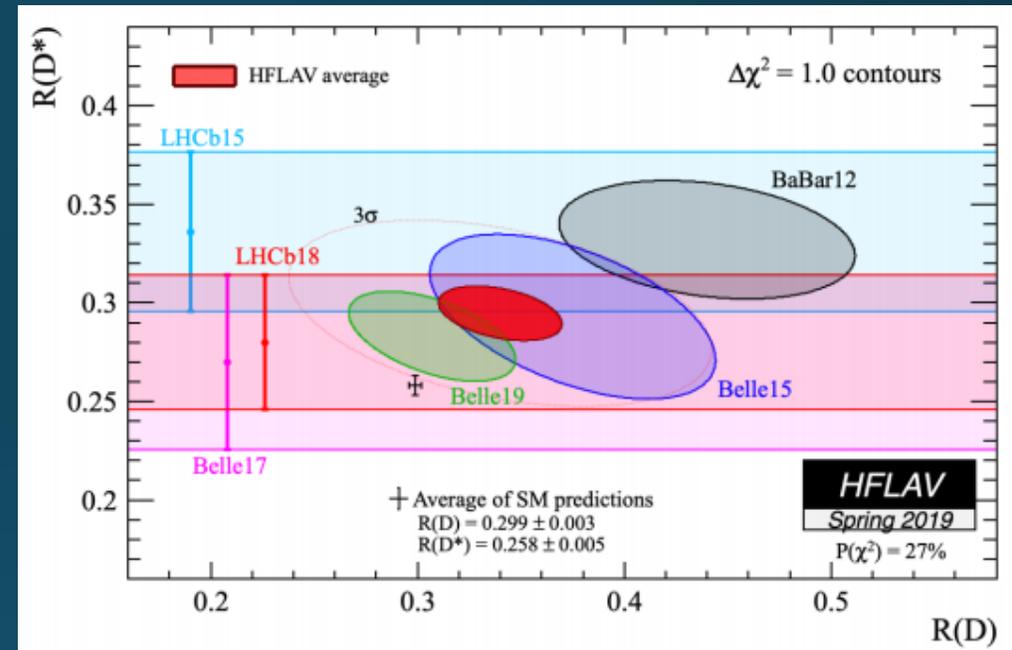
- Intensively verified with very high accuracy, for example

$$\left(\frac{g_\tau}{g_\mu}\right)^4 = 0.178 \left(\frac{m_\mu}{m_\tau}\right)^5 \left(\frac{\tau_\mu}{\tau_\tau}\right) \longrightarrow \frac{g_\tau}{g_\mu} = 0.999 \pm 0.003$$

- It was consistent with all experimental results,,, until recently

Flavor anomaly

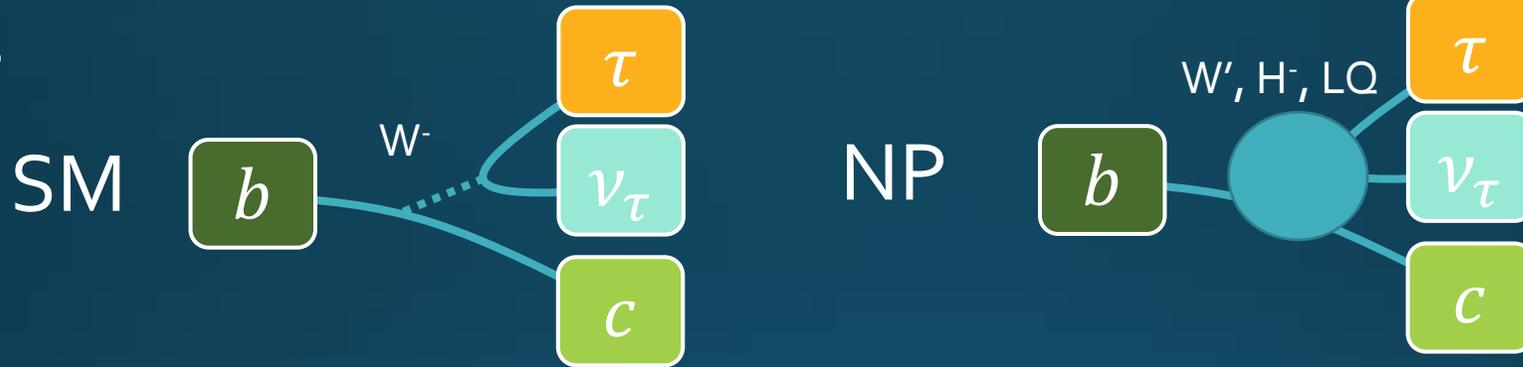
$$R(D) = \frac{\mathcal{B}(B \rightarrow \tau \nu_\tau D)}{\mathcal{B}(B \rightarrow \mu \nu_\mu D)}$$



Possible contribution from new physics in heavy flavors!?

New physics effect?

B decays

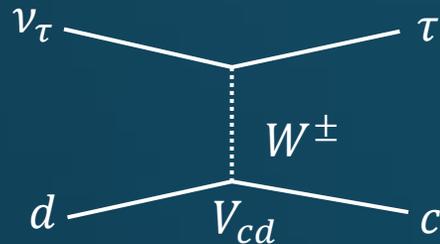


Neutrino CC beauty production



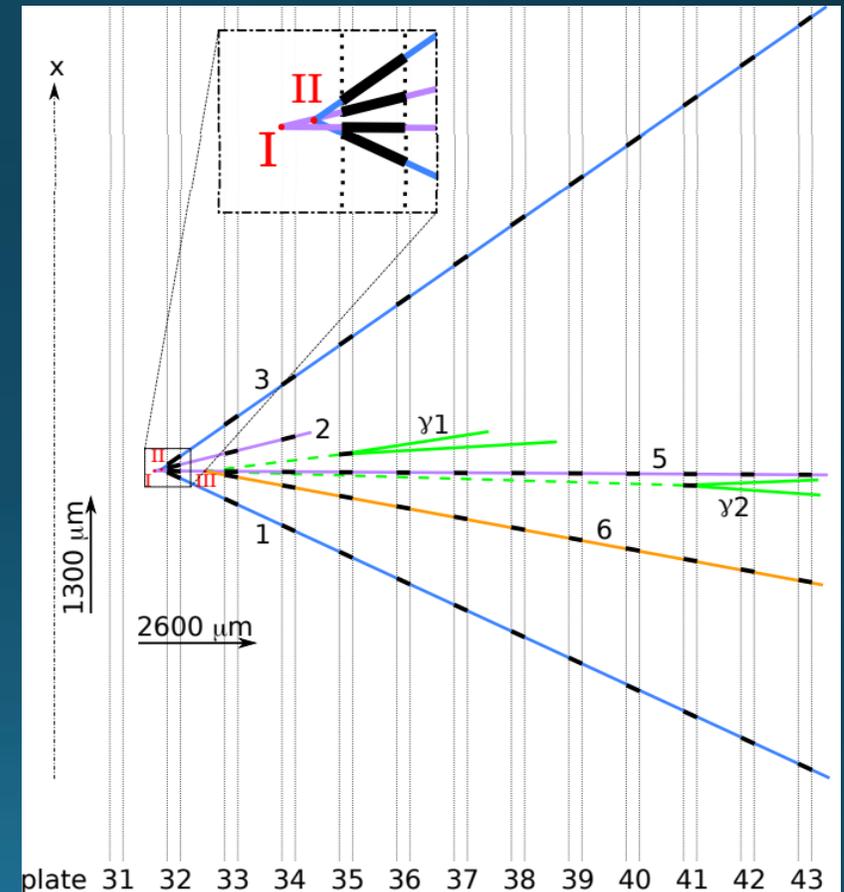
OPERA's ν_τ induced charm production event

SM process,
charm production
via mixing

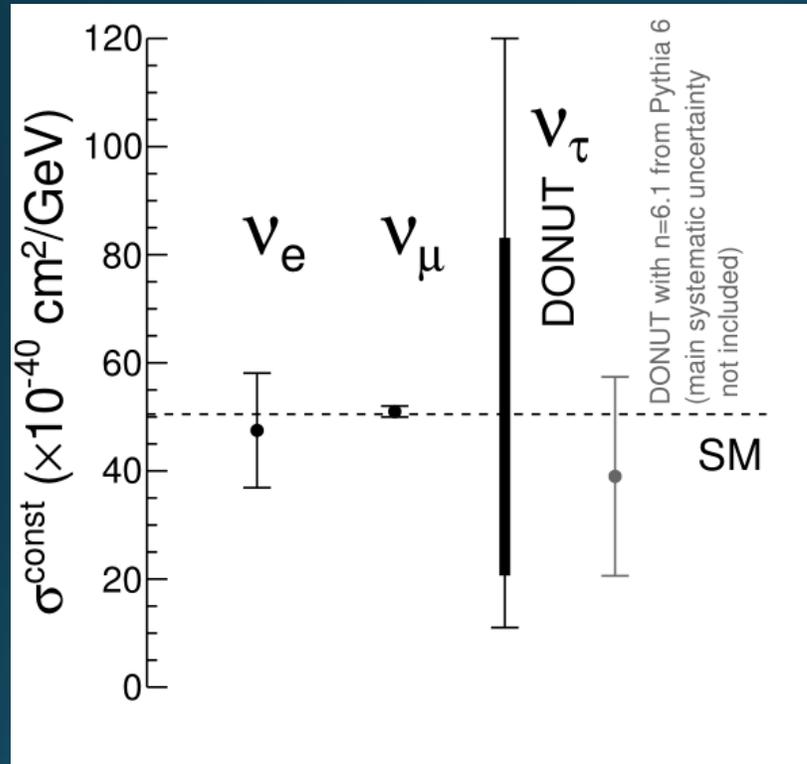


Well measured for ν_μ

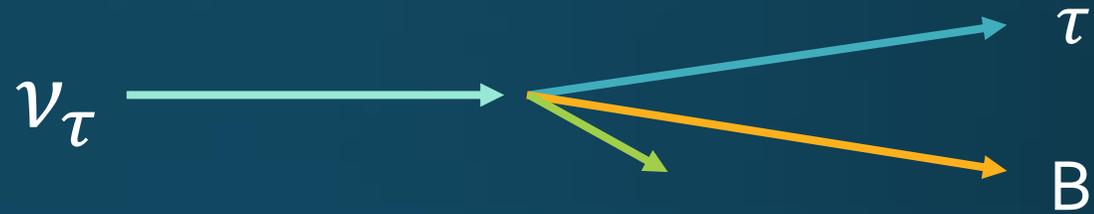
- 1 event was observed with surprise
- Expectation:
 - Signal 0.04
 - Background < 0.05
- Could also be a hint of new physics!?



Status of Lepton Universality testing in neutrino scattering



Poor constraint for ν_τ

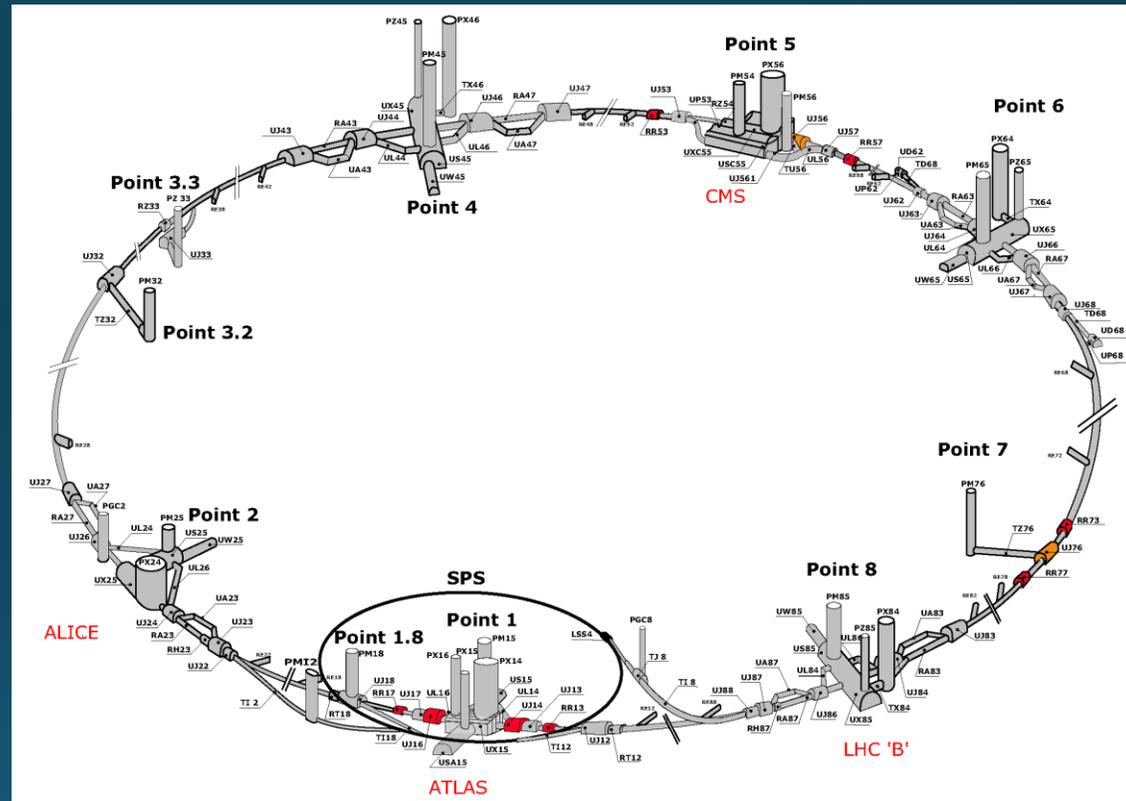


High energy neutrinos ($E_\nu > 100 \text{ GeV}$) is required to access heavy flavor channels

→ Need high statistics and high energy beam experiment!

LHC as neutrino source?

Large Hadron Collider
27 km circumference
7 TeV + 7 TeV

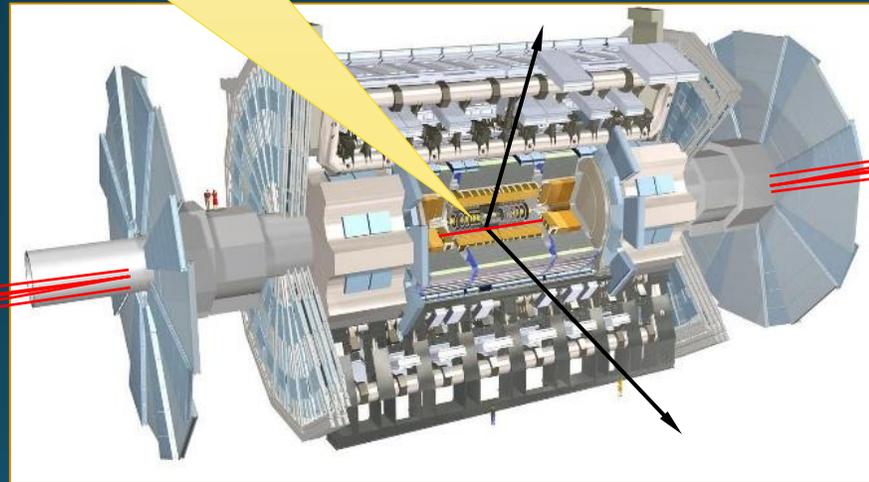


Let's open new domain of research! **Neutrino**

Wait! There is no neutrino beamline!!

LHC as a neutrino source

14 TeV p-p collision



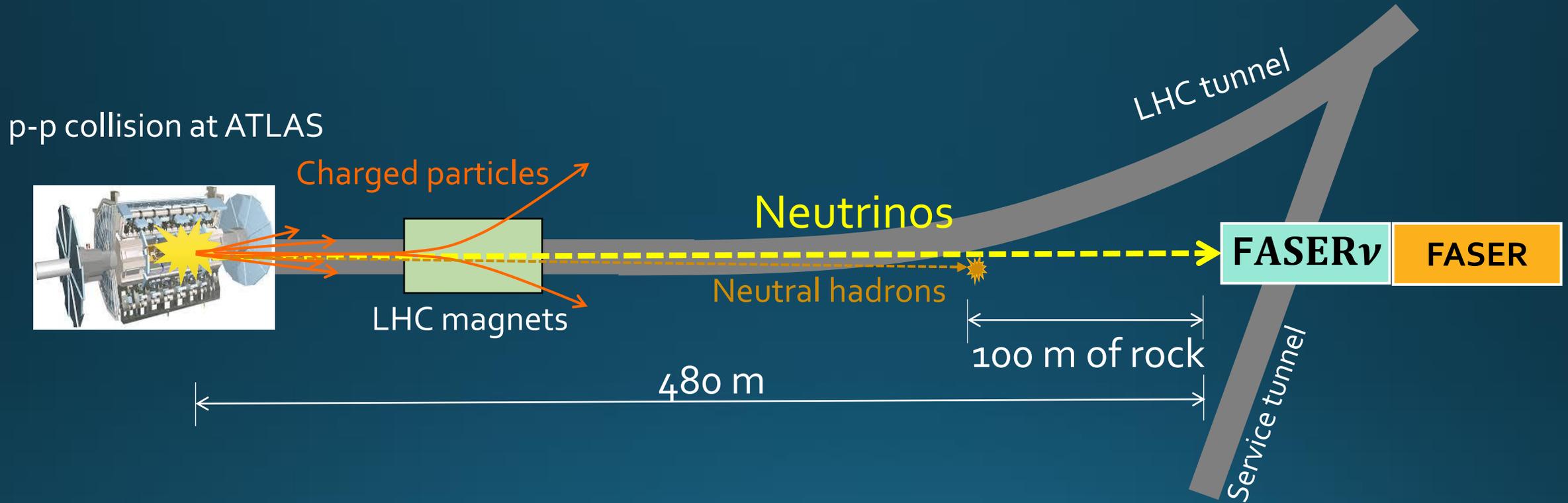
No experiment has sought neutrinos at the LHC so far!

Intense neutrino beam (+ long lived particles, LLPs) here!



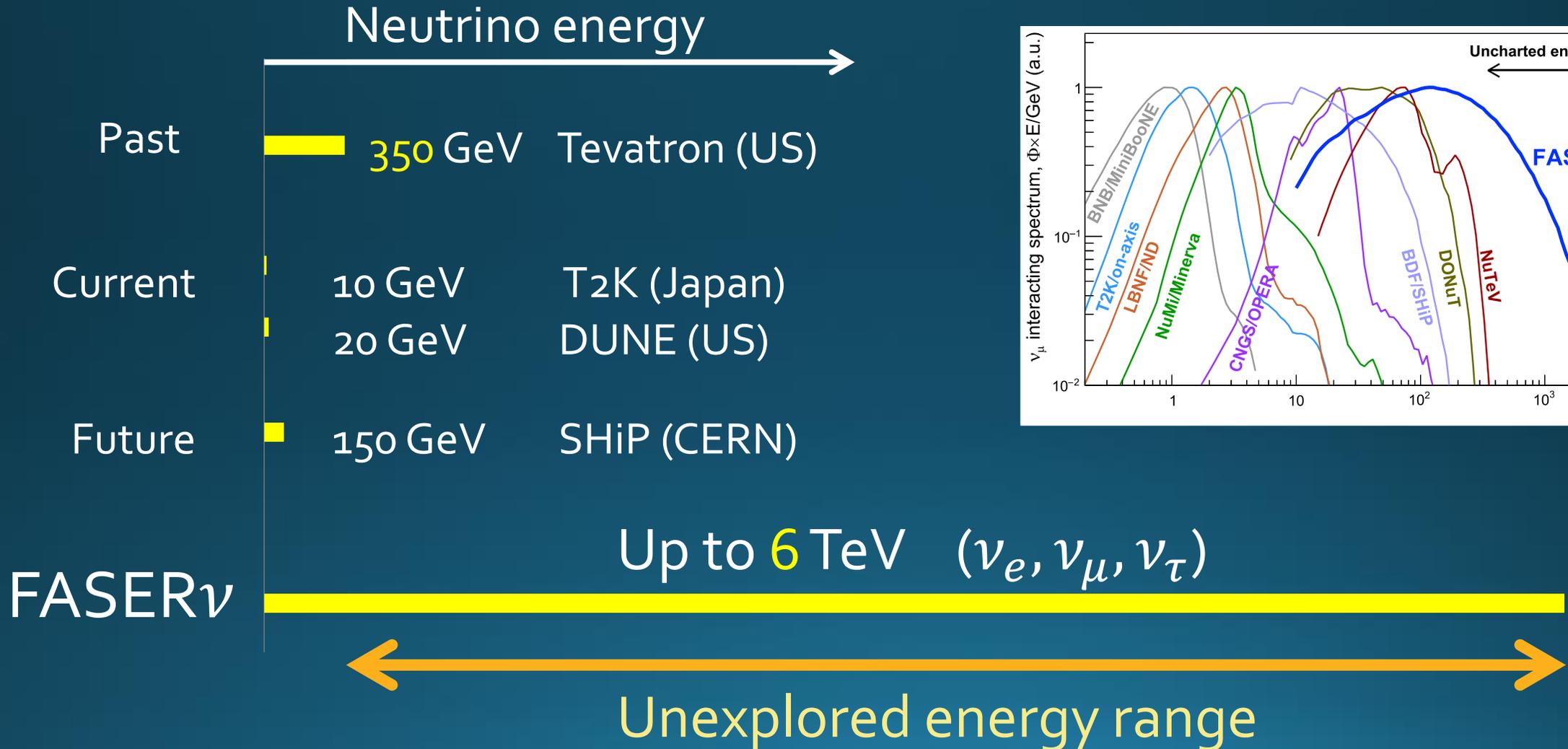
FASER (new particle searches) was approved by CERN in Mar 2019
FASER ν (neutrino program) was approved by CERN in Dec 2019

Novel “neutrino beamline”

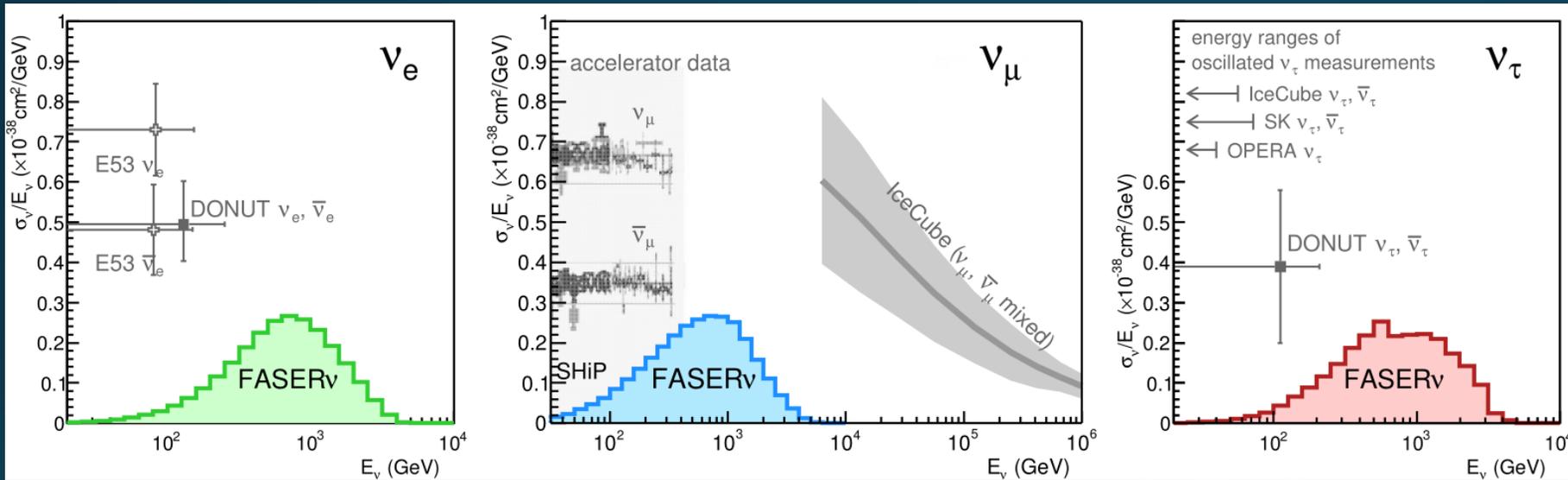


Negligible cost for infrastructure

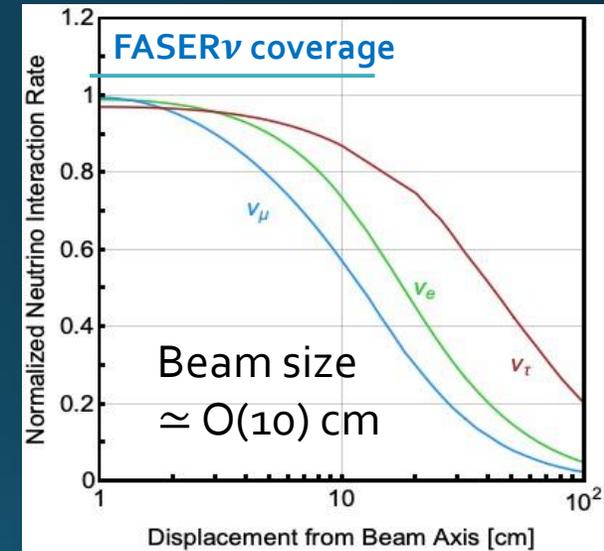
High energy frontier



Neutrino spectrum at FASER ν



Unexplored energy regime for all three flavors



Collimated beam

Neutrinos at the LHC: New domain of neutrino research!

- Neutrinos by **collider method**
- **High energy frontier** ~ TeV
- Study of production, propagation and interactions of high energy neutrinos

Production

14 TeV p-p collision \equiv 100 PeV int
in fixed target ($\sqrt{s} \sim 10$ TeV)

Prompt neutrino production \rightarrow
Input for neutrino telescopes

QCD (charm/gluon PDF,
intrinsic charm)

Propagation

Unique energy and baseline,
 $L/E \sim 10^{-3}$ m/MeV

Neutrino oscillation at
 $\Delta m^2 \sim 1000$ eV²

Interaction

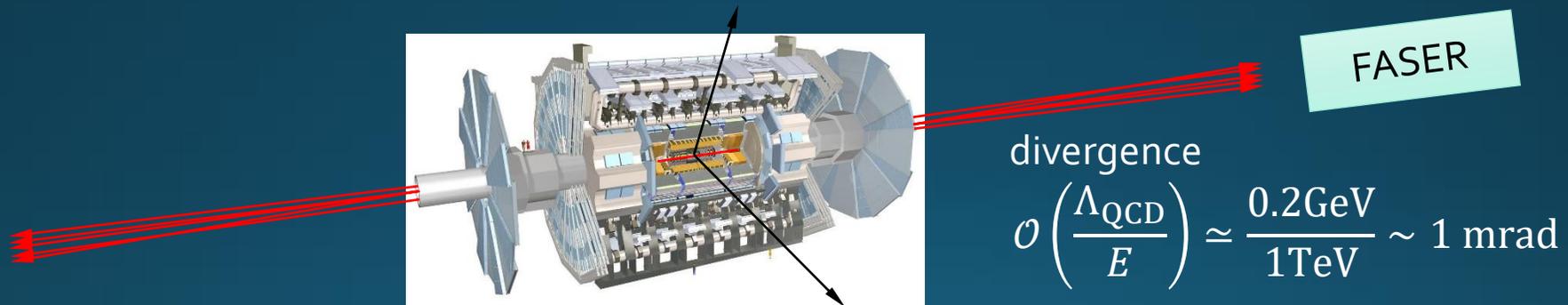
3-flavor neutrino cross sections
in unexplored energy range

Neutrino induced heavy quark
productions

New physics effects

FASER: Forward Search Experiment at the LHC

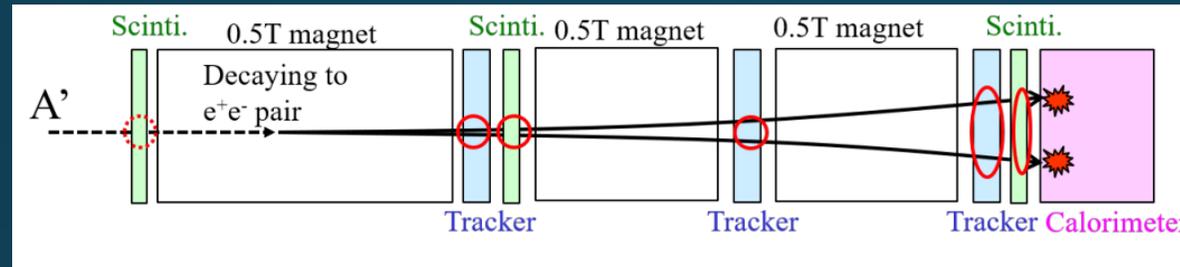
- ATLAS and CMS searches focus on high $p_T \rightarrow$ appropriate for heavy, strongly interacting particles
 - No evidence of new particles is detected so far.
- If new particles are **light and weakly interacting** to the SM particles (e.g. **dark photon**), they could be long-lived and collimated in the very forward region \rightarrow FASER arXiv:1708.09389 , 1811.12522



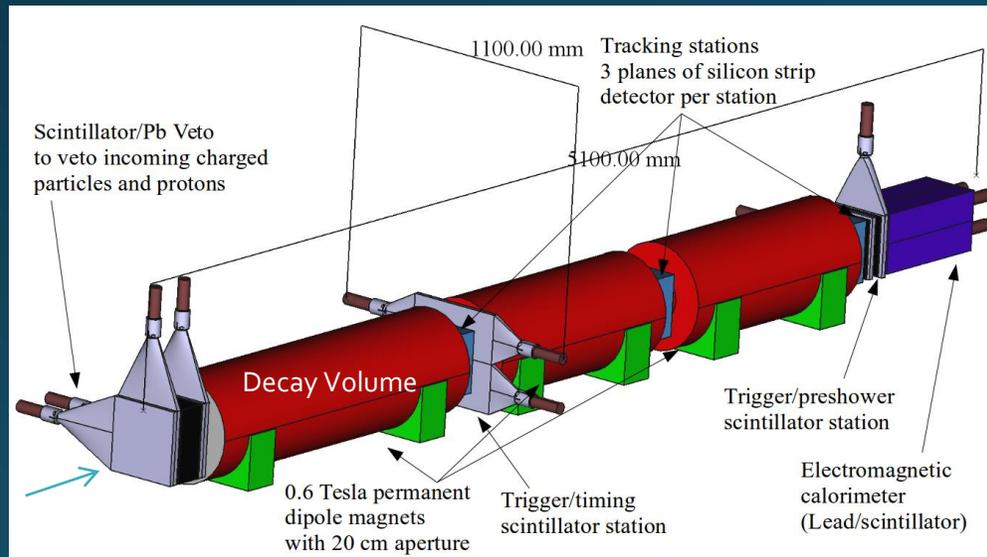
- The LOI (July 2018) and technical proposal (October 2018) were submitted. **Approved by CERN in March 2019.**
- Preparing for physics run in 2021 (Run3 of the LHC operation)

FASER detector & sensitivity

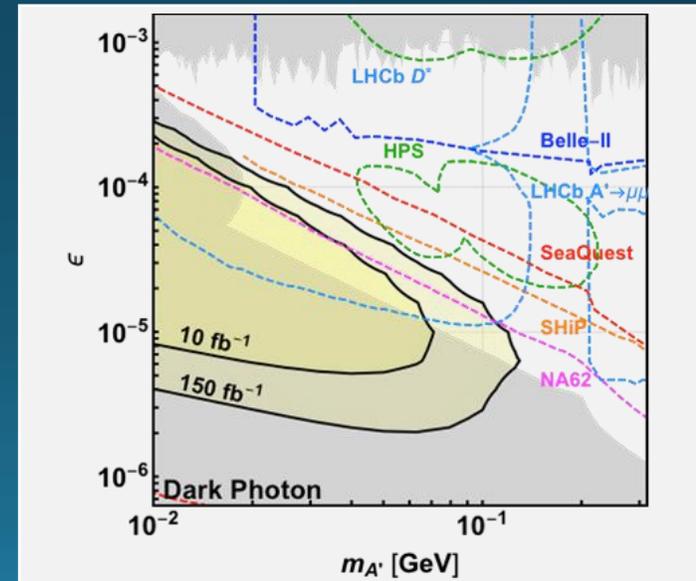
- Dark photon: Photon in dark sector, and it has mass
- Signal: Dark photon decay into e^+e^- pair



Detector schematic (original one without FASERnu)



Sensitivity for dark photon search in Run 3



THE FASER COLLABORATION

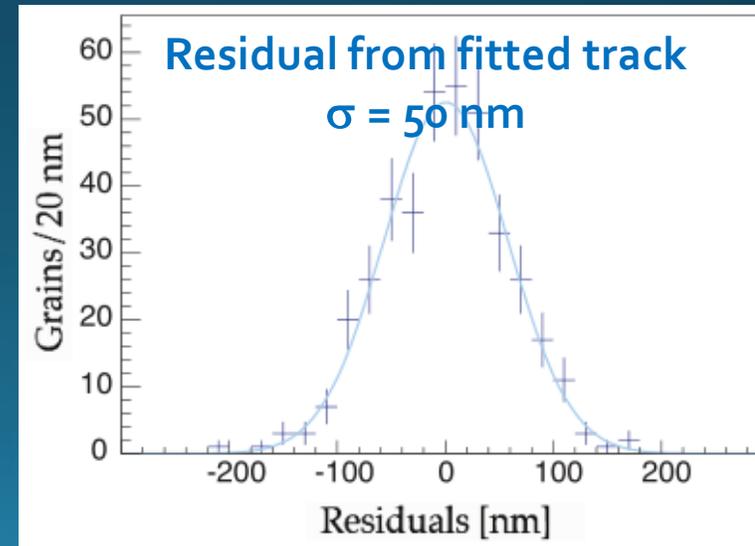
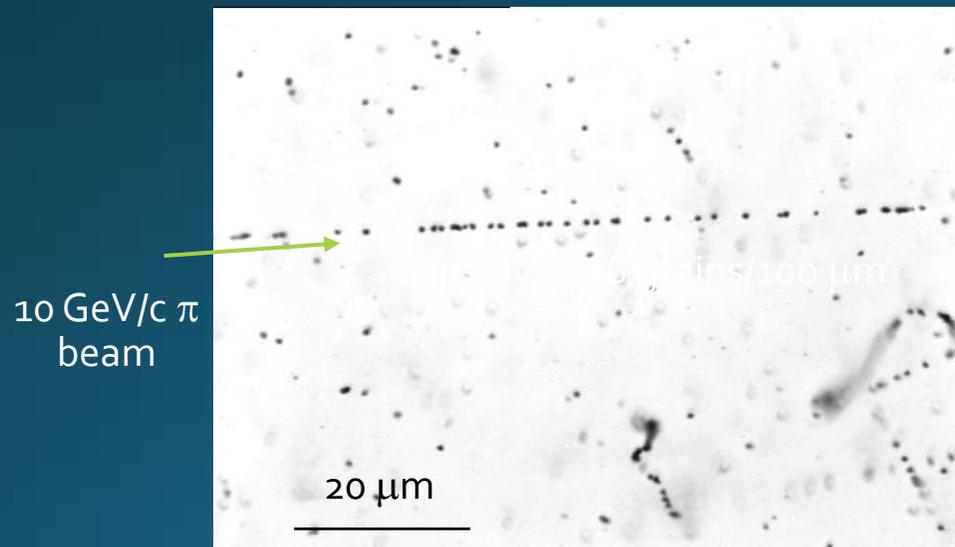
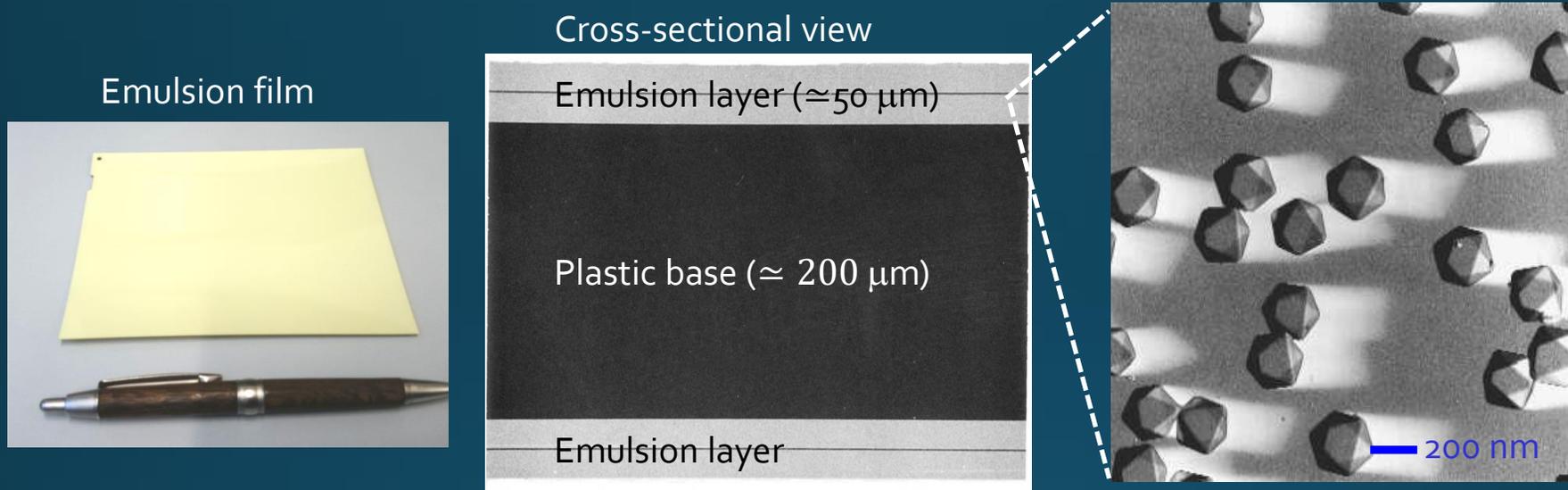
- 64 collaborators, 20 institutions, 8 countries

Henso Abreu (Technion), Yoav Afik (Technion), Claire Antel (Geneva), Akitaka Ariga (Bern), Tomoko Ariga (Kyushu/Bern), Florian Bernlochner (Bonn), Jamie Boyd (CERN), Lydia Brenner (CERN), Dave Casper (UC Irvine), Franck Cadoux (Geneva), Xin Chen (Tsinghua), Andrea Coccaro (INFN), Candan Dozen (Tsinghua), Yannick Favre (Geneva), Deion Fellers (Oregon), Jonathan Feng (UC Irvine), Didier Ferrere (Geneva), Iftah Galon (Rutgers), Stephen Gibson (Royal Holloway), Sergio Gonzalez-Sevilla (Geneva), Shih-Chieh Hsu (Washington), Zhen Hu (Tsinghua), Peppe Iacobucci (Geneva), Sune Jakobsen (CERN), Enrique Kajomovitz (Technion), Felix Kling (SLAC), Umut Kose (CERN), Susanne Kuehn (CERN), Helena Lefebvre (Royal Holloway), Lorne Levinson (Weizmann), Ke Li (Washington), Jinfeng Liu (Tsinghua), Chiara Magliocca (Geneva), Josh McFayden (CERN), Sam Meehan (CERN), Dimitar Mladenov (CERN), Mitsuhiro Nakamura (Nagoya), Toshiyuki Nakano (Nagoya), Marzio Nessi (CERN), Friedemann Neuhaus (Mainz), Hidetoshi Otono (Kyushu), Carlo Pandini (Geneva), Hao Pang (Tsinghua), Brian Petersen (CERN), Francesco Pietropaolo (CERN), Markus Prim (Bonn), Michaela Queitsch-Maitland (CERN), Filippo Resnati (CERN), Jakob Salfeld-Nebgen (CERN), Osamu Sato (Nagoya), Paola Scampoli (Bern), Kristof Schmieden (Mainz), Matthias Schott (Mainz), Anna Sfyrta (Geneva), Savannah Shively (UC Irvine), Jordan Smolinsky (Florida), Yosuke Takubo (KEK), Ondrej Theiner (Geneva), Eric Torrence (Oregon), Sebastian Trojanowski (Sheffield), Serhan Tufanli (CERN), Benedikt Vormwald (CERN), Dengfeng Zhang (Tsinghua), Gang Zhang (Tsinghua)



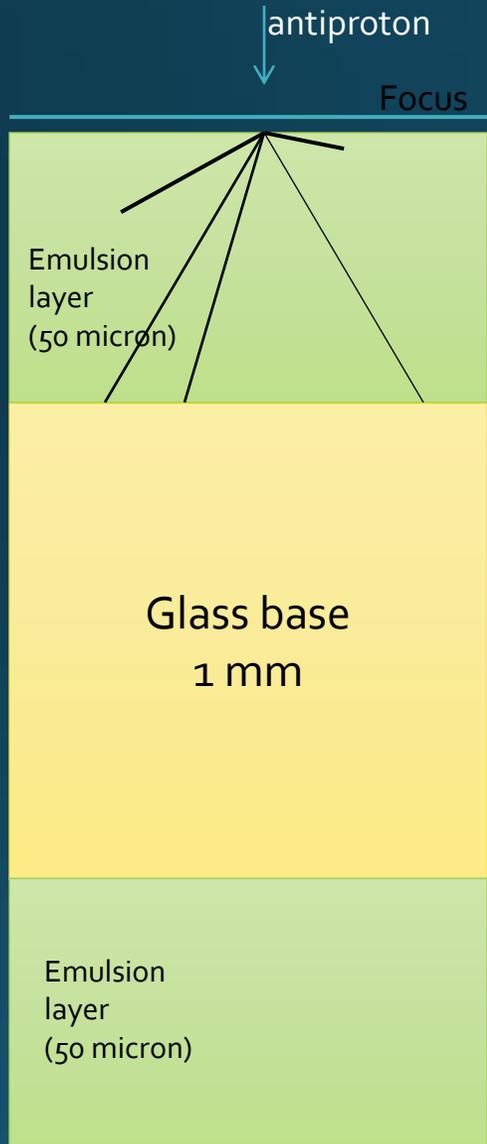
Emulsion detectors: 3D tracking device with 50 nm precision

AgBr crystal = detector
 10^{14} channels/film or 10^{14} channels/cm³



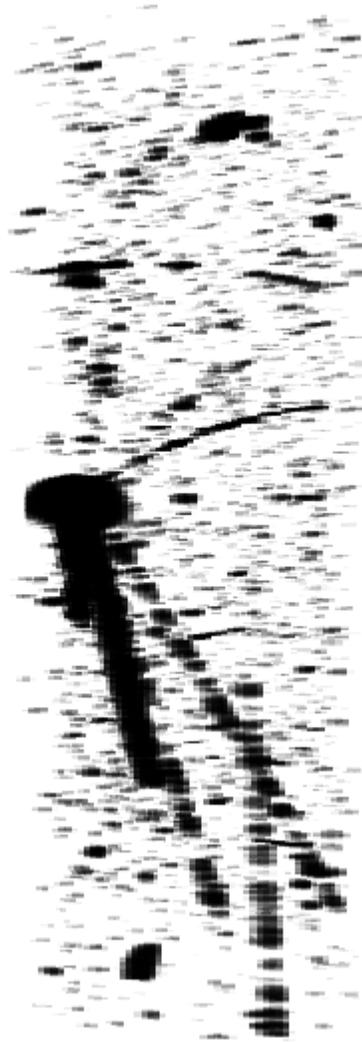
Antiproton annihilation in emulsion

Antiproton annihilation taken in AEgIS 2012



← 200 microns →

3D view of emulsion detector



- 3D high resolution hits
- Work as tracker
- dE/dx proportional to darkness (Number of grains)

150 μm x 120 μm x 50 μm

Emulsion = a detector with high detection channel density



$150\ \mu\text{m} \times 120\ \mu\text{m} \times 50\ \mu\text{m}$

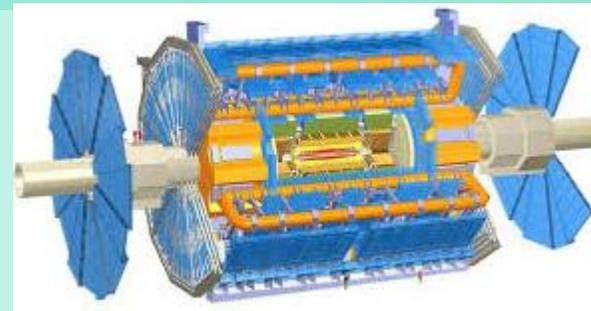
1.2×10^8 channels (crystal) in this volume.

1 film = 10^{14} channels

ATLAS-IBL pixel sensor
FE-14

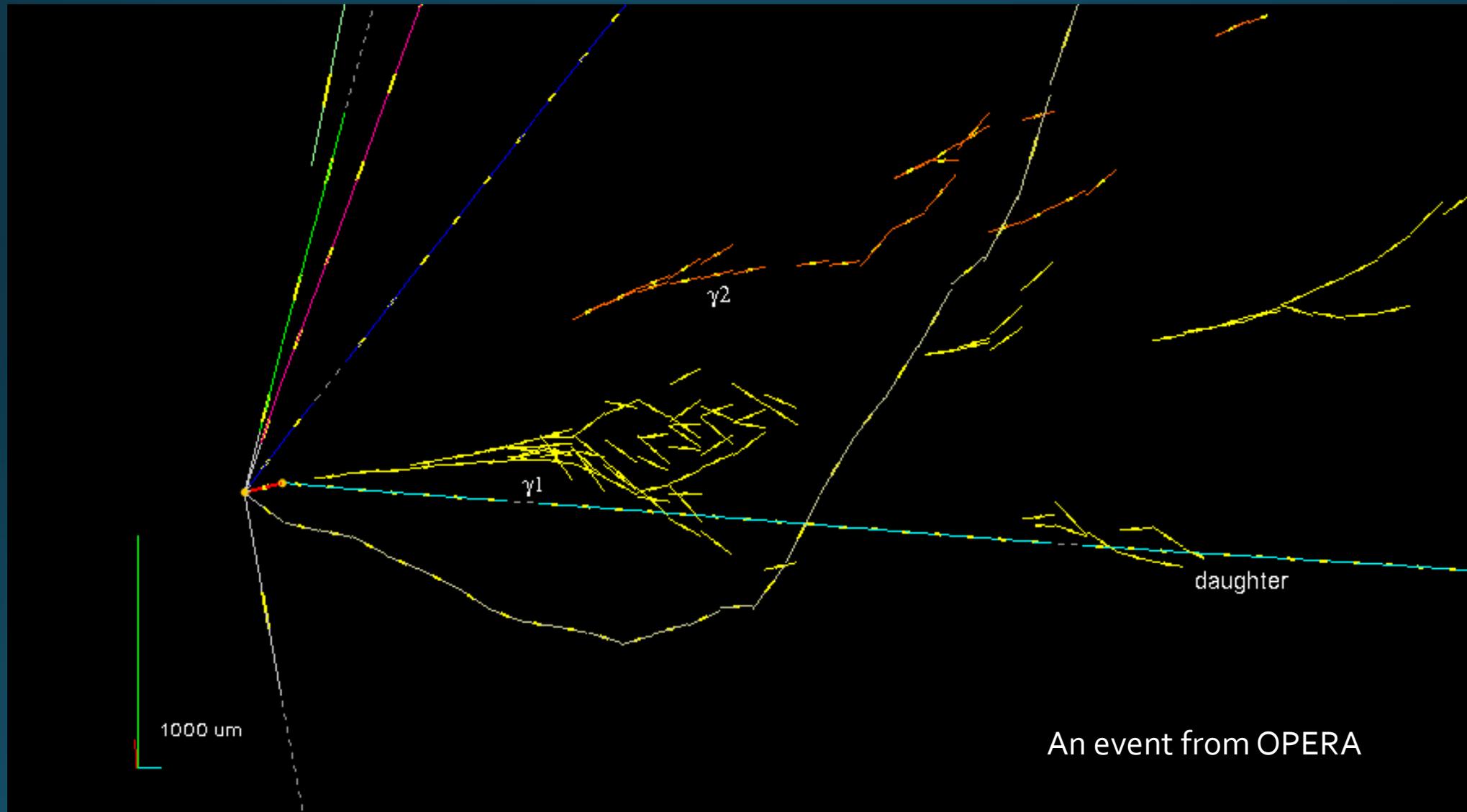
1 pixel =
 $250\ \mu\text{m} \times 50\ \mu\text{m} \times 200\ \mu\text{m}$

Sum of all channels in ATLAS = $\sim 10^8$

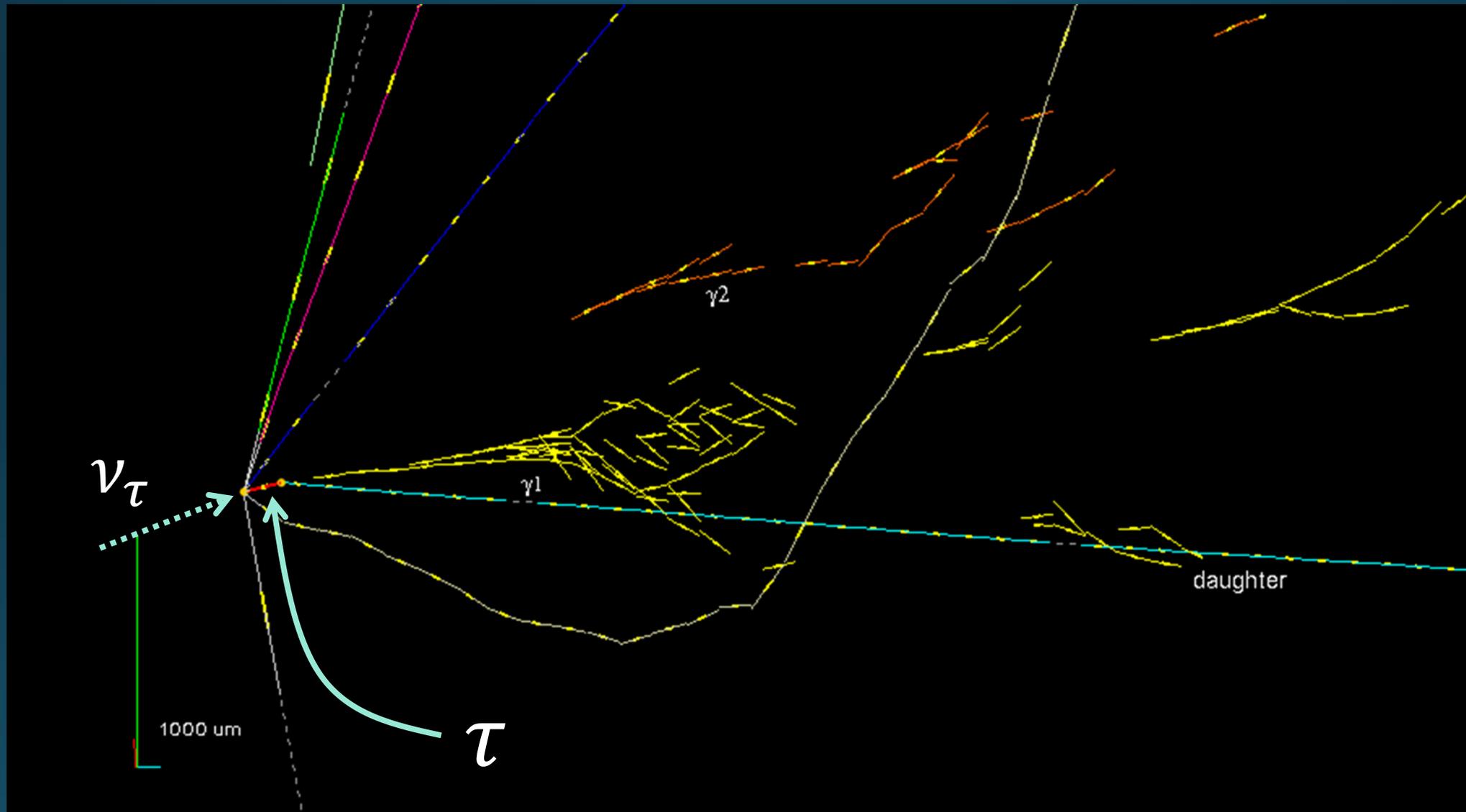


High density of detection channels, $O(10^{14})$ channels/cc, makes emulsion attractive for many purposes.

Emulsion-based neutrino detector

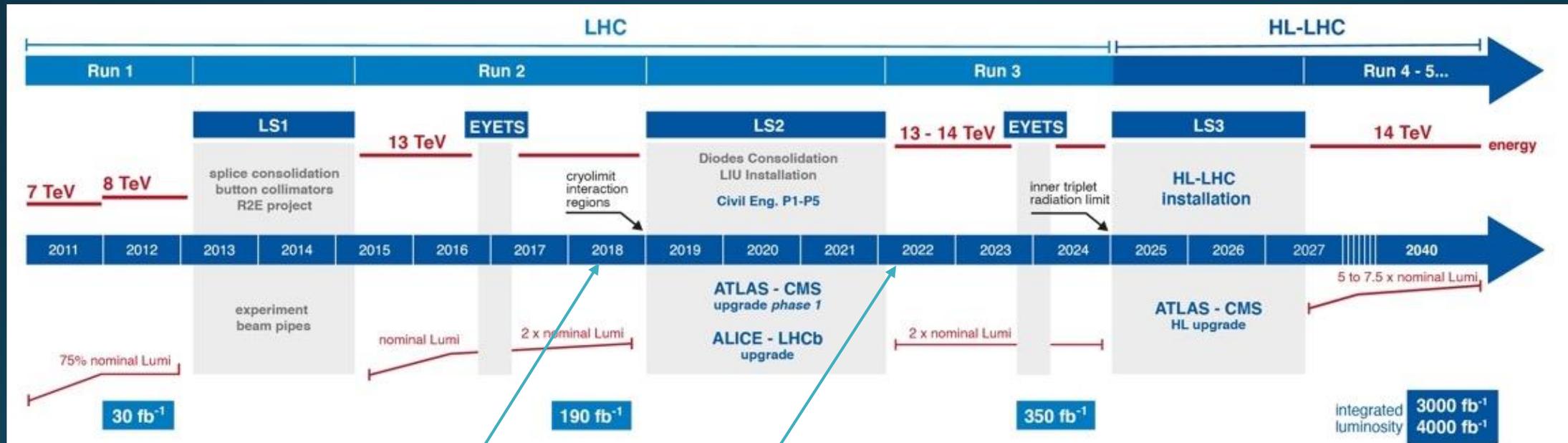


Emulsion-based neutrino detector



FASERnu Schedule

We are in LS2. Pilot run was performed in 2018.
Physics run will start in 2022.



BG measurement,
pilot run in 2018

Physics run will start in
2022 (~150 fb⁻¹)

FASER ν history (personal view)

-2011

- ν_τ with Tevatron? \rightarrow shutdown

2013-

- ν_τ with CERN SPS \rightarrow NA65/DsTau, SHiP

2018, in Run 2 of LHC operation

- April, first contact with FASER project
- May, joined to take BG data
- June, install emulsion detector
- July, \rightarrow emulsion can work!
- Sep-Oct, install a pilot neutrino detector and data taking
- Aug, FASER LOI
- Nov, FASER TP

2019

- Jan, First neutral interactions
- Aug, FASERnu LOI
- Oct, FASERnu Technical proposal
- Dec, FASERnu Approval
- Mar, FASER approval

10.1140/epjc/s10052-020-7631-5

CERN-EP-2019-160, KYUSHU-RCAPP-2019-003, SLAC-PUB-17460, UCI-TR-2019-19



Detecting and Studying High-Energy Collider Neutrinos with FASER at the LHC

FASER Collaboration

Henso Abreu,¹ Claire Antel,² Akitaka Ariga,^{3,*} Tomoko Ariga,^{3,4,*} Jamie Boyd,⁵ Franck Cadoux,² David W. Casper,⁶ Xin Chen,⁷ Andrea Coccaro,⁸ Candan Dozen,⁷ Peter B. Denton,^{9,†} Yannick Favre,² Jonathan L. Feng,⁶ Didier Ferrere,² Iftah Galon,¹⁰ Stephen Gibson,¹¹ Sergio Gonzalez-Sevilla,² Shih-Chieh Hsu,¹² Zhen Hu,⁷ Giuseppe Iacobucci,² Sune Jakobsen,⁵ Roland Jansky,² Enrique Kajomovitz,¹ Felix Kling,^{6,13,*} Susanne Kuehn,⁵ Lorne Levinson,¹⁴ Congqiao Li,¹² Josh McFayden,⁵ Sam Meehan,⁵ Friedemann Neuhaus,¹⁵ Hidetoshi Otono,⁴ Brian Petersen,⁵ Helena Pikhartova,¹¹ Michaela Queitsch-Maitland,⁵ Osamu Sato,¹⁶ Kristof Schmieden,⁵ Matthias Schott,¹⁵ Anna Sfyrla,² Savannah Shively,⁶ Jordan Smolinsky,⁶ Aaron M. Soffa,⁶ Yosuke Takubo,¹⁷ Eric Torrence,¹⁸ Sebastian Trojanowski,¹⁹ Callum Wilkinson,^{20,†} Dengfeng Zhang,⁷ and Gang Zhang⁷

¹Department of Physics and Astronomy,
Technion—Israel Institute of Technology, Haifa 32000, Israel

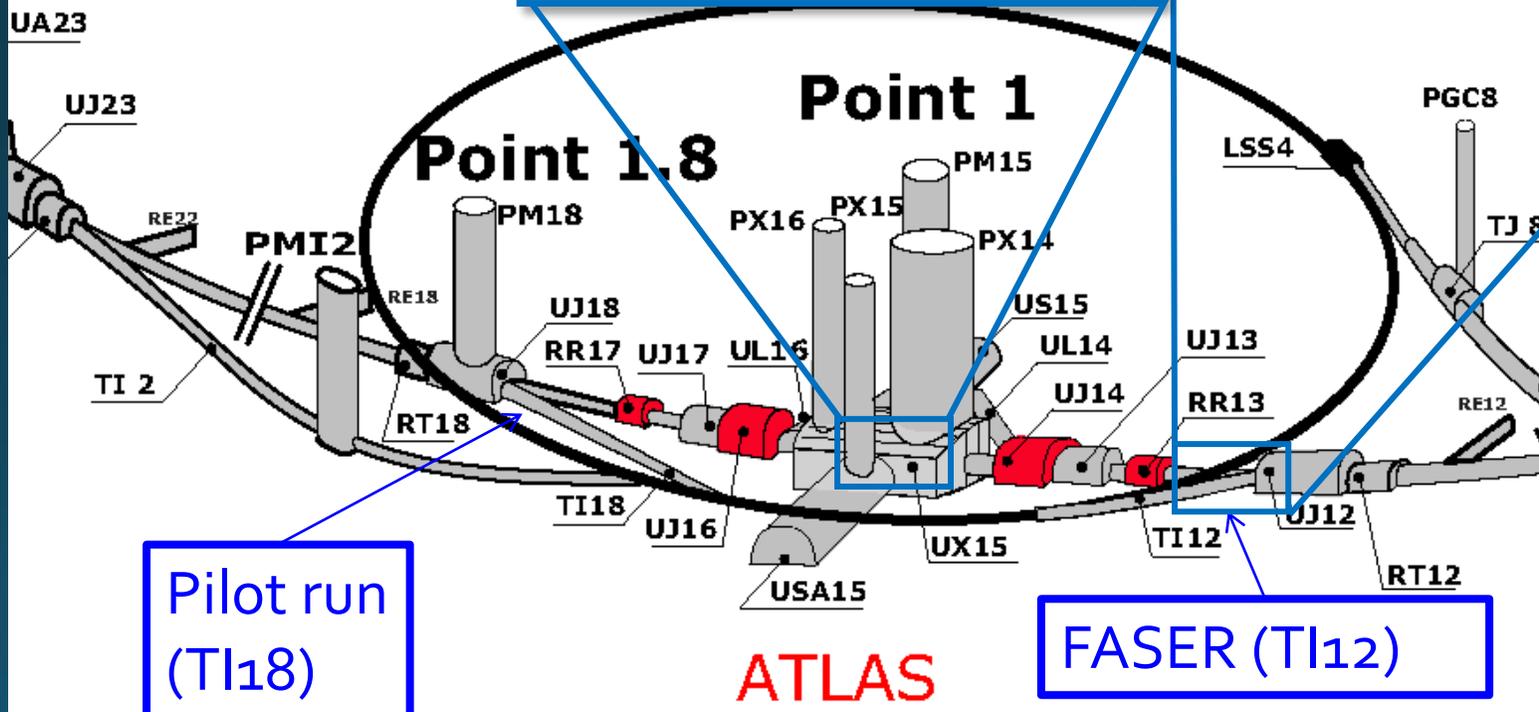
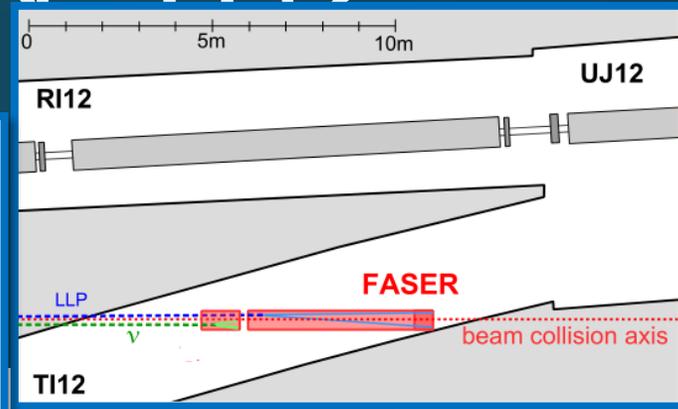
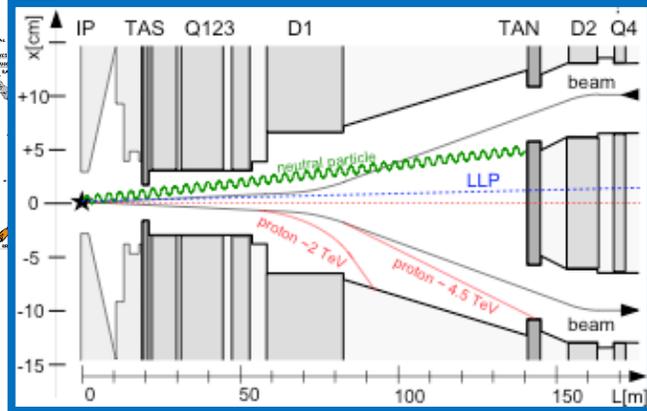
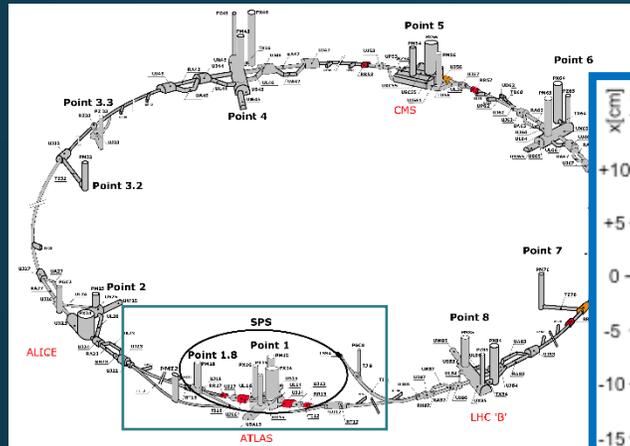
²Département de Physique Nucléaire et Corpusculaire,
University of Geneva, CH-1211 Geneva 4, Switzerland

³Universität Bern, Sidlerstrasse 5, CH-3012 Bern, Switzerland

⁴Kyushu University, Nishi-ku, 819-0395 Fukuoka, Japan

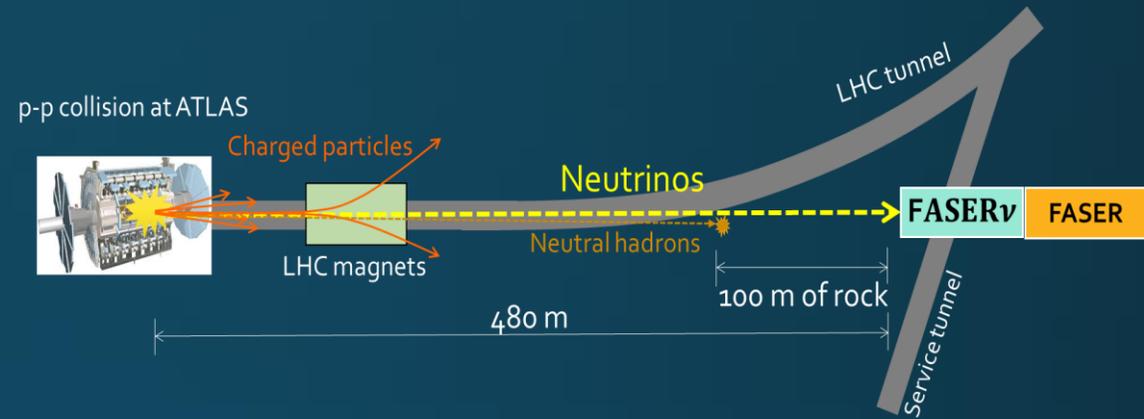
arXiv:1908.02310v1 [hep-ex] 6 Aug 2019

FASER LOCATION - TI12

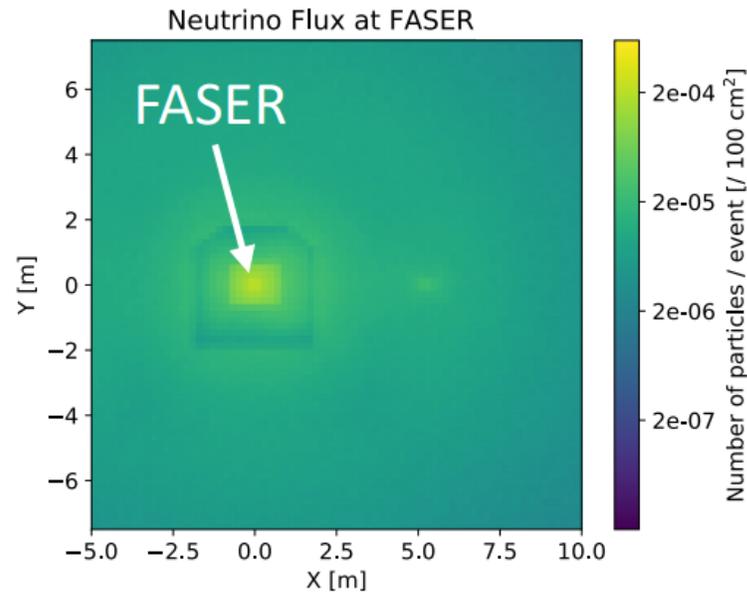
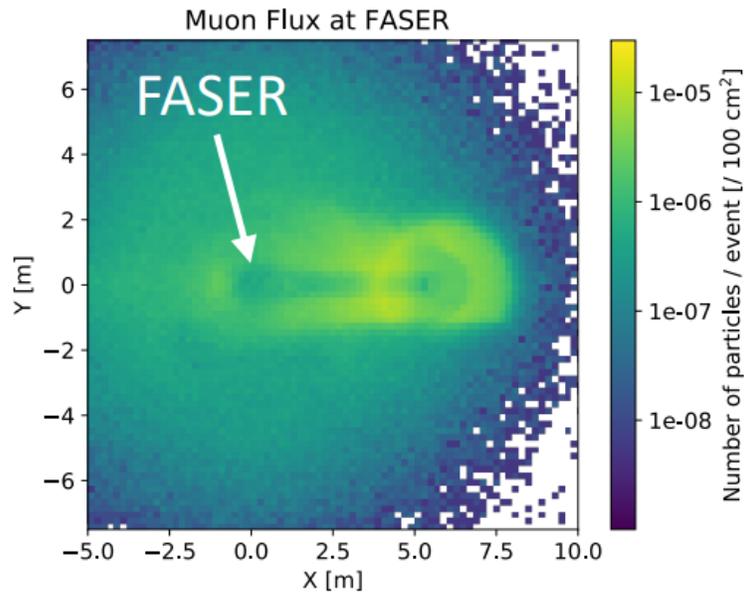


Particle fluence at the site

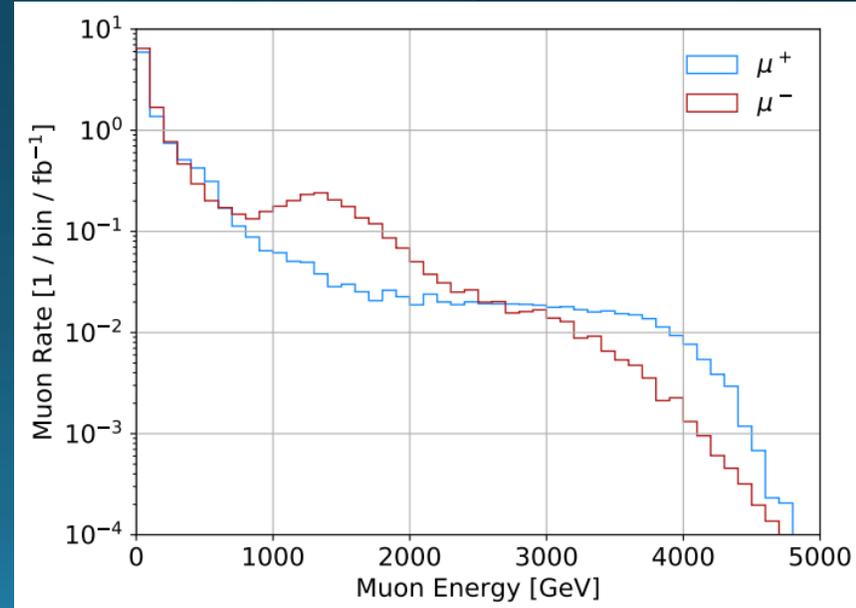
- Crucial for both neutrinos and LLP searches
- Simulation through the LHC infrastructures by FLUKA and BDSim
- Minimum muons, maximum neutrinos

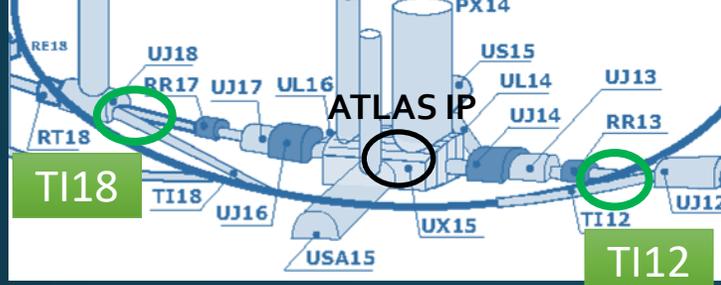


BDSim result for Tl12, Lefebvre ICHEP2020



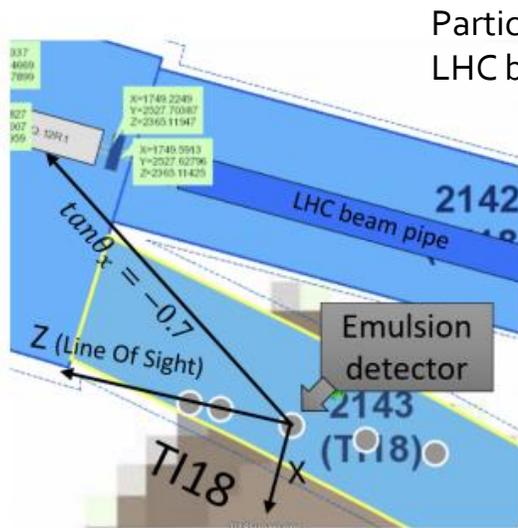
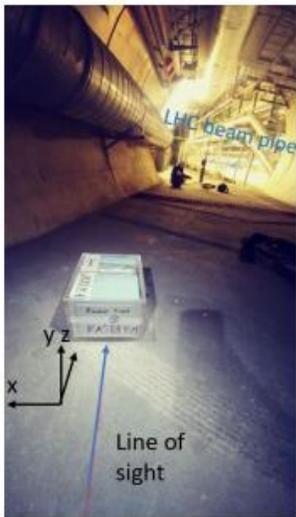
Muon energy (at 409m from IP, pilot run)
Simulated by CERN-STI group with FLUKA



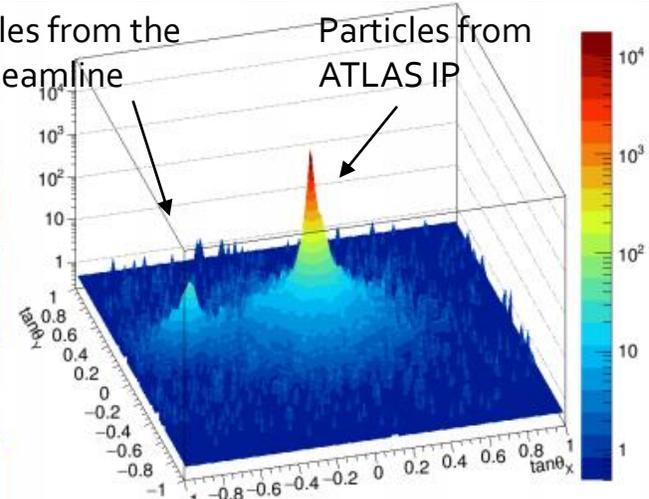


In situ measurements in 2018: Charged particle background

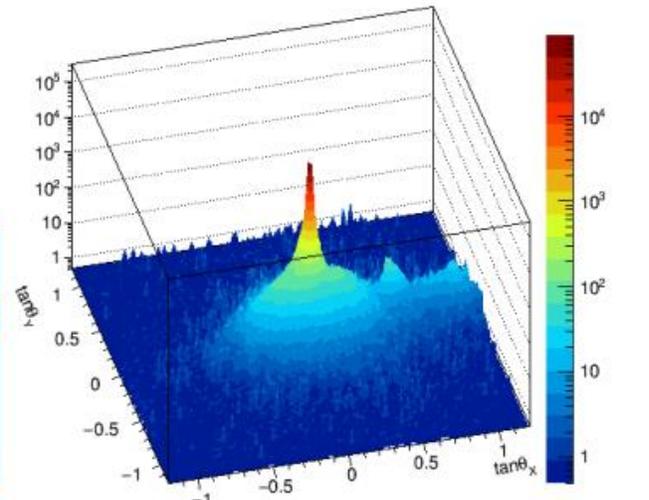
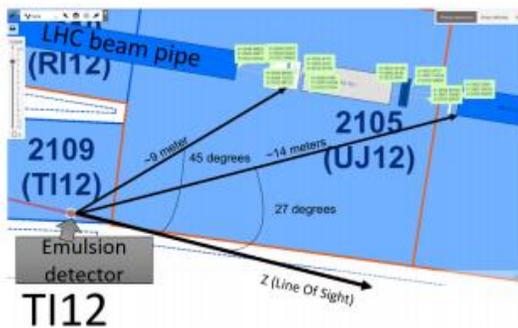
TI18



Particles from the LHC beamline
Particles from ATLAS IP



TI12

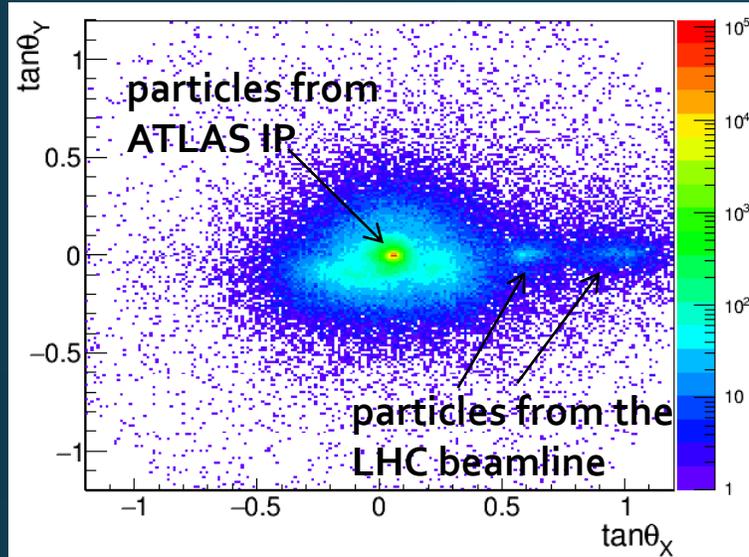


- Emulsion detectors were installed to investigate TI18 and TI12.
- Low background was confirmed.
- Few hadron tracks
- Consistent with the FLUKA prediction.

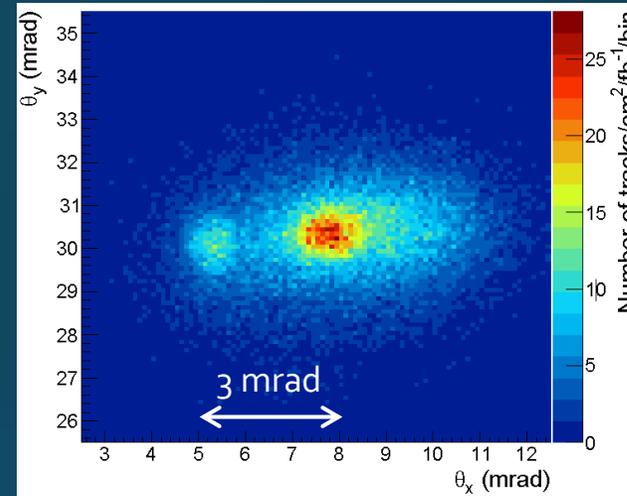
	Normalized flux (tracks/fb ⁻¹ /cm ²)
TI18	$(2.6 \pm 0.7) \times 10^4$
TI12	$(3.0 \pm 0.3) \times 10^4$

Emulsion detector can work at the actual environment!
(up to $\sim 10^6/\text{cm}^2 \approx 30 \text{ fb}^{-1}$ of data)

Angular distributions of beam backgrounds



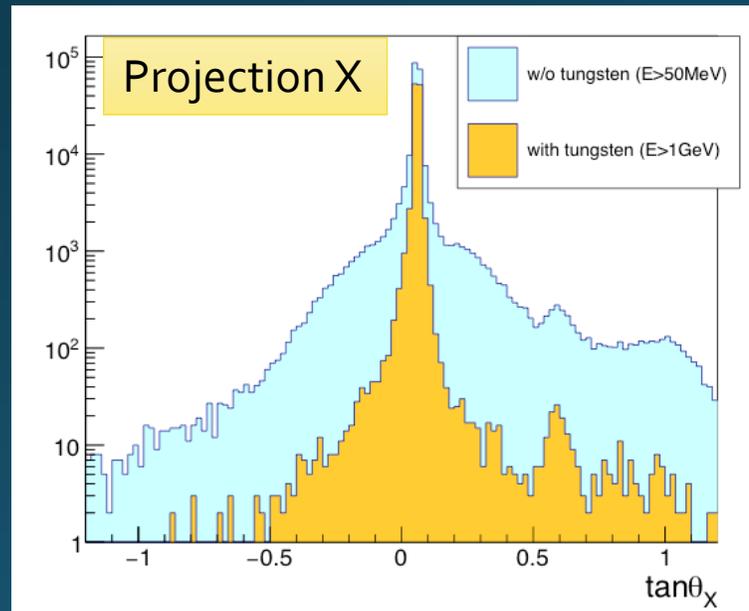
Close up to the main peak



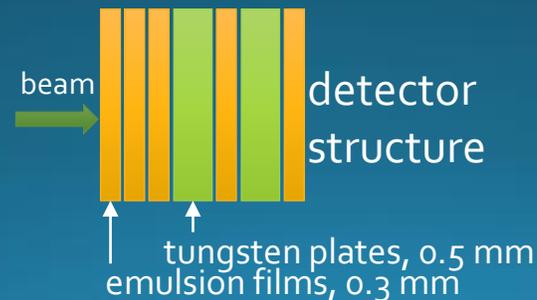
2 peak structure

$$\sigma = 0.6 \text{ mrad}$$

After 100 m of rock, it scatters only 0.6 mrad.
 → ~700 GeV



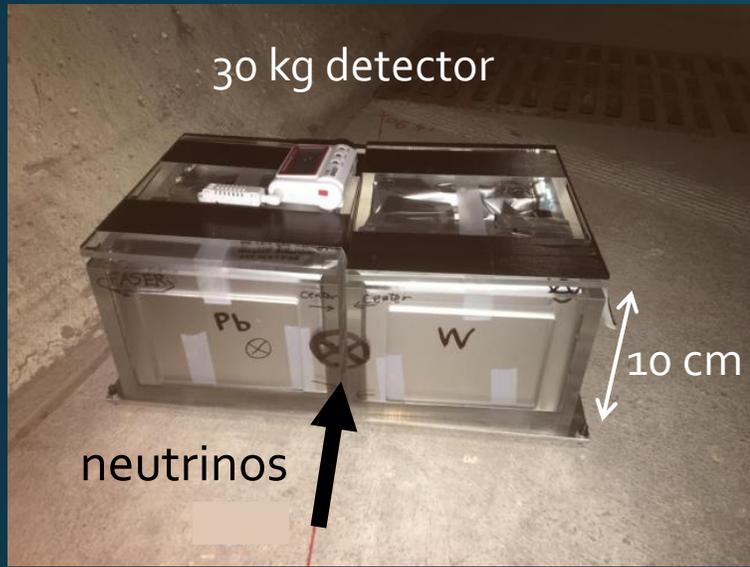
	Flux all [fb/cm ²]	Flux in main peak [fb/cm ²]
Tl18 data	$2.6 \pm 0.7 \times 10^4$	$1.2 \pm 0.4 \times 10^4$
Tl18 pilot		$1.7 \pm 0.1 \times 10^4$
Tl12 data	$3.0 \pm 0.3 \times 10^4$	$1.9 \pm 0.2 \times 10^4$
FLUKA MC		2.5×10^4



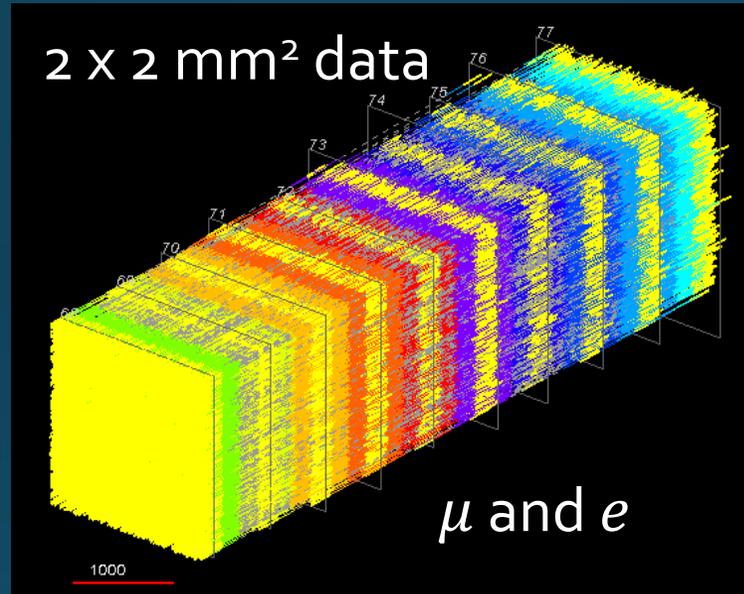
Data and the FLUKA (uncertainty 50%) prediction agrees within their uncertainties.

Pilot run in 2018

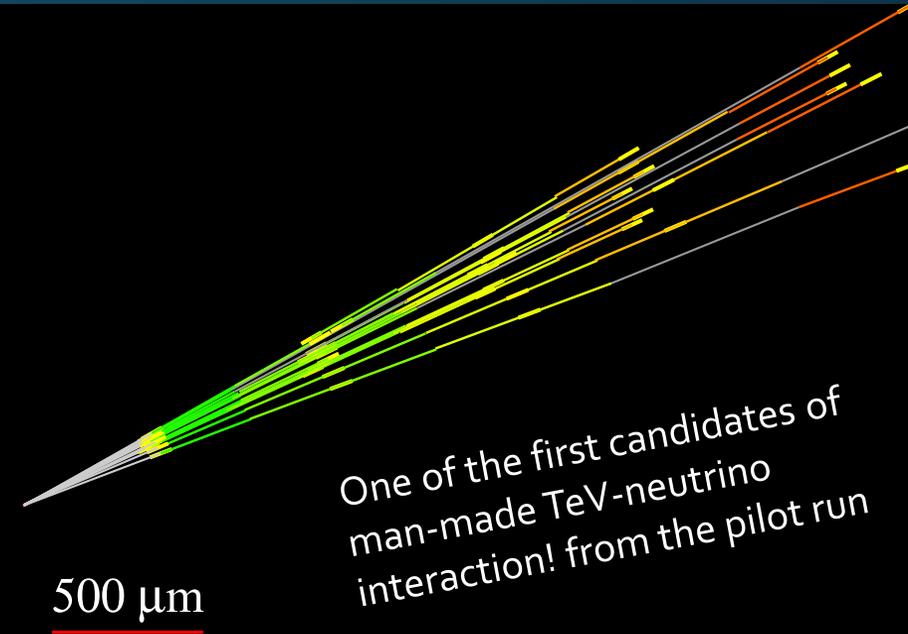
Aiming to demonstrate the feasibility of detection of collider neutrinos



6 weeks, 12.2 fb^{-1}

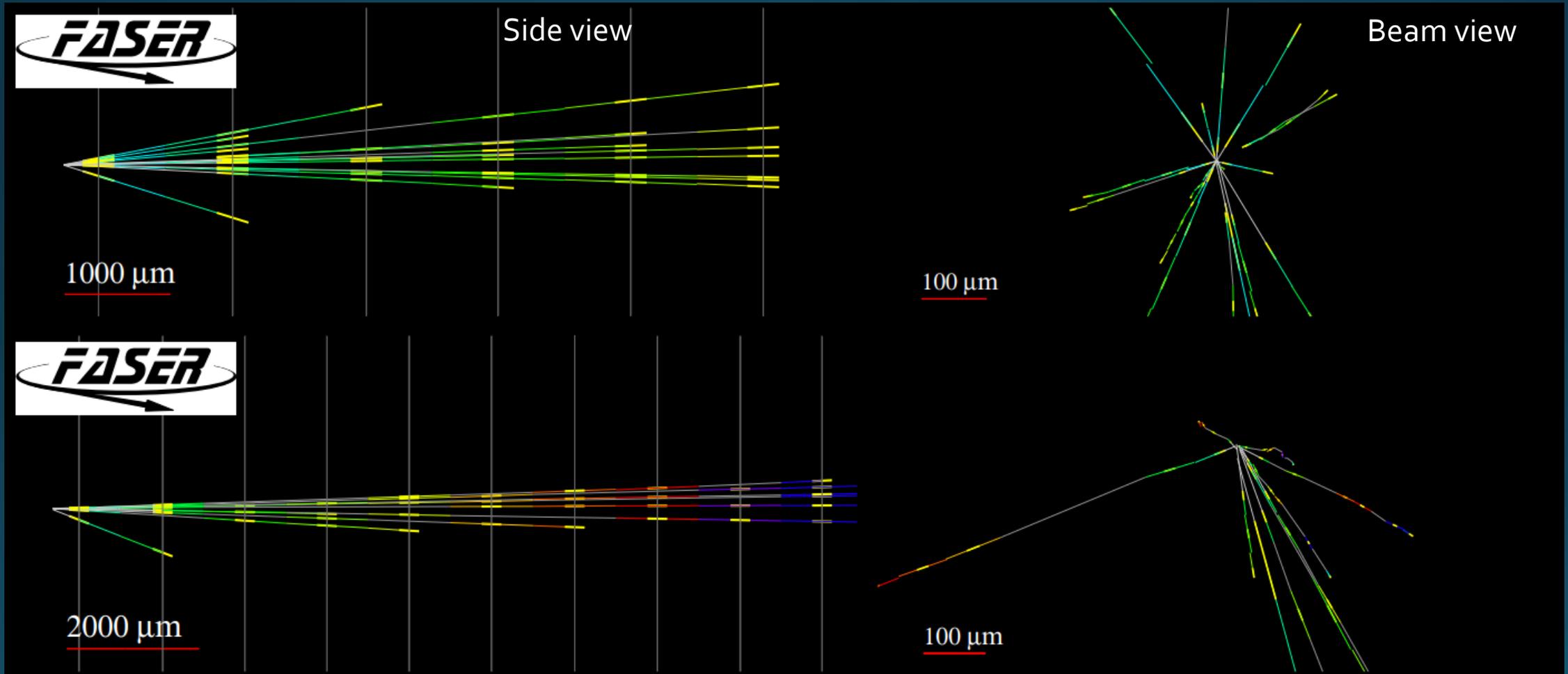


$\approx 3 \times 10^5 \text{ tracks/cm}^2$



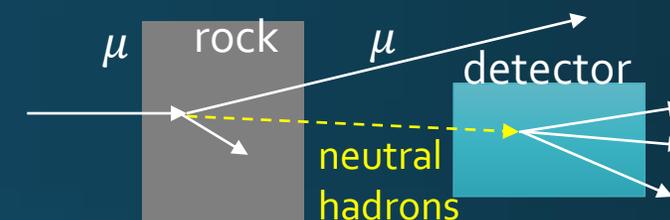
- A 30 kg emulsion based (lead, tungsten target) detector was installed on axis, 12.2 fb^{-1} of data was collected in Sep-Oct 2018 (6 weeks)

Neutrino interaction candidates



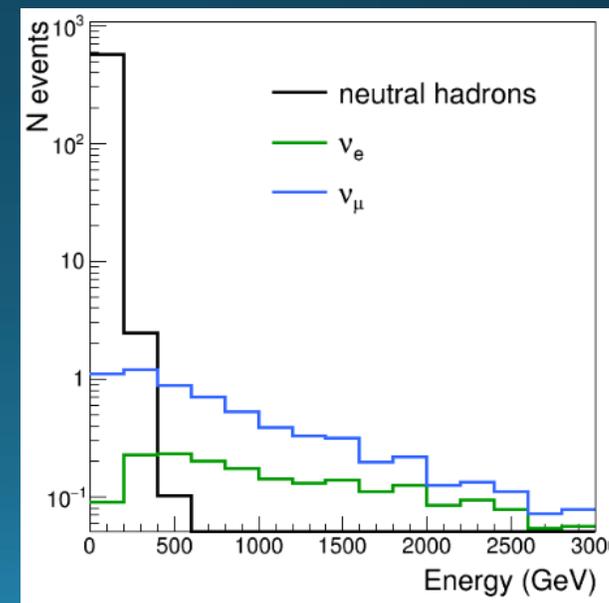
Background for neutrino analysis

- Muons rarely produce neutral hadrons in upstream rock or in detector, which can mimic neutrino interaction vertices
 - Probability of $O(10^{-5})$
- The produced neutral hadrons are low energy \rightarrow Discriminate by vertex topology
- (For physics run, Lepton ID will kill most of background)



	Negative Muons	Positive Muons
K_L	3.3×10^{-5}	9.4×10^{-6}
K_S	8.0×10^{-6}	2.3×10^{-6}
n	2.6×10^{-5}	7.7×10^{-6}
\bar{n}	1.1×10^{-5}	3.2×10^{-6}
Λ	3.5×10^{-6}	1.8×10^{-6}
$\bar{\Lambda}$	2.8×10^{-6}	8.7×10^{-7}

Production rate per muon ($E_{\text{had}} > 10 \text{ GeV}$)

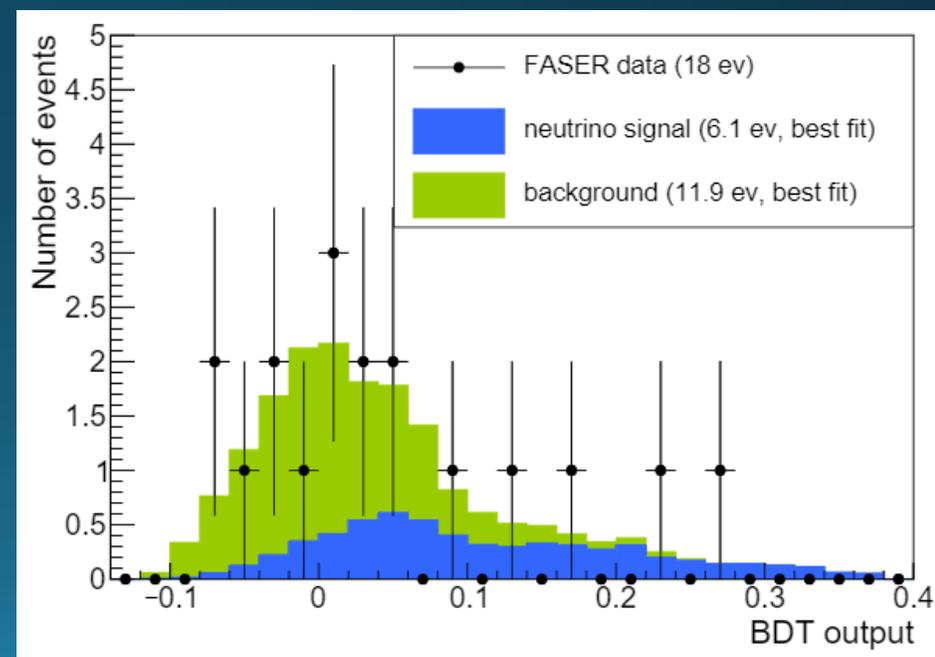


Pilot run event statistics

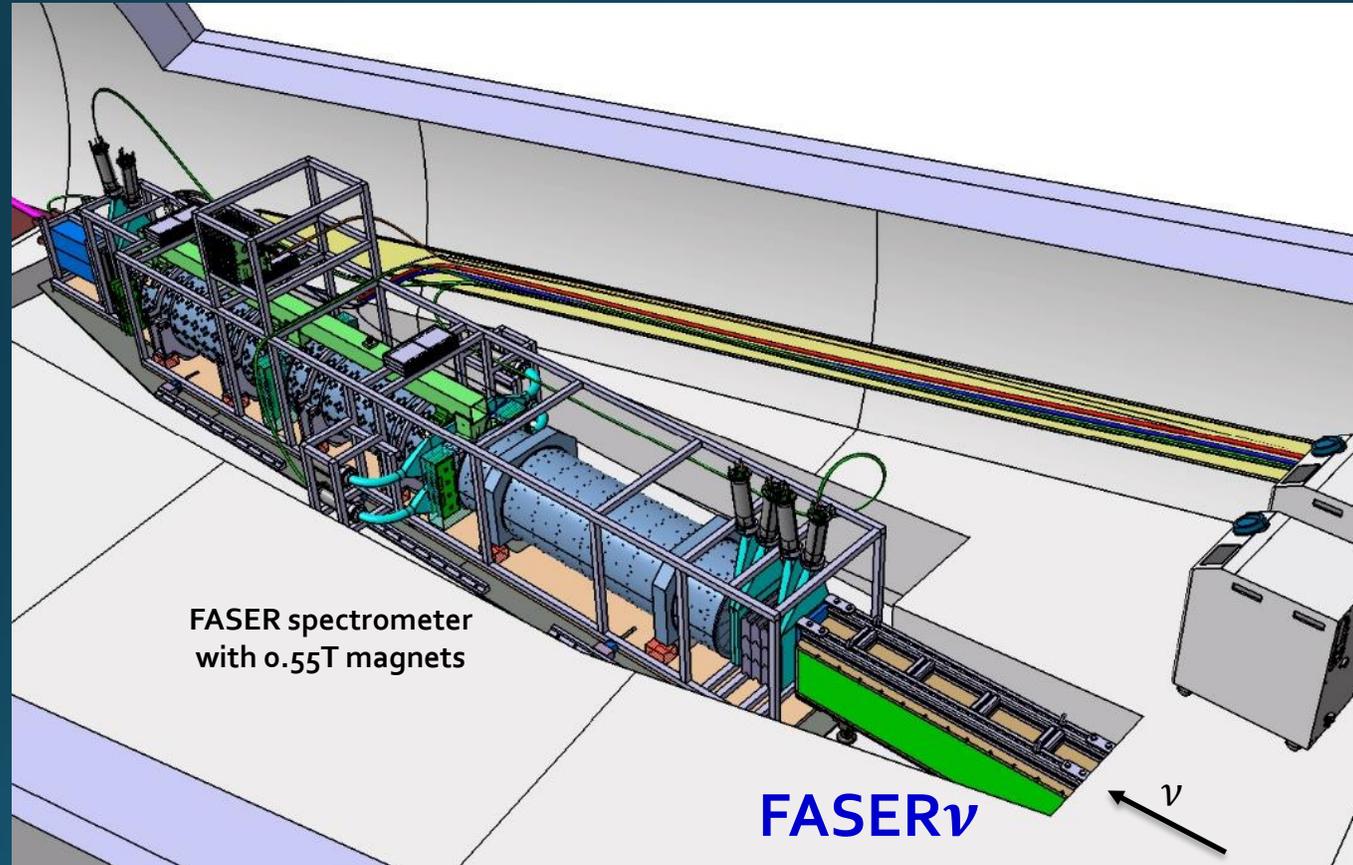
- Analyzed target mass of **11 kg**
- Pilot neutrino detector doesn't have lepton ID
 - Separation from neutral hadron BG (produced by muons) is challenging → tighter cuts
- Expected signal = $3.3^{+1.7}_{-0.95}$ events, BG = 11.0 events
- 18 neutral vertices were selected
 - by applying # of charged particle ≥ 5 , etc.
- In BDT analysis, an excess of neutrino signal is observed. Statistical significance = **2.7 sigma** from null hypothesis
- This result demonstrates the detection of neutrinos from the LHC

Vertex detection efficiency

Signal		Background		
		FTFP_BERT	QGSP_BERT	
ν_e	0.490	K_L	0.017	0.015
$\bar{\nu}_e$	0.343	K_S	0.037	0.031
ν_μ	0.377	n	0.011	0.012
$\bar{\nu}_\mu$	0.266	\bar{n}	0.013	0.013
ν_τ	0.454	Λ	0.020	0.021
$\bar{\nu}_\tau$	0.368	$\bar{\Lambda}$	0.018	0.018



Detector in the LHC Run3 (2021-2024)



FASER spectrometer
with 0.55T magnets

FASER ν

1.2 tons

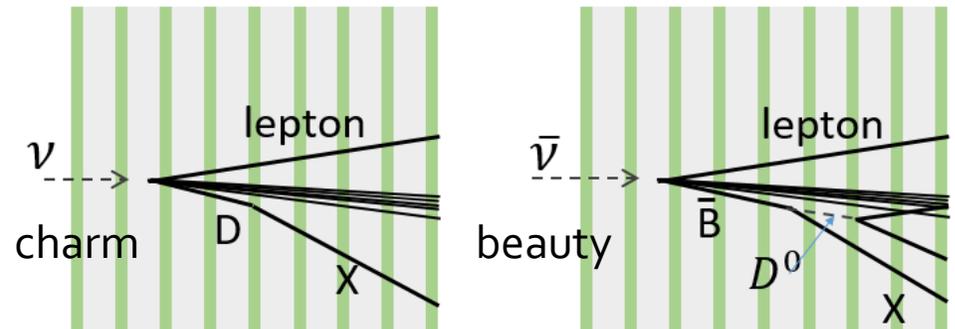
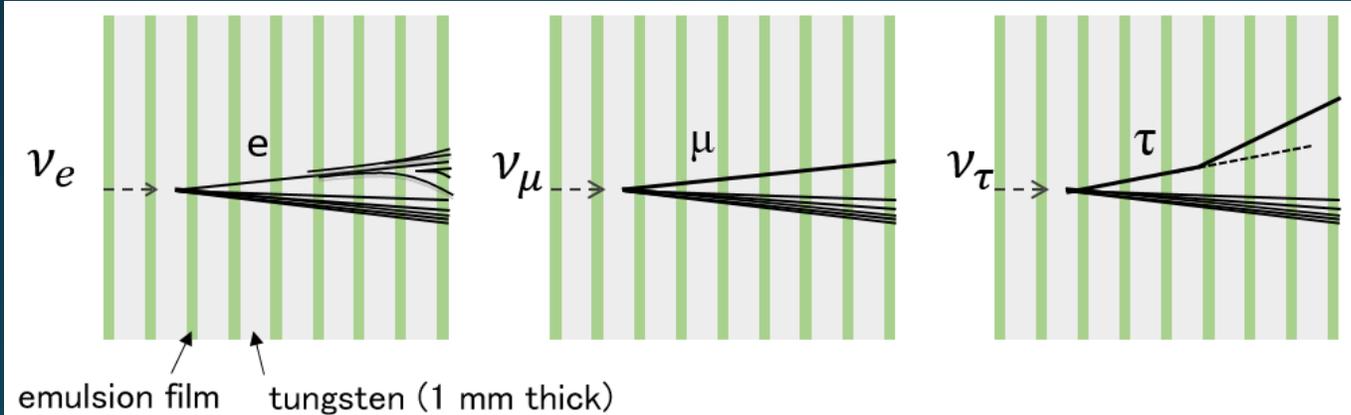
Conceptual detector design

Emulsion films + tungsten plates

ν
->

770 layers
25 cm x 25 cm x 1.1 m

1.2 tons, 220 X_0 , 8 λ_{int}



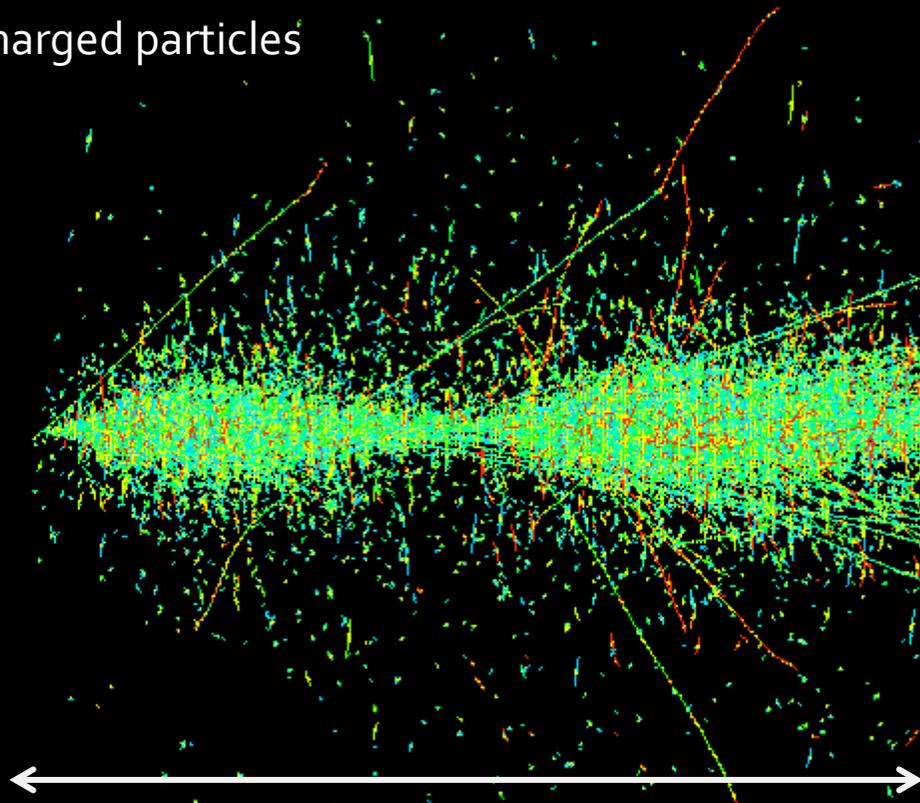
Exchange emulsions 9 times

2 0 2 2	3 ex	$\sim 80 \text{ fb}^{-1}$
2 0 2 3	3 ex	$\sim 80 \text{ fb}^{-1}$
2 0 2 4	3 ex	$\sim 80 \text{ fb}^{-1}$



Simulated 1 TeV ν_μ CC interaction

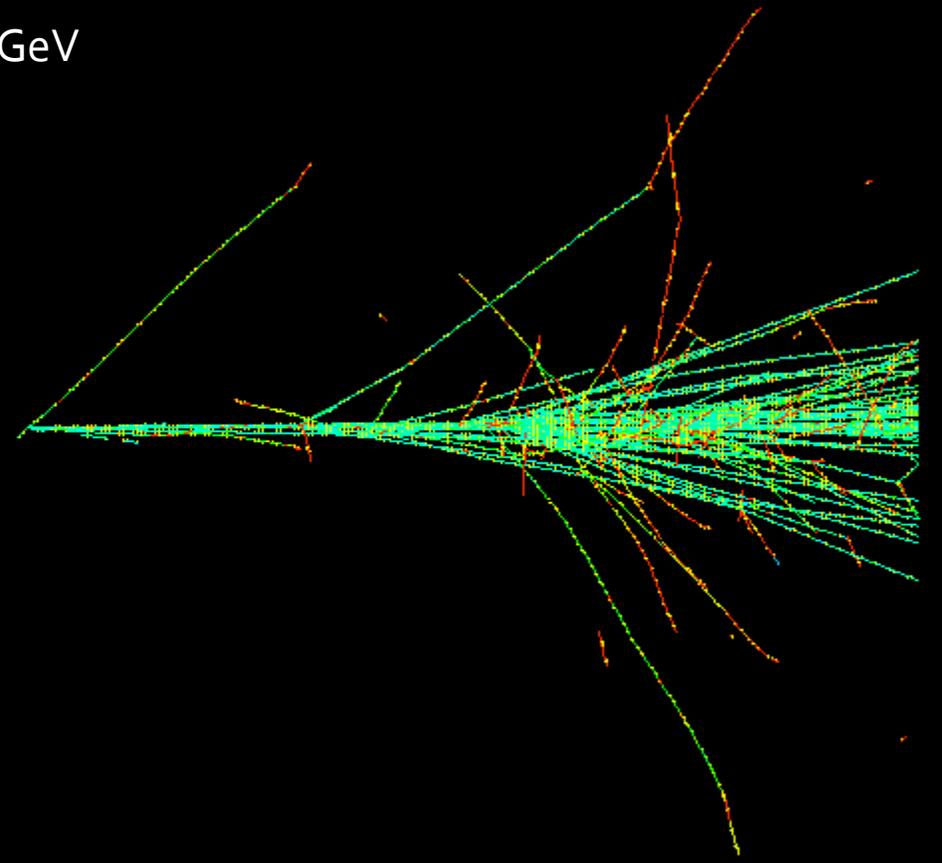
All charged particles



200 tungsten plates (27 cm)
 $\sim 57 X_0, \sim 2 \lambda_{int}$

50000 μm

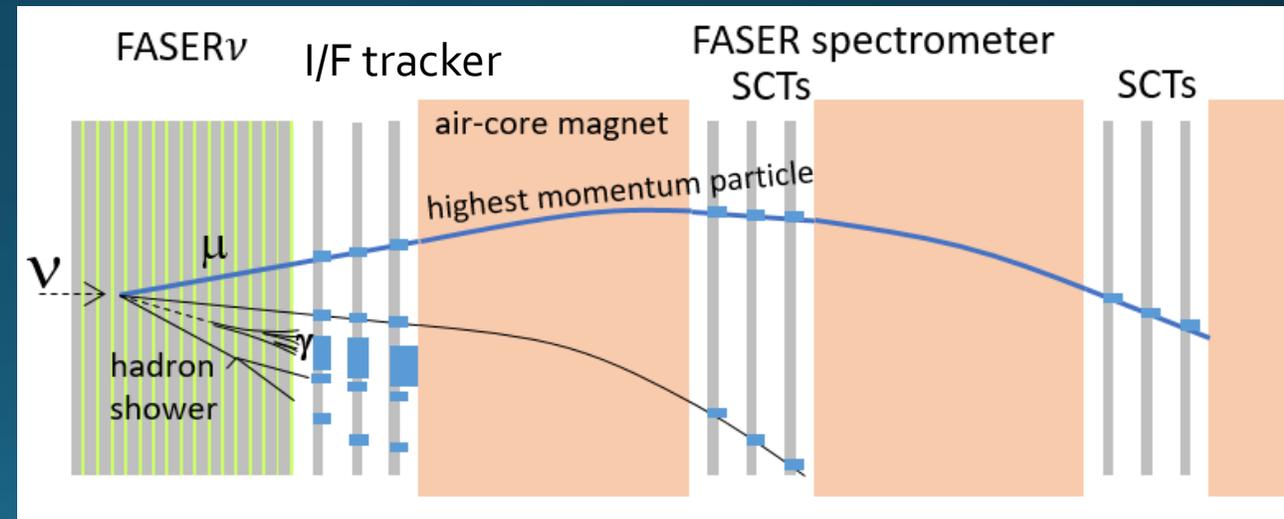
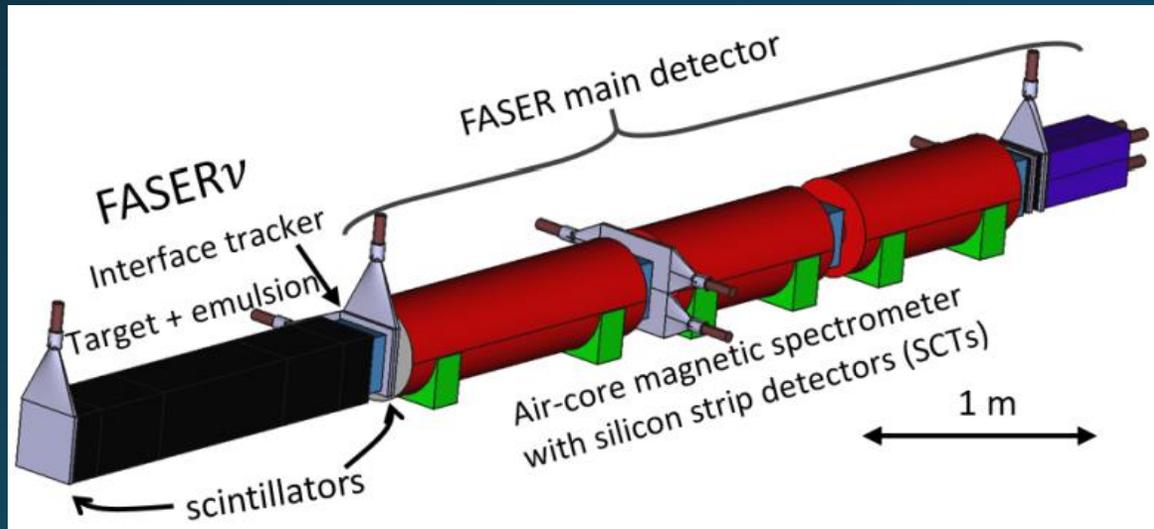
$P > 0.3$ GeV



50000 μm

FASER ν + FASER, hybrid configuration

- Muon charge identification
- Distinguish ν_μ and $\bar{\nu}_\mu \rightarrow$ Wider physics cases
- Improve neutrino energy reconstruction



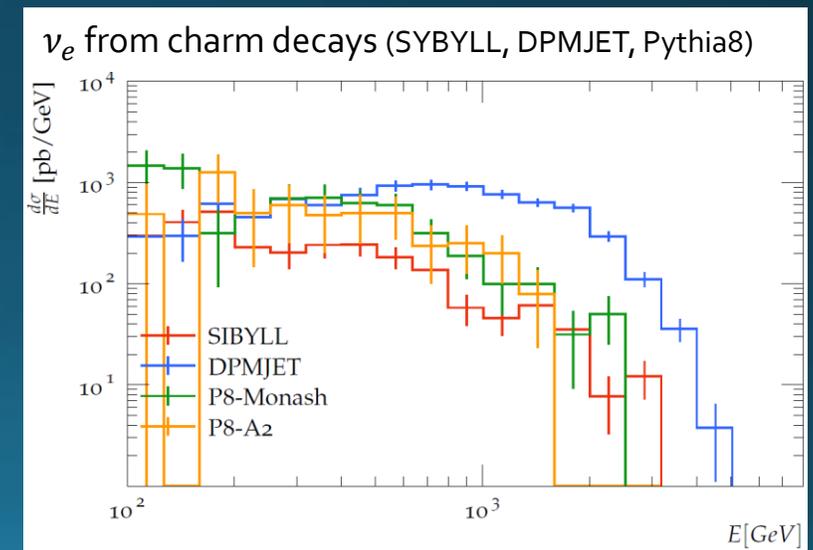
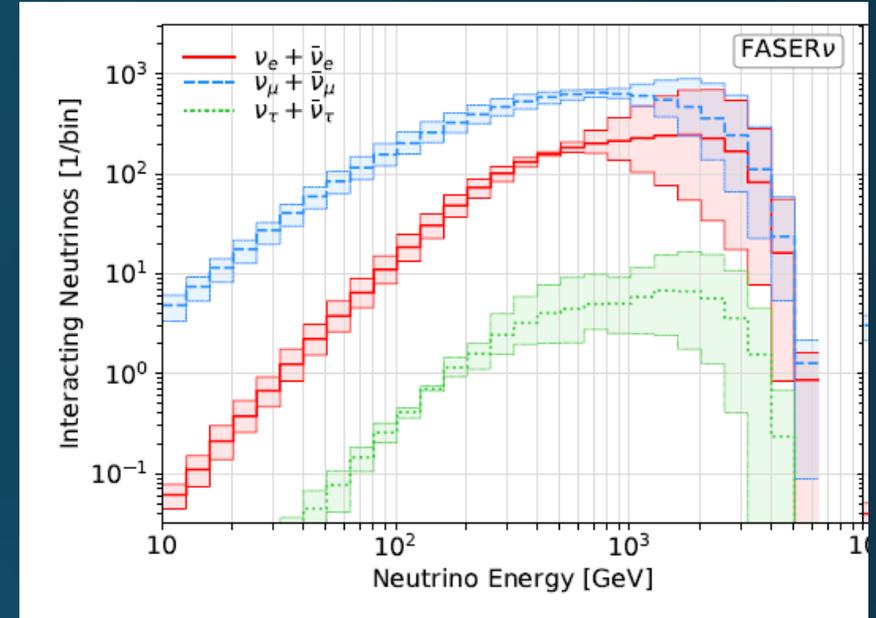
Neutrino event rate (2021-2024)

- **Small detector, but a lot of interactions ($\sim 10^4$ CC) are expected during Run3**
- **Neutrino fluxes are being cross-checked among different simulations**
 - Differences due to **hadron generators** and **beamline infrastructure reproduction** were identified. Currently, differences at hadron generators level is dominant

Expected number of CC interactions in FASER ν in Run3 (14 TeV LHC, 150 fb^{-1})

	SIBYLL	Pythia 8	DPMJET (used in FLUKA)
$\nu_e, \bar{\nu}_e$	800, 452	826, 477	3390, 1024
$\nu_\mu, \bar{\nu}_\mu$	6571, 1653	7120, 2178	8437, 2737
$\nu_\tau, \bar{\nu}_\tau$	16, 6	22, 11	111, 43

- **Work in progress for quantifying and reducing these uncertainties**
 - Creating a dedicated forward physics tune with Pythia8, using forward data (LHCf, FASER's muon measurements, etc.)



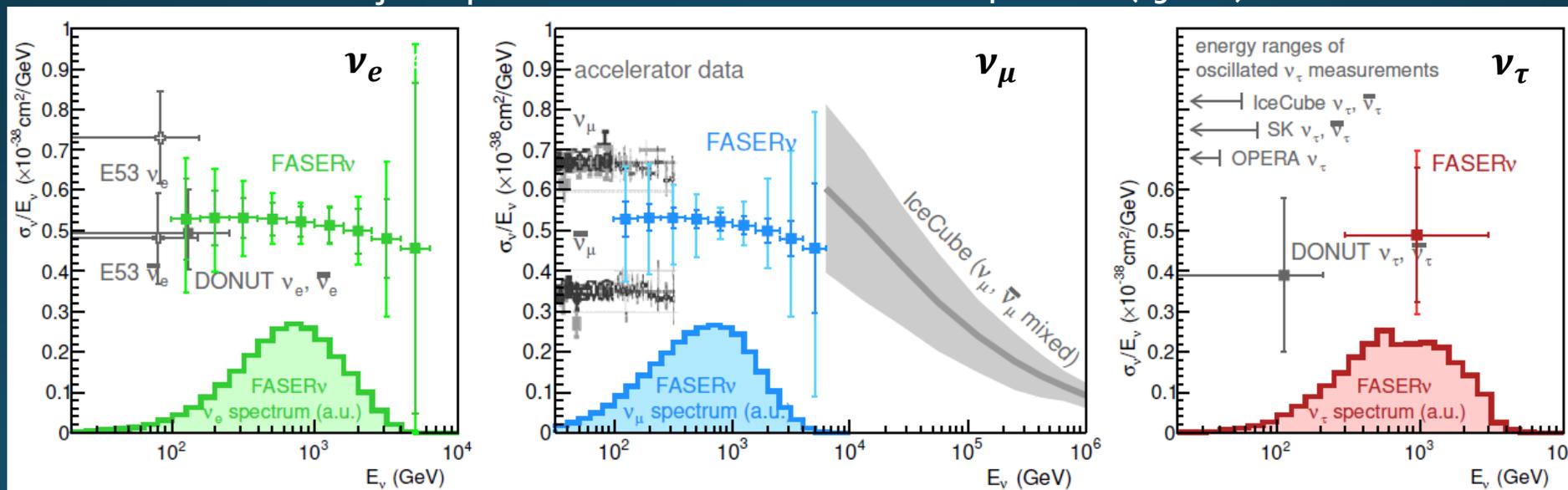
Large variation between different hadron production models (at p-p collision)

Physics studies in the LHC Run 3 (1): Cross sections

FASER Collaboration,
Eur. Phys. J. C 80 (2020) 61,
arXiv:1908.02310

- Neutrino cross section measurement at unexplored energy range
 - ν_e, ν_τ at the highest energy
 - Fill the gap between accelerator and cosmic data for ν_μ

Projected precision of FASER ν measurement at 14-TeV LHC (150 fb $^{-1}$)

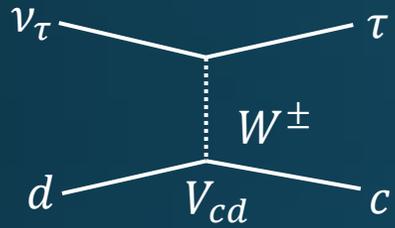


inner error bars: statistical uncertainties, outer error bars: uncertainties from neutrino production rate corresponding to the range of predictions obtained from different MC generators.

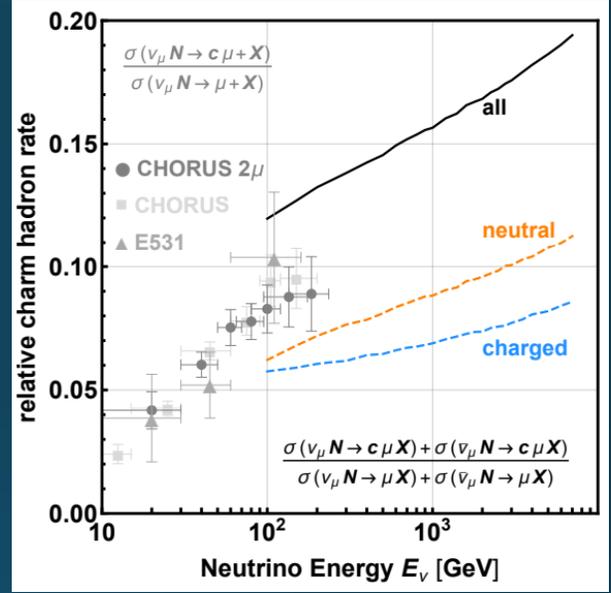
Physics studies in the LHC Run 3 (2):

Heavy-flavor-associated channels

- **Measure charm** production channels
 - Large rate $\sim 10\%$ ν CC events, $\mathcal{O}(1000)$ events
 - First measurement of ν_e induced charm prod.

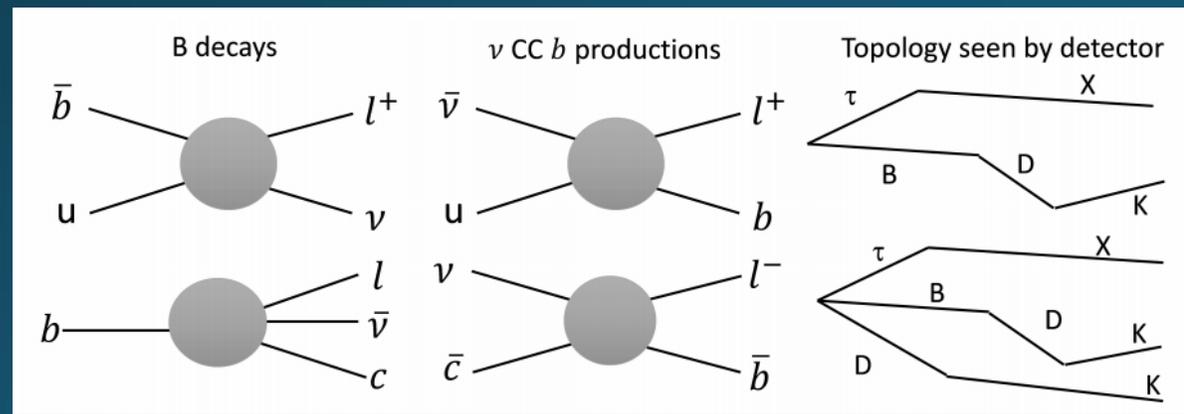


$$\frac{\sigma(\nu_\ell N \rightarrow \ell X_c + X)}{\sigma(\nu_\ell N \rightarrow \ell + X)} \quad \ell = e, \mu$$



- **Search for Beauty** production channels

- Expected SM events (ν_μ CC b production) are $\mathcal{O}(0.1)$ events in Run 3, due to CKM suppression, $V_{ub}^2 \approx 10^{-5}$



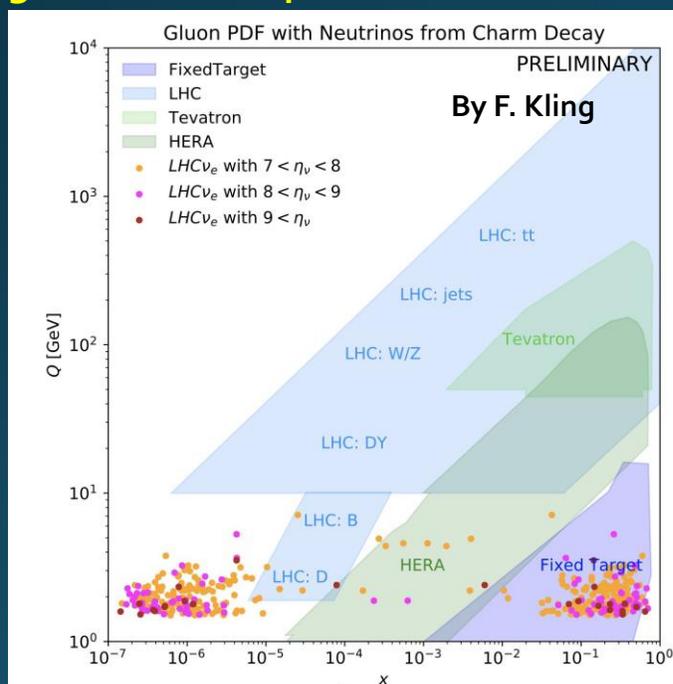
$$\bar{\nu}N \rightarrow \ell \bar{B}X$$

$$\nu N \rightarrow \ell BDX$$

Physics studies in the LHC Run 3 (3): QCD

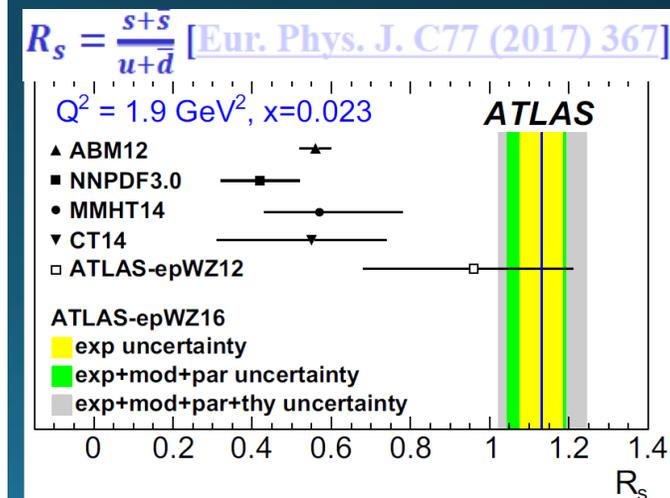
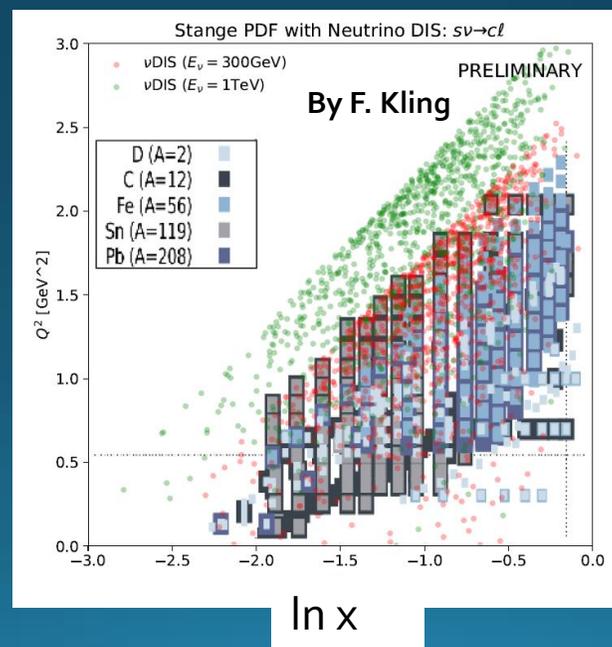
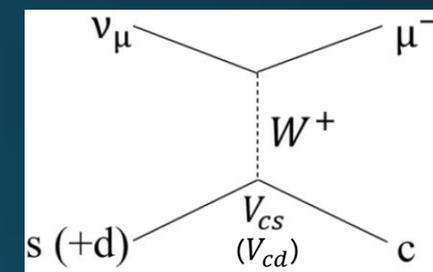
PDF in proton (neutrino production)

- Forward particle production is poorly constrained by other LHC experiments. FASER ν 's **neutrinos flux measurements** will provide novel complimentary constraints that can be used to validate/improve MC generators.
- Neutrinos from charm decay could allow to **test transition to small-x factorization, constrain low-x gluon PDF and probe intrinsic charm.**



PDF in target (neutrino interaction)

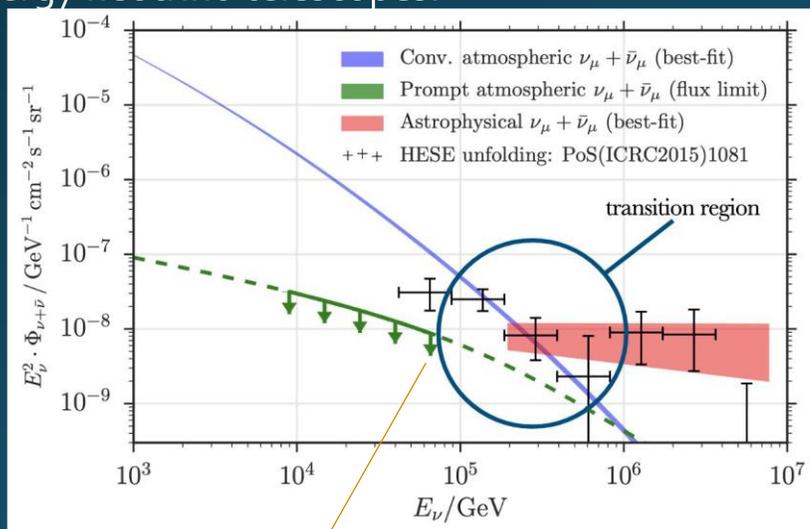
- It is also interesting to probe (nuclear) PDFs via DIS neutrino scattering. In particular, **charm associated neutrino events ($\nu s \rightarrow l c$) are sensitive to the poorly constrained strange quark PDF.**



Physics studies in the LHC Run 3 (4): Cosmic rays and neutrino

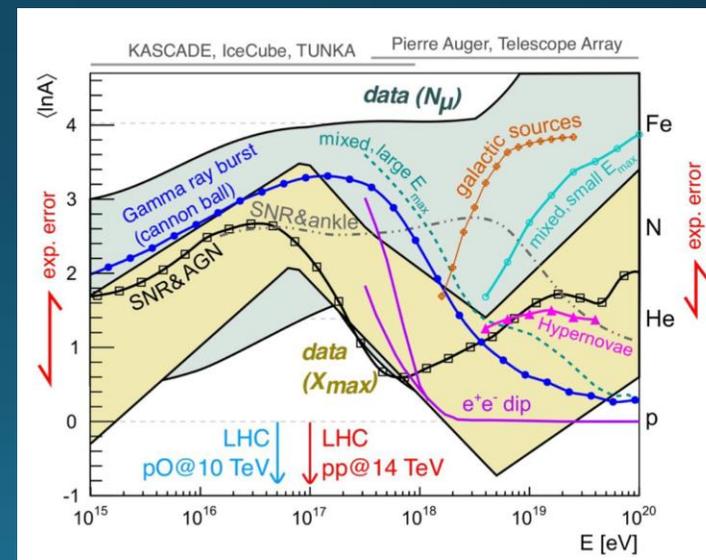
- In order for IceCube **to make precise measurements of the cosmic neutrino flux**, accelerator measurements of high energy and large rapidity charm production are needed.
- As 7+7 TeV p - p collision corresponds to 100 PeV proton interaction in fixed target mode, a direct **measurement of the prompt neutrino production at FASER ν** would provide important basic data for current and future high-energy neutrino telescopes.

- Muon problem in CR physics: **cosmic ray experiments have reported an excess in the number of muons** over expectations computed using extrapolations of hadronic interaction models tuned to LHC data at the few σ level. **New input from LHC is crucial to reproduce CR data consistently.**



prompt atmospheric neutrinos

IceCube Collaboration,
Astrophys. J. 833 (2016)

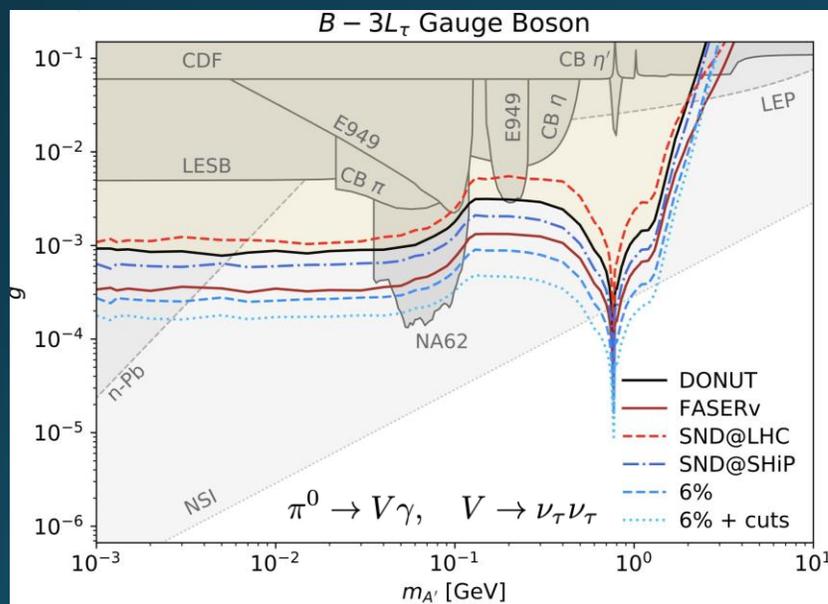


K.H. Kampert, M. Unger, *Astropart. Phys.* 35, 660 (2012),
H.P. Dembinski et al., *EPJ Web Conf.* 210, 02004 (2019)

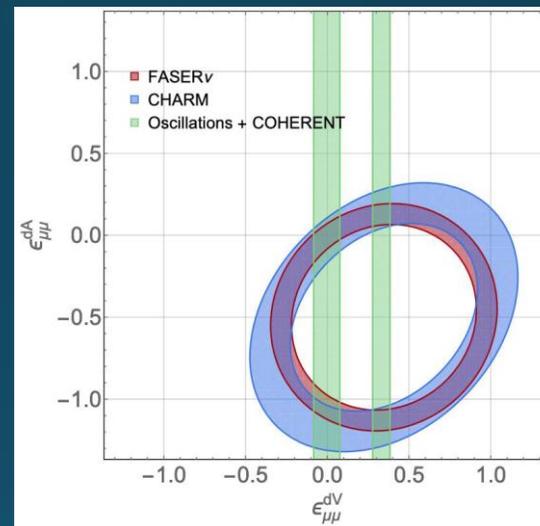
Physics studies in the LHC Run 3 (5): BSM Physics

- The tau neutrino flux is small in SM. A **new light weakly coupled gauge bosons** decaying into tau neutrinos could significantly enhance the tau neutrino flux.

F. Kling, Phys. Rev. D 102, 015007 (2020), arXiv:2005.03594



- NC measurements at FASER ν could constrain **neutrino non-standard interactions (NSI)**.



A. Ismail, R.M. Abraham, F. Kling, arXiv: 2012.10500

- Sterile neutrinos** with mass ~ 40 eV can cause oscillations at FASER ν and the spectrum deformation may be seen.

FASER Collaboration, Eur. Phys. J. C 80 (2020) 61, arXiv:1908.02310

- If DM is light, the LHC can produce an energetic and collimated DM beam towards FASER ν . FASER ν could also search for **DM scattering**.

B. Batell, J. Feng, S. Trojanowski, 2020, in preparation

Emulsion detector preparation

- Emulsion gel and film production facilities in Nagoya have been set up in 2020. We are testing mass production
- Chemical compatibility of tungsten plates with emulsion film were tested



Tungsten plate



Emulsion film

Solution tanks



Crystal formation tank

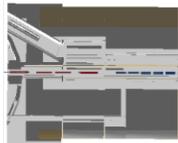


Film production facility

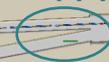


Experimental site

ATLAS



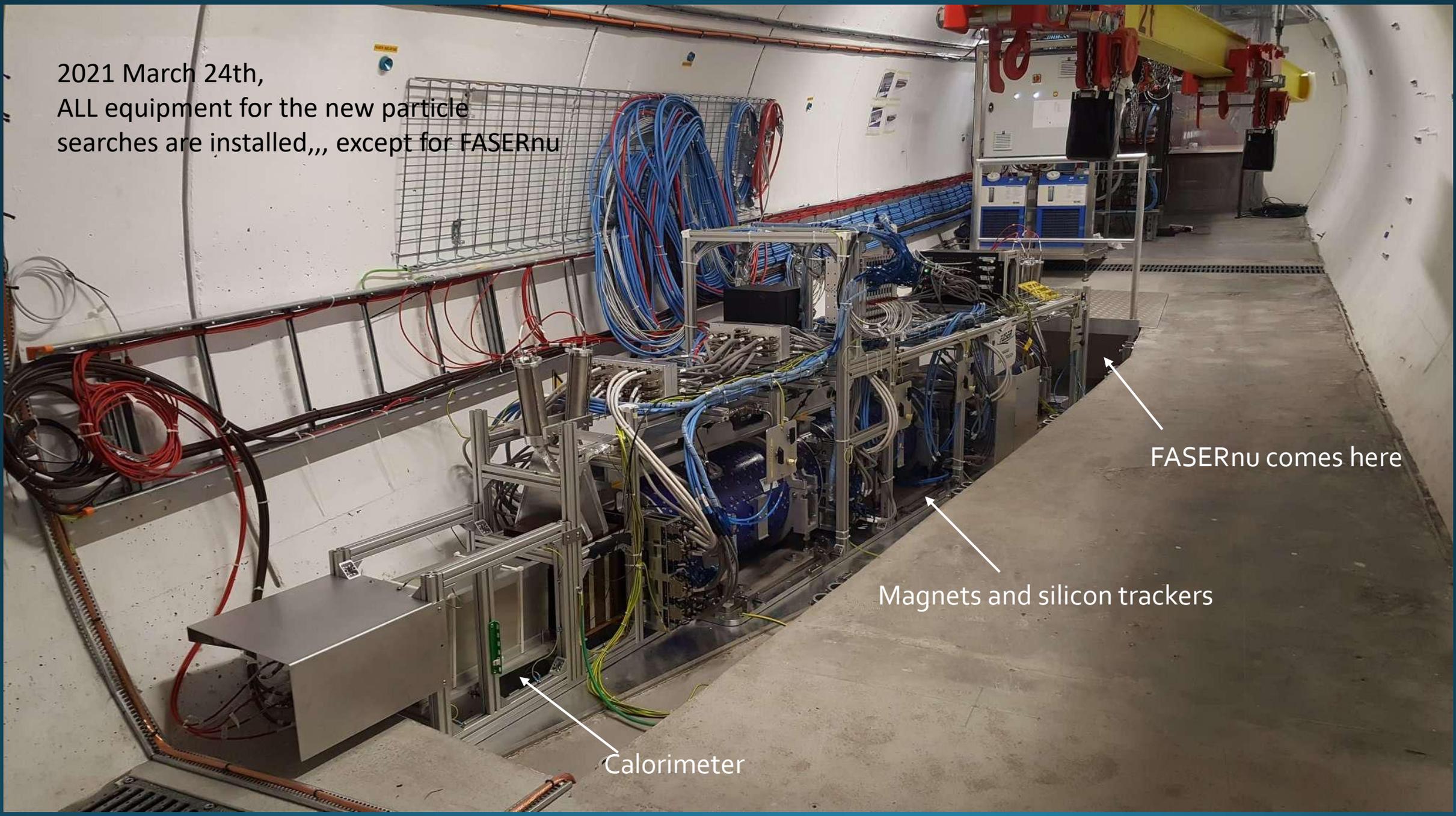
FASER



Evolution of T112 tunnel for FASER installation



2021 March 24th,
ALL equipment for the new particle
searches are installed,,, except for FASERnu

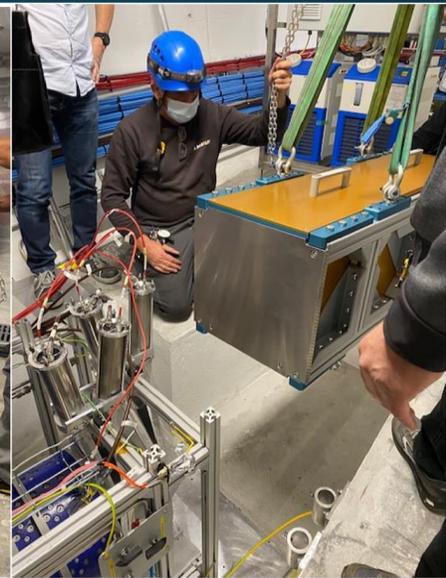


FASERnu comes here

Magnets and silicon trackers

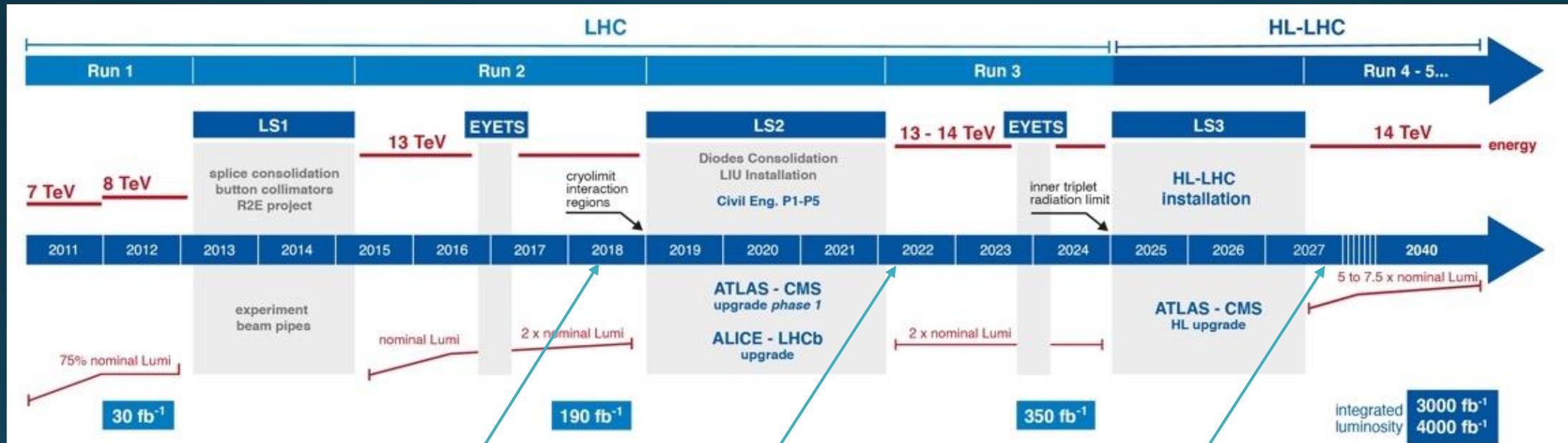
Calorimeter

FASER ν installation test in April 2021



LHC Schedule

- LHC Run-3 will start in 2022, aiming to double the integrated luminosity
- HL-LHC, starting in 2027, will deliver 10 times more integrated luminosity



BG measurement,
pilot run in 2018

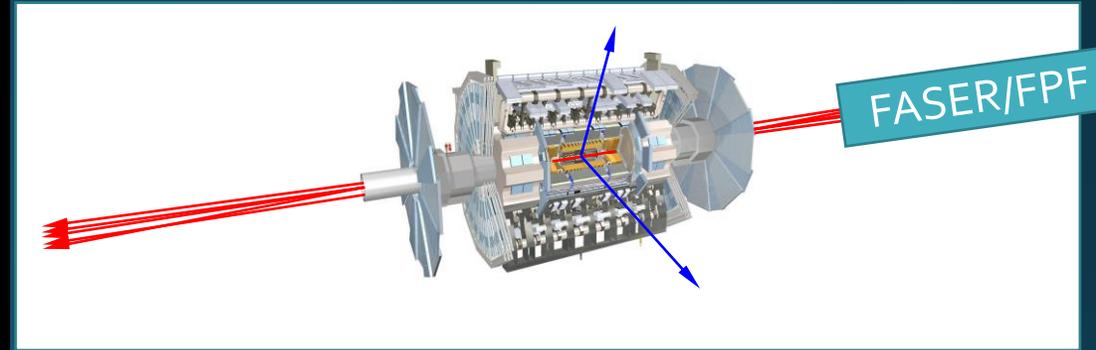
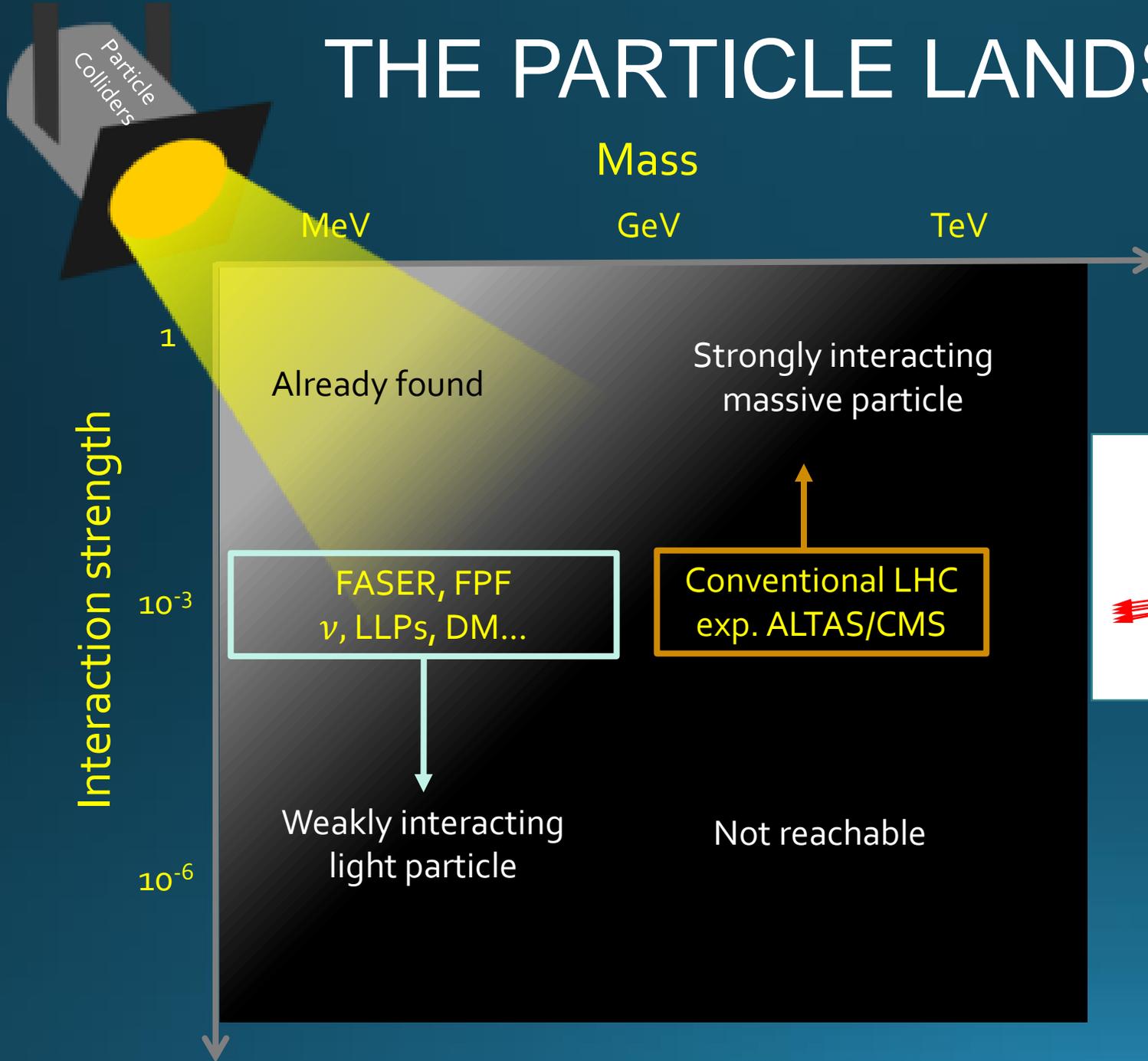
Physics run will start in
2022 (~150 fb⁻¹)

Forward experiment
in HL-LHC

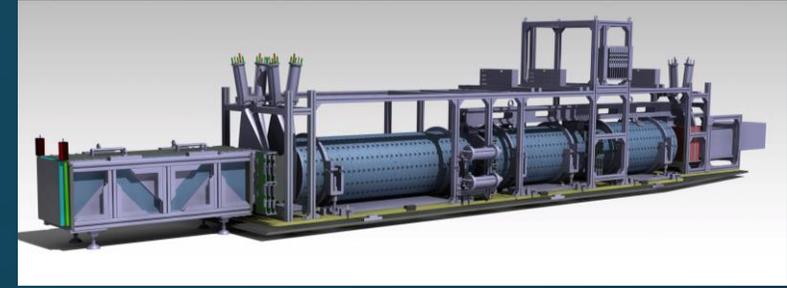
Motivation to Forward Physics Facility (FPF)

- LHC is currently the high energy frontier, and in next 15 years.
- The high luminosity run (HL-LHC) will start in 2027. What is the best way to exploit it?
 - Conventional LHC exps (ATLAS, CMS) studies “Physics with high Pt, small cross sections (fb, pb, nb)
 - However, the total cross section is **100 mb**, mostly at **far forward direction** (small Pt). Why not to use this abundant events?
- Far forward physics = unexplored physics domain, but explorable with a relatively small investment thanks to the existence of the LHC, as pioneered by FASER
- Proposal: “Let’s build a **Forward Physics Facility** and host variety of experiments”
 - SM: tau neutrino, QCD, cosmic ray
 - BSM: LLPs, FIPs, dark sector particles, milli-charged particles

THE PARTICLE LANDSCAPE

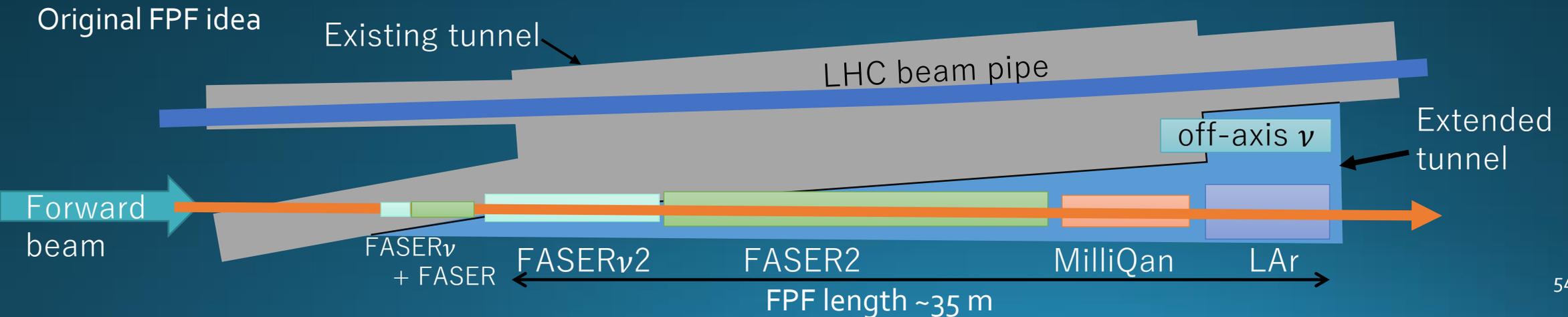


Idea of FPF



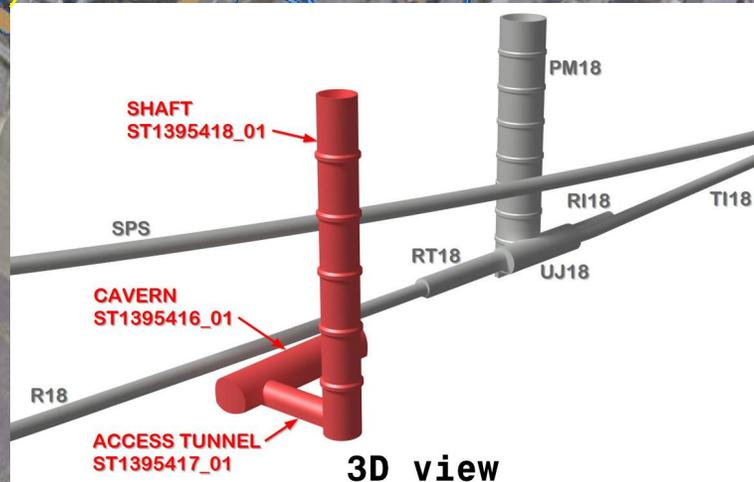
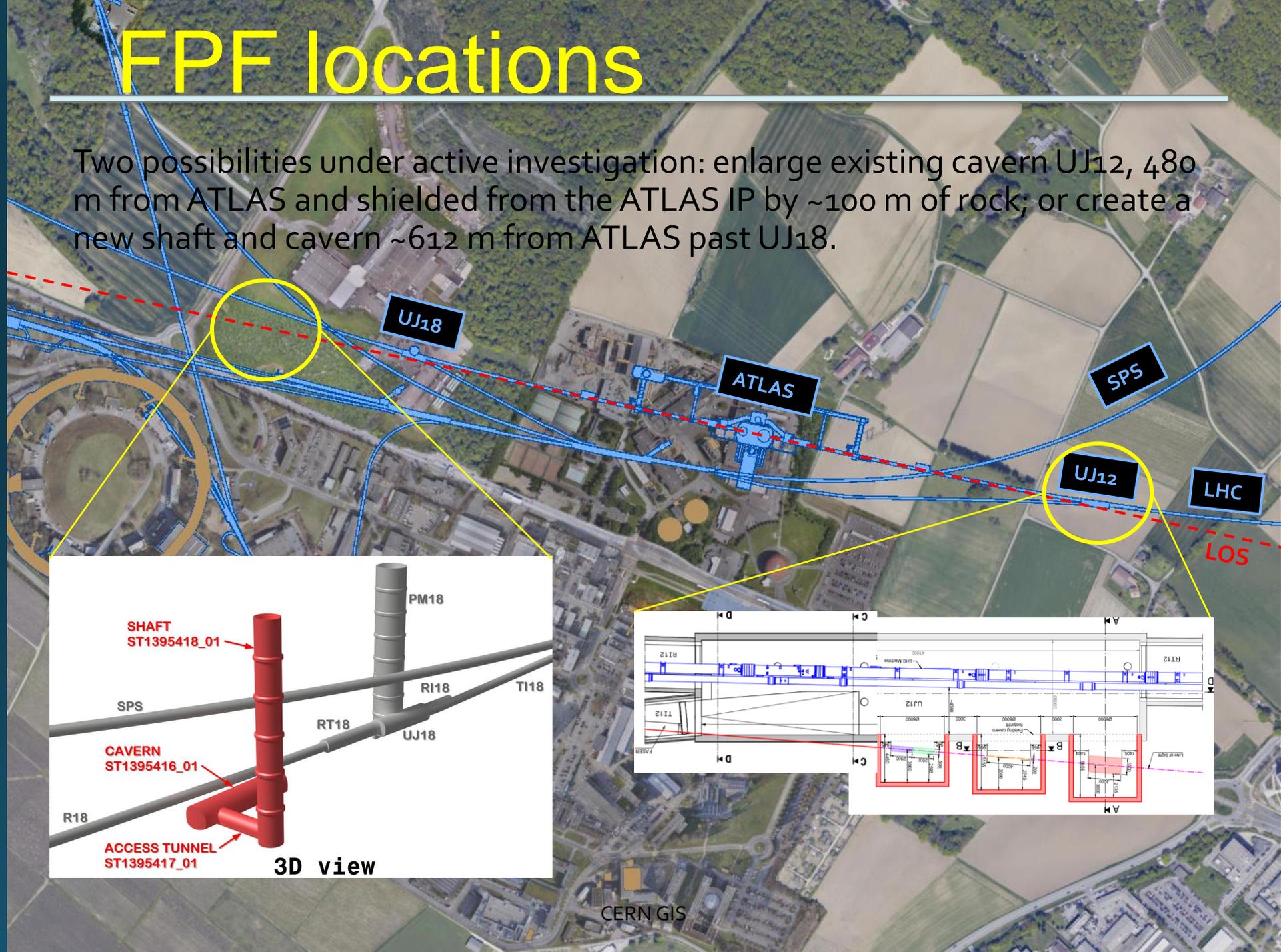
FASER

- Multiple single-purpose detectors
 - LLPs: FASER2
 - Neutrino, FIP: FASER ν 2 (on-axis), SND2 (off-axis), LAr
 - Milli-charged particle : MilliQan
- An experimental hall to host these experiments → **Forward Physics Facility (FPF)**
- Neutrino experiment with x200 statistics (x10 detector x20 beam)
 - Focus on **tau neutrinos**



FPF locations

Two possibilities under active investigation: enlarge existing cavern UJ12, 480 m from ATLAS and shielded from the ATLAS IP by ~100 m of rock; or create a new shaft and cavern ~612 m from ATLAS past UJ18.



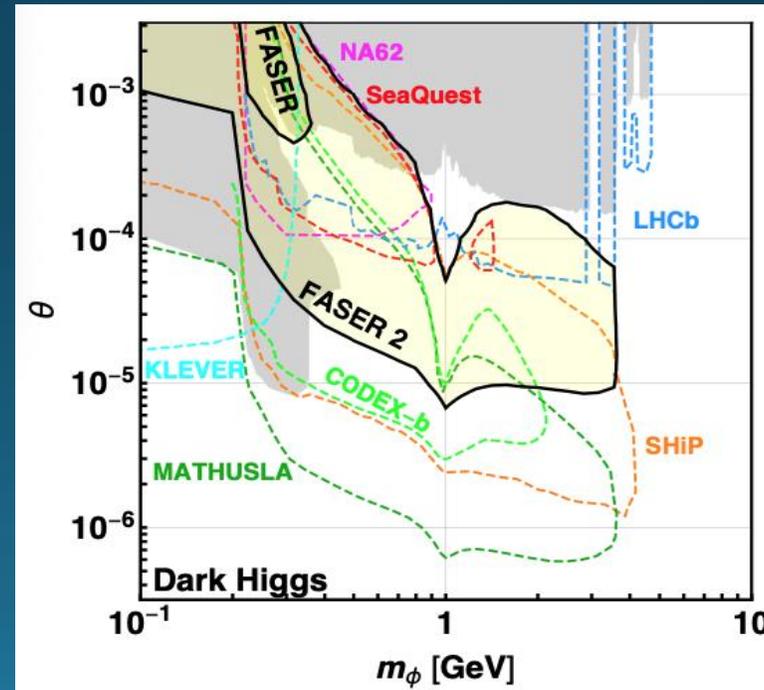
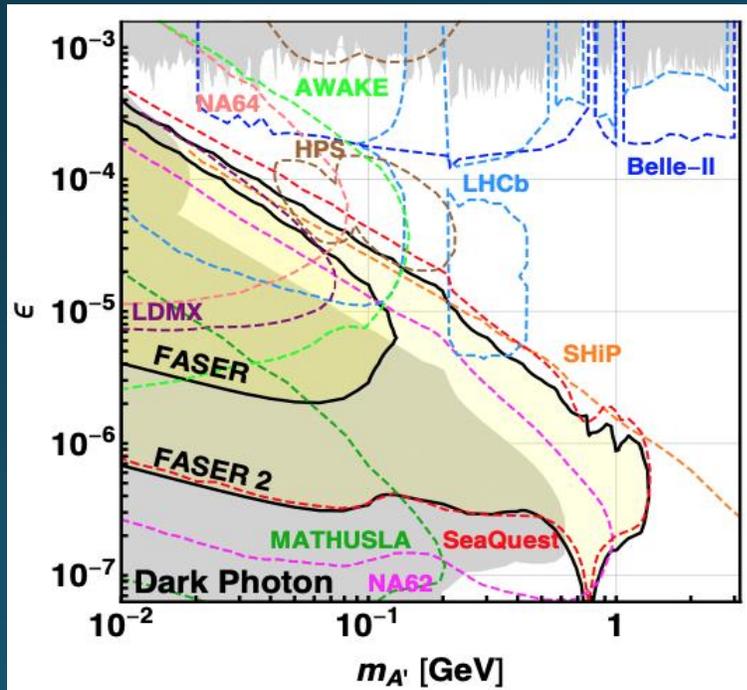
FASER ν 2: Neutrino physics

- FASER ν @ LHC-Run 3 (1.2 ton)
 - Unexplored TeV energy $\sim 1000 \nu_e, \sim 10,000 \nu_\mu, \sim 10 \nu_\tau$ CC events
 - Also SND@LHC (off-axis)
- FASER ν 2 @HL-LHC (~ 10 ton)
 - FASER ν 2: Beam x 20, ~ 10 tons mass \rightarrow 200 times FASER ν
 - $\sim 10^5 \nu_e, 10^6 \nu_\mu, 10^3 \nu_\tau$ CC events
- Tau neutrino physics, precise measurement of cross sections, rare process

FASER2: New particle searches (Long Lived Particles)

- FASER2, New larger detector at Forward Physics Facility
 - FASER (R=10cm, L=1.5m, Run 3) → FASER 2 (R=1m, L=5m, HL-LHC)
 - Largely explore unexplored parameter space
- x 300 decay volume
x 20 beam

Dark photon



Dark higgs

Current status of FPF

- FPF fits European and US's strategy
 - Update of the European Strategy(2020), diversity of particle physics, maximum use of the LHC
 - CERN's Physics Beyond Colliders, <https://indico.cern.ch/category/7885/>
 - US's Snowmass community study and P5 prioritization.
 - FPF <https://doi.org/10.5281/zenodo.4059893> (over 200 signatures within 1 week)
 - LOIS FASER2: https://www.snowmass21.org/docs/files/summaries/EF/SNOWMASS21-EF9_EF6-NF3_NF6-RF6_RF0-CF7_CFO-AF5_AFO_FASER2-038.pdf
 - FASERnu2: https://www.snowmass21.org/docs/files/summaries/NF/SNOWMASS21-NF10_NF6-EF6_EF9-IF0_FASERnu2-006.pdf
 - Neutrino detector: https://www.snowmass21.org/docs/files/summaries/NF/SNOWMASS21-NF10_NFO-EFO_EFO_Ariga-072.pdf
- FPF kick-off workshop (9-10 November 2020)
 - 40 talks, lively discussions over wide topics
 - <https://indico.cern.ch/event/955956>
- Second workshop in 2 weeks <https://indico.cern.ch/event/1022352/>
- HL-LHC is going to start 2027. Now is the time to discuss physics and feasibility of FPF.

Summary

- The **FASER** experiment is a new experiment at the LHC with 2 pillars
 - FASER: Search for new particles
 - FASER ν : Neutrinos
- FASER ν is the first neutrino experiment with a collider
 - Beam at new kinematical regime, including 3 flavors
 - Detector with flavor sensitivity
 - Data taking in 2022-2024
 - **Detection of neutrinos from the LHC was demonstrated with the pilot detector in 2018**
- Future projects (FPF) at the HL-LHC are under discussion

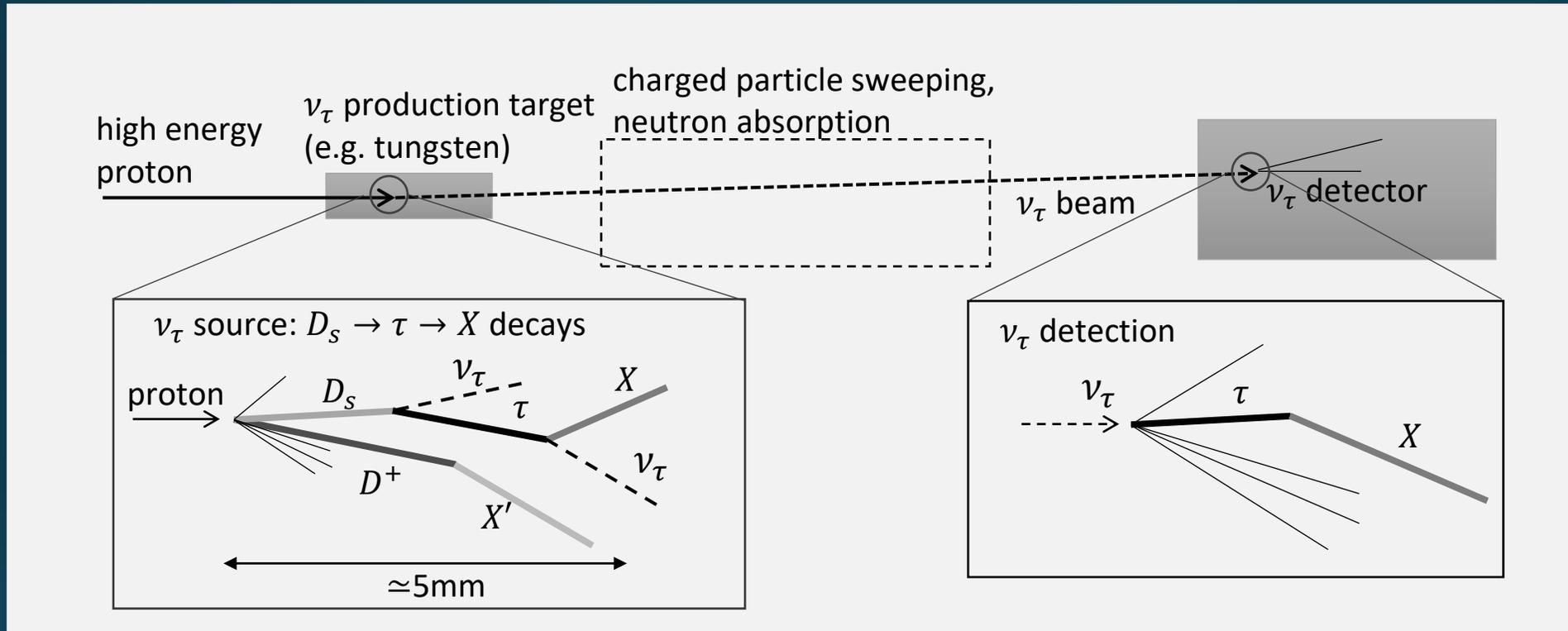
Publications on FASER/FASERnu

- Publications of the FASER Collaboration
 - FASER Letter of Intent at [CERN document server](#) and in [arXiv](#)
 - FASER Technical Proposal at [CERN document server](#) and in [arXiv](#)
 - FASER's Physics Reach for Long-Lived Particles in [Physical Review D](#) and in [arXiv](#)
 - Input to the European Strategy for Particle Physics Update in [arXiv](#)
 - Detecting and Studying High-Energy Collider Neutrinos with FASER at the LHC in [European Physical Journal C](#) and in [arXiv](#)
 - Technical Proposal of FASERν neutrino detector at [CERN document server](#) and in [arXiv](#)
 - **First neutrino interaction candidates at the LHC** in [arXiv](#) *New since last week!*
- Conference talks on FASERnu
 - [Neutrinos at CERN](#), NEUTRINO 2020, 24 June 2020, Tomoko Ariga
 - [FASERnu](#), ICHEP 2020, 28 July - 6 August, Akitaka Ariga

Lepton non-universality?

	channel	Lepton universality
W decay	$W \rightarrow \tau \nu_\tau$	$\Delta(2.8 \sigma)$
B decays	full leptonic $B \rightarrow \tau \nu_\tau$	Δ
	R_D : semi leptonic $B \rightarrow D^{(*)} \tau \nu_\tau$	$\times (3\sigma)$
	R_K : neutral semileptonic $B \rightarrow K \ell^+ \ell^-$	$\times (3\sigma)$
B_s decay	$B_s \rightarrow D_s \tau \nu_\tau$	Δ
	$B_c \rightarrow J/\psi \tau \nu_\tau / B_c \rightarrow J/\psi \mu \nu_\mu$	$\Delta (2\sigma)$
Charm decay	full leptonic $D_s \rightarrow \tau \nu_\tau / D_s \rightarrow \mu \nu_\mu$	$O (1\sigma \text{ excess})$
Lepton leptonic decay	$\tau \rightarrow \mu \nu \nu / \mu \rightarrow e \nu \nu$	\odot
Kaon decay	$K \rightarrow e \nu / K \rightarrow \mu \nu$	\odot
Pion decay	$\pi \rightarrow \mu \nu / \pi \rightarrow e \nu$	\odot
tau CC interaction	never measured	-
ν_τ CC interaction	$\nu_\tau N \rightarrow \tau X$	Δ (too few statistics)

Accelerator-based ν_τ cross section measurement



ν_τ production study: DsTau (NA65)

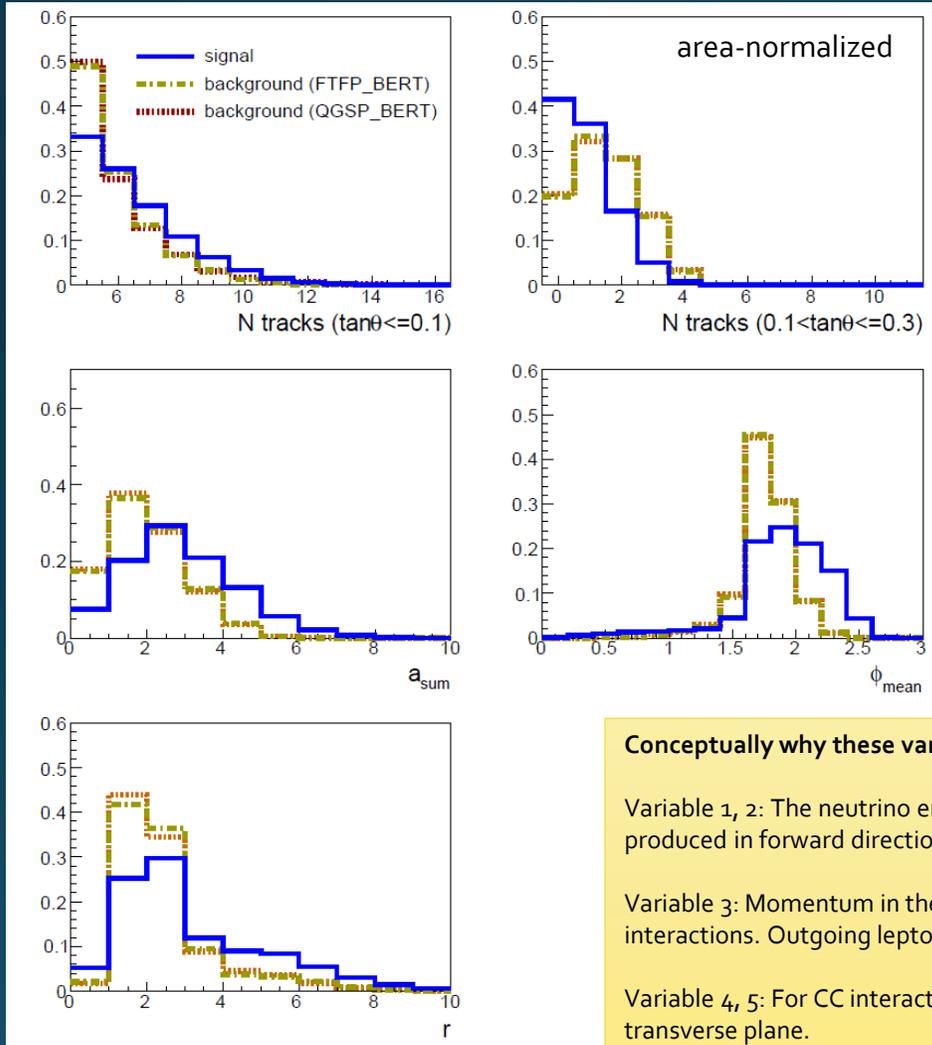
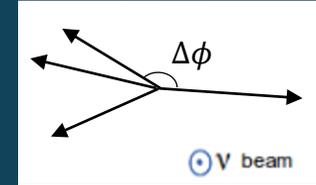
- No experimental data on the D_s differential cross section
- Large systematic uncertainty ($\sim 50\%$) in the ν_τ flux prediction

ν_τ detection: e.g. DONuT, SHiP, FASER ν

- Statistical uncertainty 33% in DONUT
- Will be reduced to the 2% level in future experiments

Variables for MVA

Expected distributions of the variables



5 variables used in the analysis

1. the number of tracks with $\tan\theta \leq 0.1$ with respect to the beam direction
2. the number of tracks with $0.1 < \tan\theta \leq 0.3$ with respect to the beam direction
3. the absolute value of vector sum of transverse angles calculated considering all the tracks as unit vectors in the plane transverse to the beam direction (a_{sum})
4. for each track in the event, calculate the mean value of opening angles between the track and the others in the plane transverse to the beam direction, and then take the maximum value in the event (ϕ_{mean})
5. for each track in the event, calculate the ratio of the number of tracks with opening angle ≤ 90 degrees and > 90 degrees in the plane transverse to the beam direction, and then take the maximum value in the event (r).

Multiplicity and Pseud rapidity distribution

Momentum balance

Back-to-back kinematics at vertex

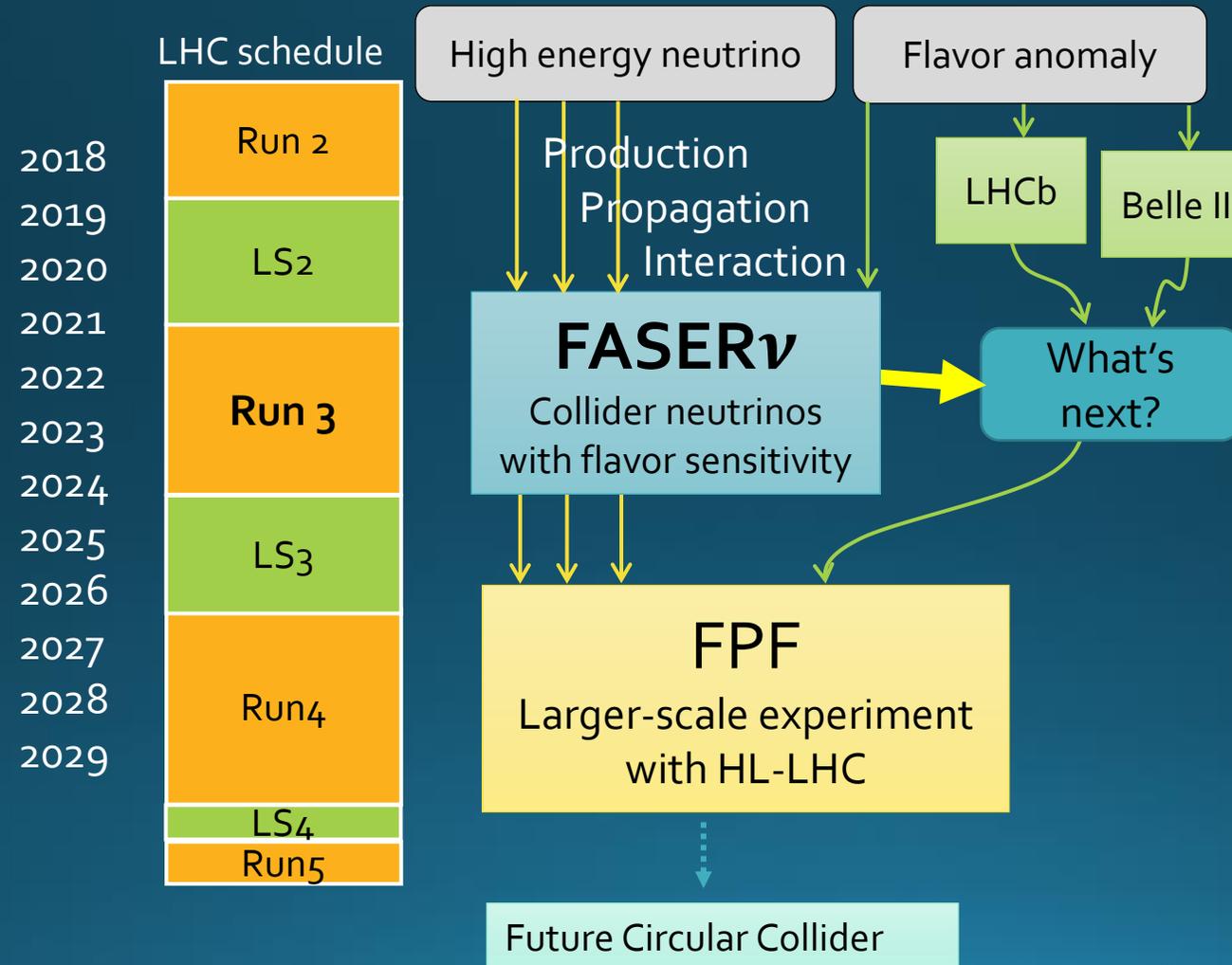
Conceptually why these variables are good:

Variable 1, 2: The neutrino energy is higher than the neutral hadron energy. Higher energy, more particles are produced in forward direction, i.e. $\tan(\theta) < 0.1$ (var 1), and higher ratio of var1/var2.

Variable 3: Momentum in the transverse plane is more balanced in hadron interactions than neutrino CC and NC interactions. Outgoing leptons in neutrino interactions take a major energy, which distorts this variable.

Variable 4, 5: For CC interactions, we expect the outgoing lepton and hadron system are back to back in the transverse plane.

Roadmap towards a future experiment



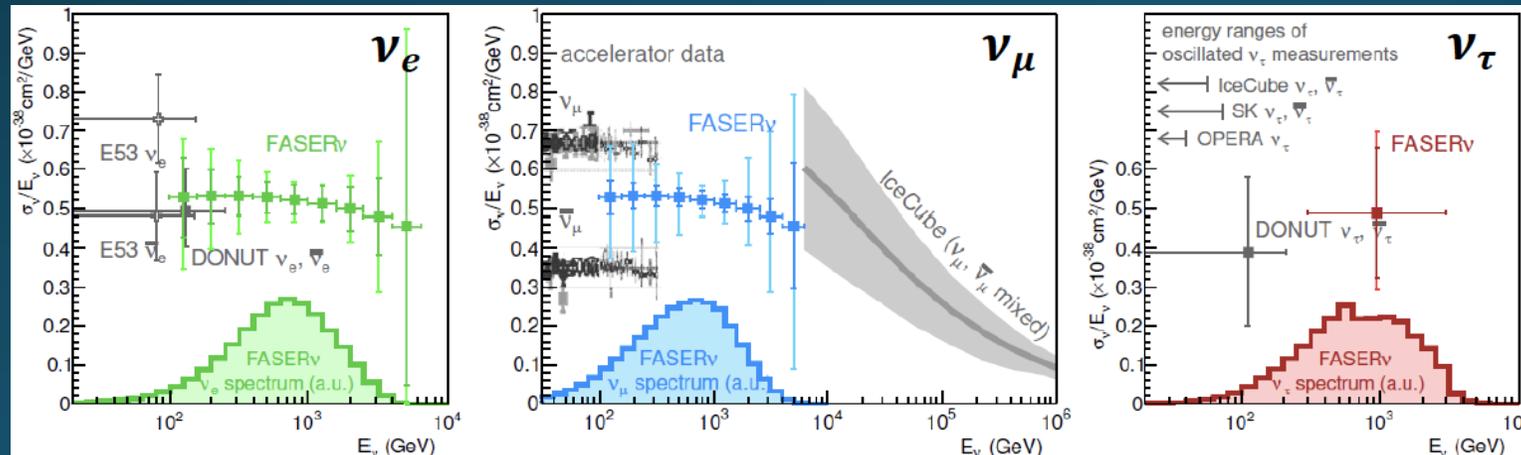
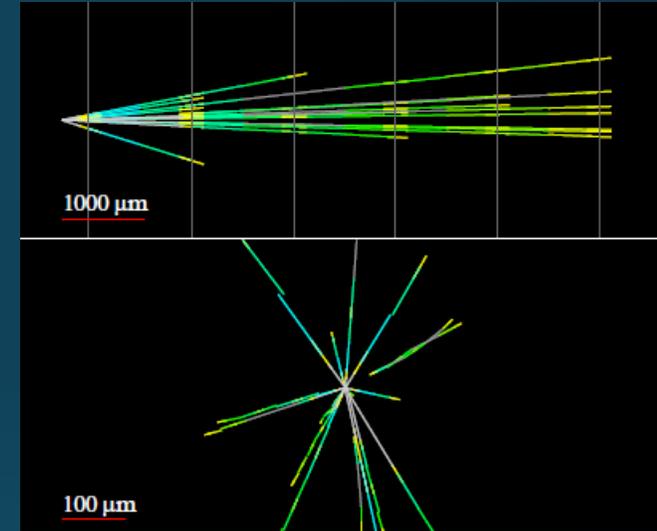
FASER₂ Physics Sensitivity

- Physics Beyond Colliders benchmark cases

Benchmark Model	FASER	FASER ₂	References
V ₁ /BC ₁ : Dark Photon	√	√	Feng, Galon, Kling, Trojanowski, 1708.09389
V ₂ /BC _{1'} : U(1) _{B-L} Gauge Boson	√	√	Bauer, Foldenauer, Jaeckel, 1803.05466 FASER Collaboration, 1811.12522
BC ₂ : Invisible Dark Photon	–	–	–
BC ₃ : Milli-Charged Particle	–	–	–
S ₁ /BC ₄ : Dark Higgs Boson	–	√	Feng, Galon, Kling, Trojanowski, 1710.09387 Batell, Freitas, Ismail, McKeen, 1712.10022
S ₂ /BC ₅ : Dark Higgs with hSS	–	√	Feng, Galon, Kling, Trojanowski, 1710.09387
F ₁ /BC ₆ : HNL with e	–	√	Kling, Trojanowski, 1801.08947 Helo, Hirsch, Wang, 1803.02212
F ₂ /BC ₇ : HNL with μ	–	√	Kling, Trojanowski, 1801.08947 Helo, Hirsch, Wang, 1803.02212
F ₃ /BC ₈ : HNL with τ	√	√	Kling, Trojanowski, 1801.08947 Helo, Hirsch, Wang, 1803.02212
A ₁ /BC ₉ : ALP with photon	√	√	Feng, Galon, Kling, Trojanowski, 1806.02348
A ₂ /BC ₁₀ : ALP with fermion	√	√	FASER Collaboration, 1811.12522
A ₃ /BC ₁₁ : ALP with gluon	√	√	FASER Collaboration, 1811.12522

FASERν2: Neutrino physics

- FASERν @ LHC-Run 3 (1.2 ton)
 - Unexplored TeV energy $\sim 1000 \nu_{e\tau}$, $\sim 10,000 \nu_{\mu\tau}$, $\sim 10 \nu_{\tau}$ CC events
 - Also SND@LHC (off-axis)
- FASERν2 @HL-LHC (~ 10 ton)
 - FASERν2: Beam $\times 20$, ~ 10 tons mass \rightarrow 200 times FASERν $\sim 10^5 \nu_{e\tau}$, $10^6 \nu_{\mu\tau}$, $10^3 \nu_{\tau}$ CC events
- Tau neutrino physics, precise measurement of cross sections, rare process



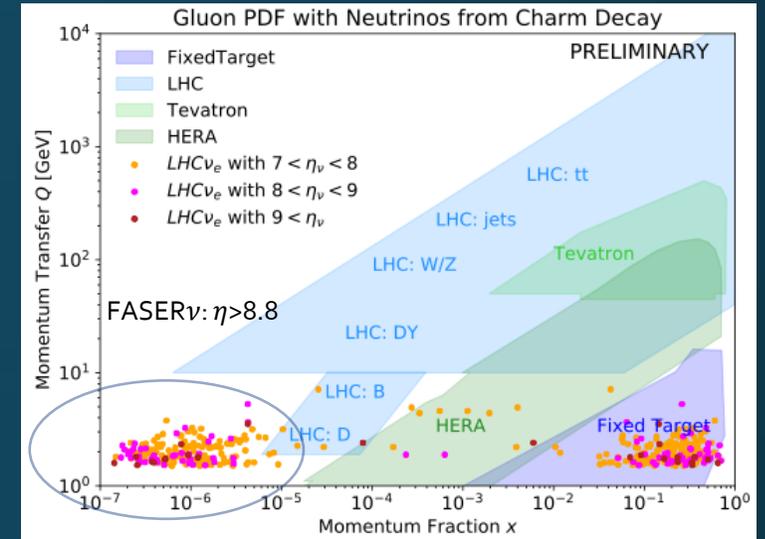
FASER ν 2:QCD physics

- 超前方ハドロン生成

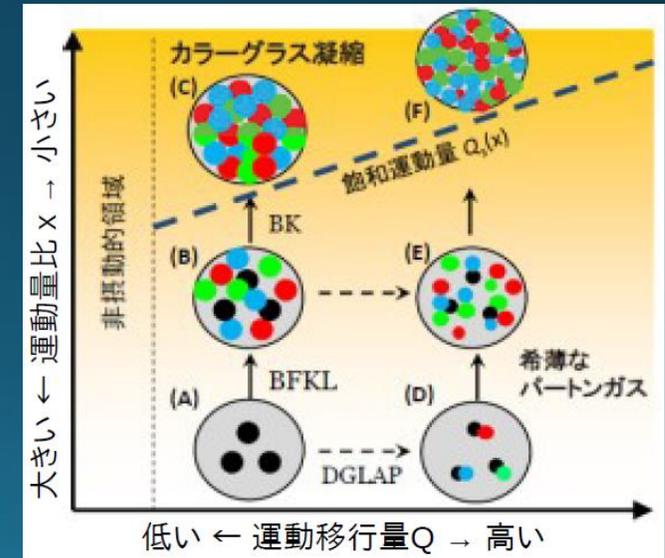
- 現状、データの欠如・大きな不定性。
 - モデルにより大きく違う(EPOS-LHC, OGSJET, DPMJET, SIBYLL, PYTHIA)
 - 陽子陽子衝突時、Small-x とLarge-xのパートンが寄与。カラーグラス凝縮、Intrinsic charm \rightarrow ニュートリノスペクトルにゆがみ
- 宇宙線物理学へのインプット。E.g.) IceCubeでの高エネルギー宇宙ニュートリノ解析へのプロンプトニュートリノ背景事象に制限

- FASER ν 2にてQCDの詳細解析

- ニュートリノ生成
 - 陽子内パートン ($K, D \rightarrow \nu_e, \pi \rightarrow \nu_\mu, D \rightarrow \nu_\tau$)
- ニュートリノ反応
 - ν CC チャーム粒子生成による標的中のストレンジネスパートンの研究 $\nu_\ell S \rightarrow \ell C$



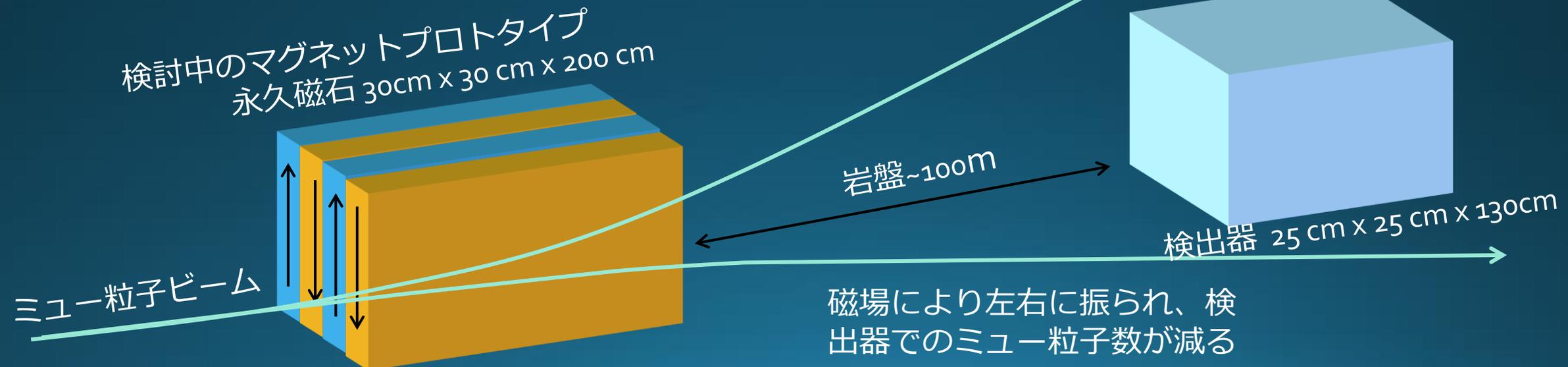
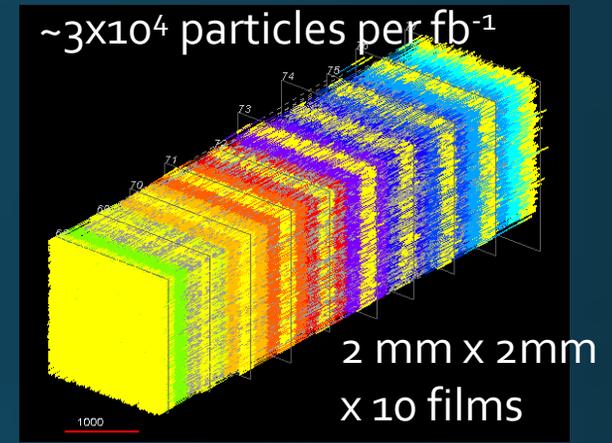
ν flux at FASER ν に寄与するパートン



ミュー粒子バックグラウンド

2018年
FASER data

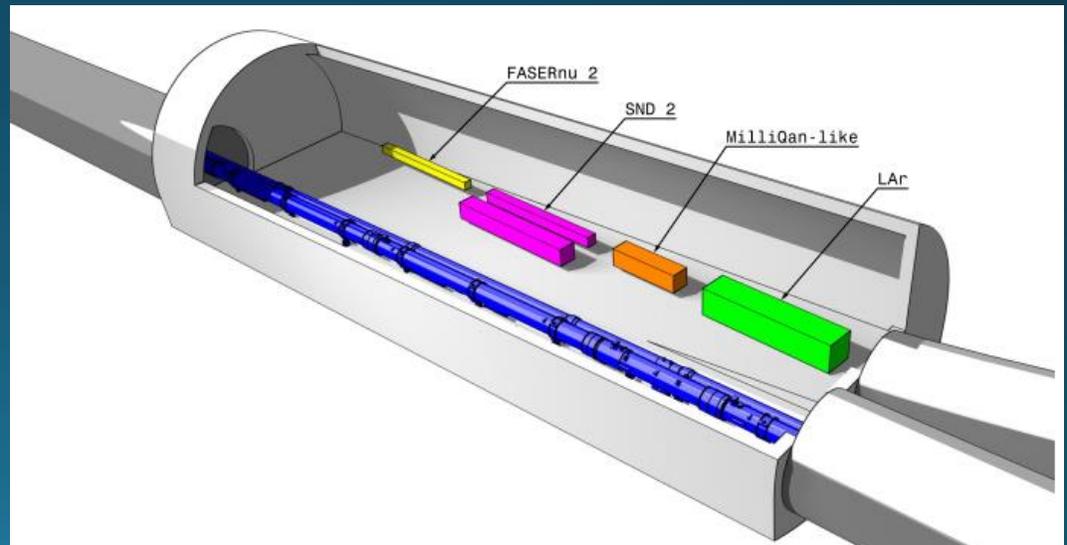
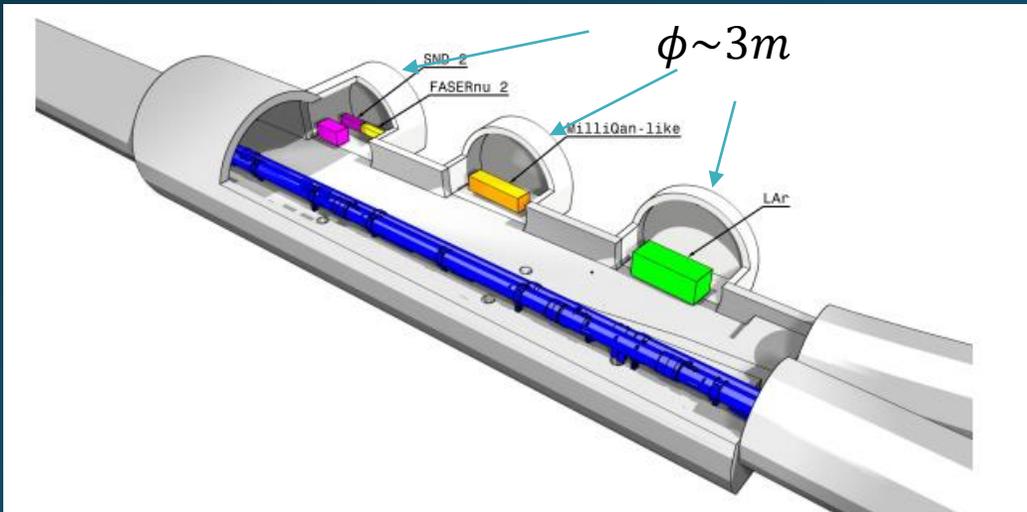
- 高いミュー粒子バックグラウンドが物理感度を制限
 - ニュートリノ・ミリチャージ実験には死活問題
- ミュー粒子除去法の開発が必要
 - 検出器上流にミュー粒子除去マグネットを設置



複数のFPFオプション ～既存のトンネルの拡張

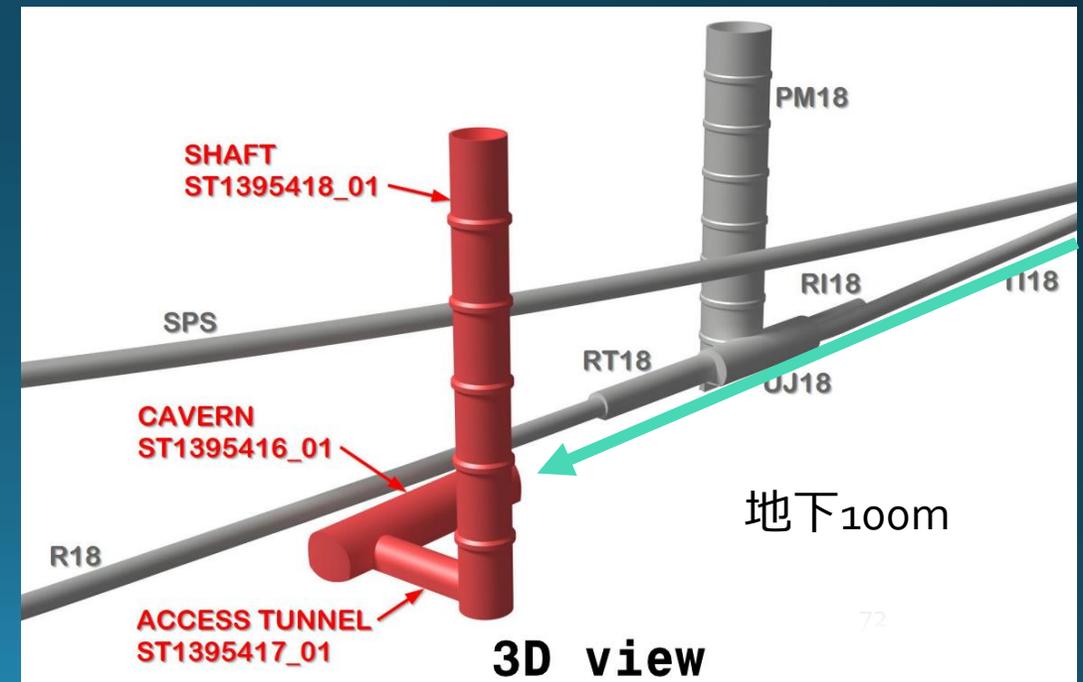
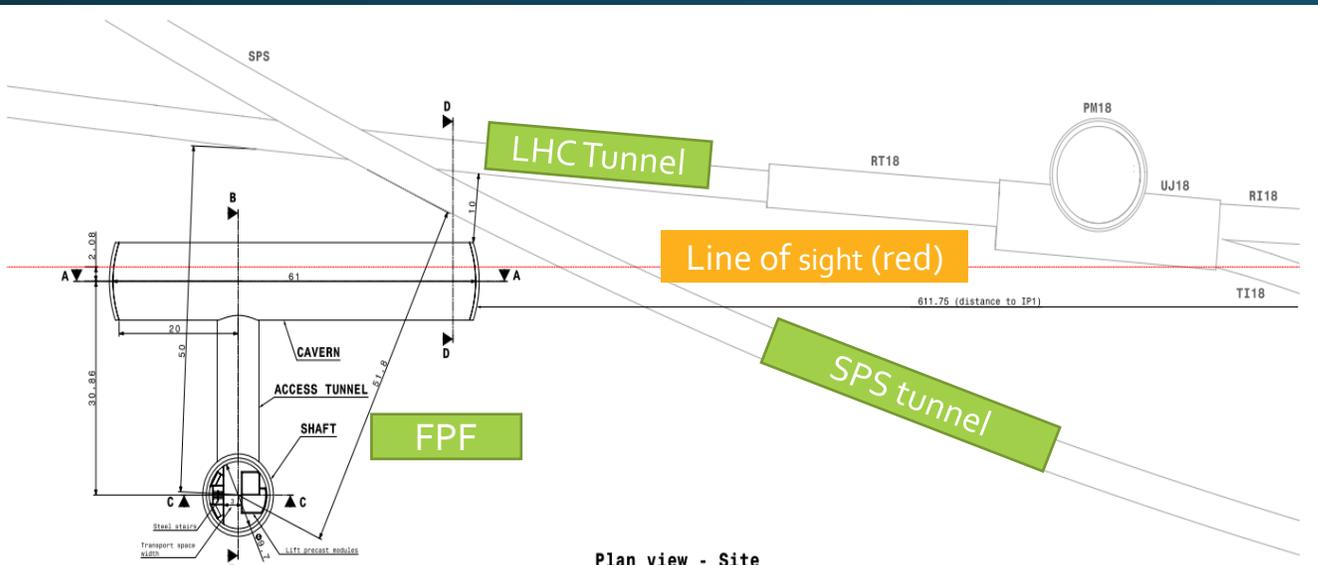
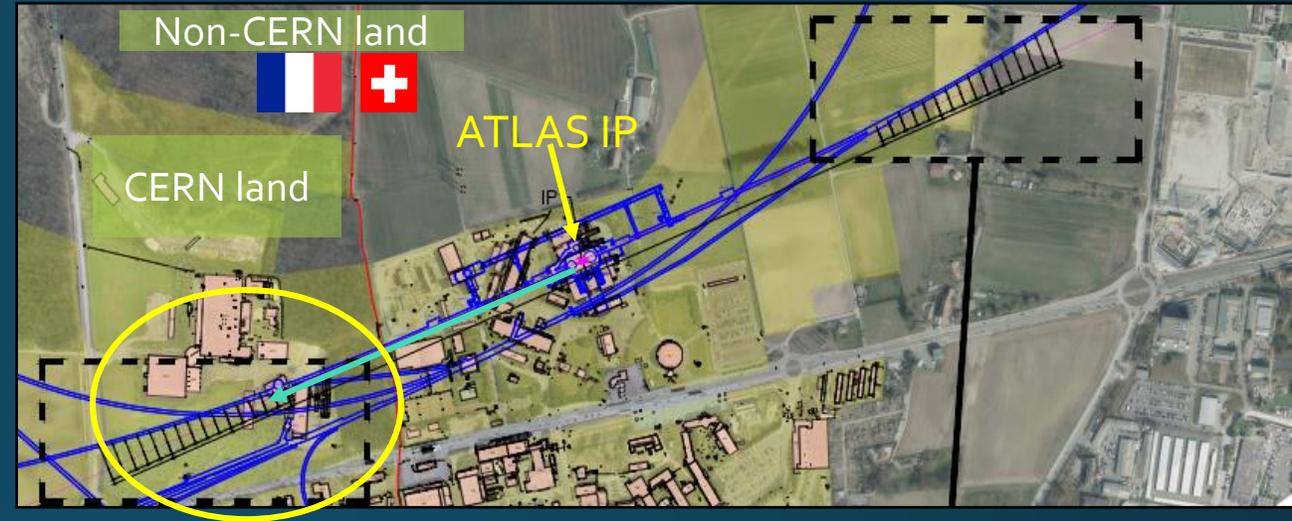
- Option 1
- 既存のLHCトンネルに横穴
- 安価だが使えるスペースが限定的

- Option 2
- 既存のトンネルを拡張
- 大規模な工事が必要
- 高価



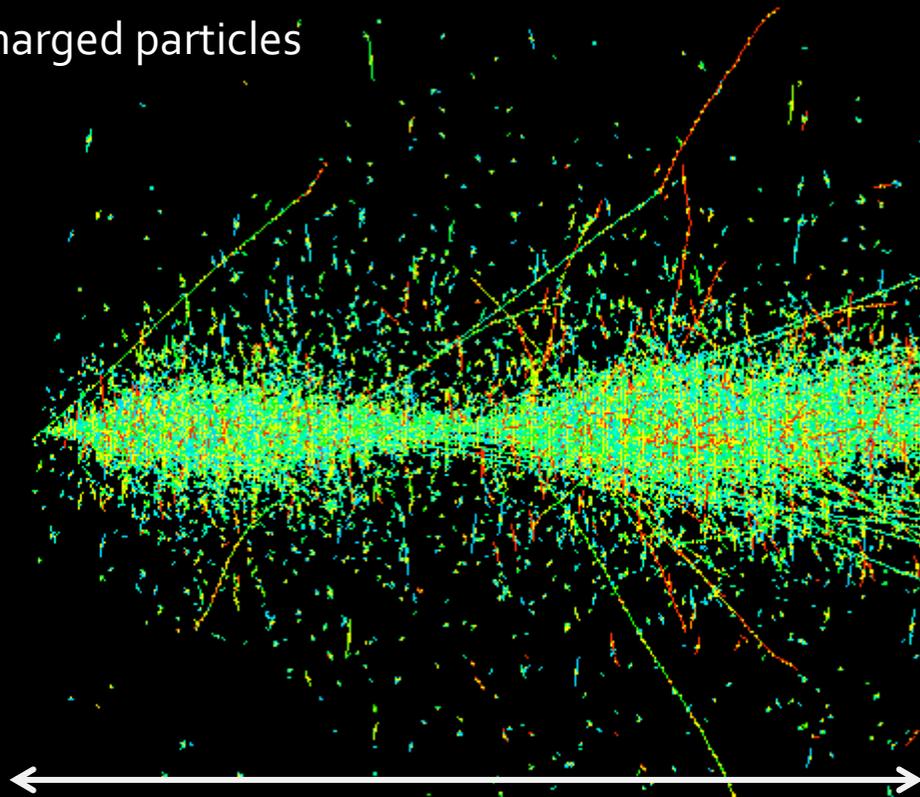
複数のFPFオプション ～新しい縦穴

- ATLASから600-800m地点に縦穴を掘り、実験ホールを作る。
- 高価だがスペースの大きさ、建設スケジュールの立てやすさ等にメリット。LHCのスケジュールに非干渉。



Simulated 1 TeV ν_μ CC interaction

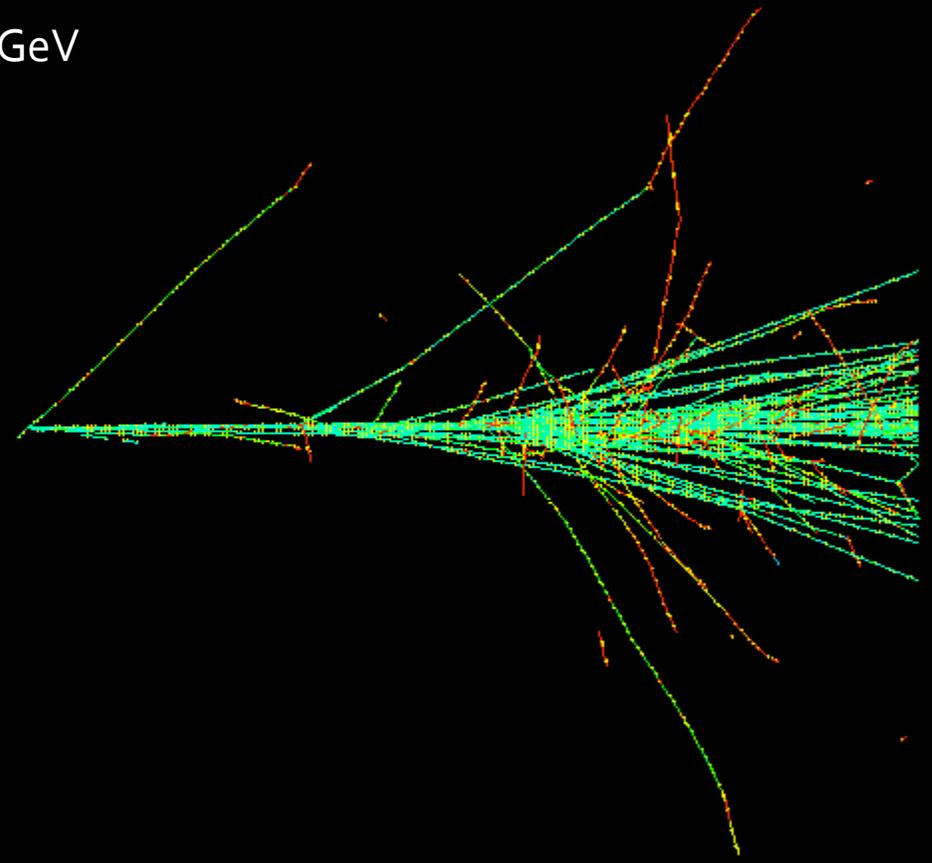
All charged particles



200 tungsten plates (27 cm)
 $\sim 57 X_0, \sim 2 \lambda_{int}$

50000 μm

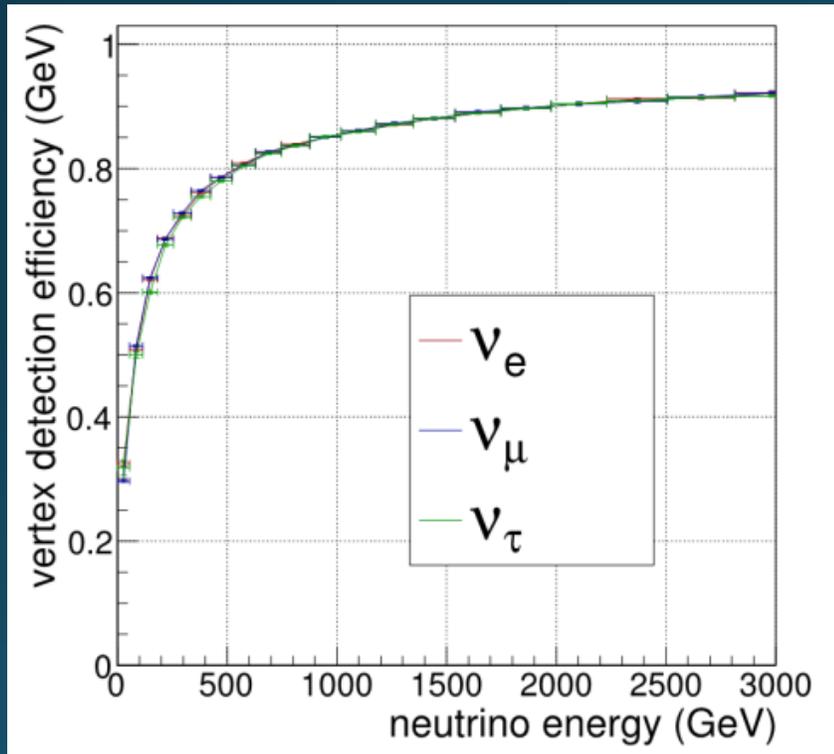
$P > 0.3$ GeV



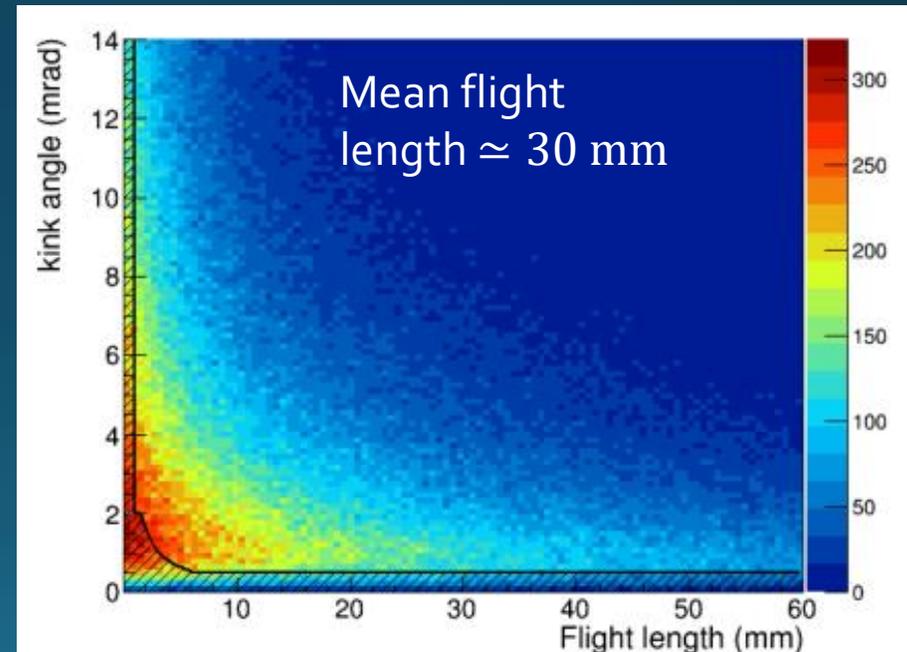
50000 μm

Detection efficiency

Vertex detection efficiency
(charged multiplicity ≥ 5)



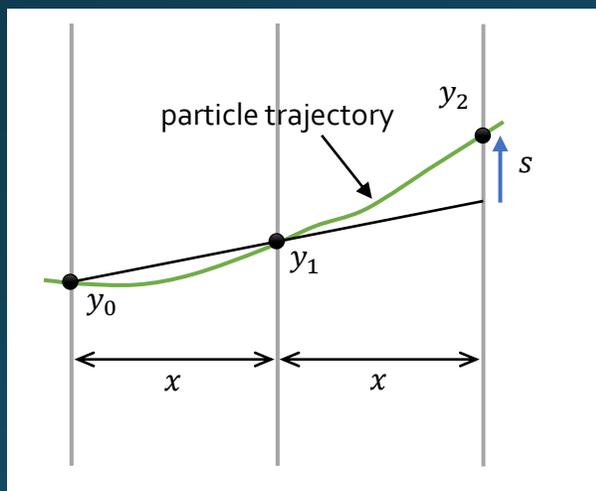
Tau decay detection efficiency
=75% ($\tau \rightarrow 1$ prong)



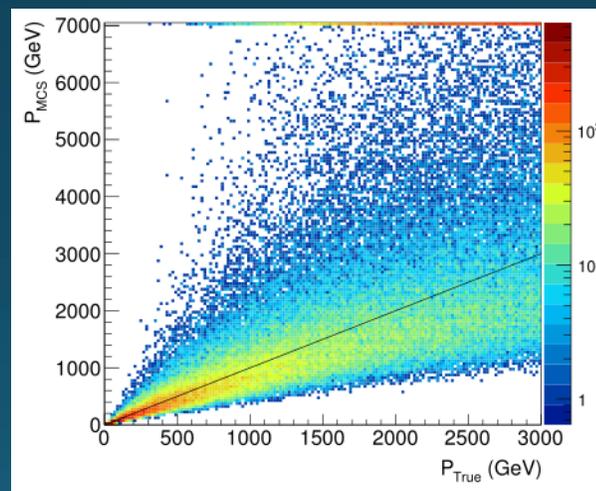
Particle momentum measurement

by multiple Coulomb scattering (MCS)

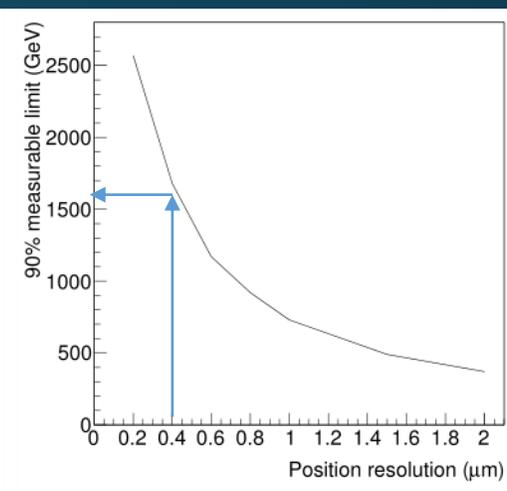
- Sub-micron precision alignment using muon tracks
 - Our experience = 0.4 μm (in the DsTau experiment)
- This allow to measure particle momenta by MCS, even above 1 TeV.



$$(s^{\text{RMS}})^2 = \left(\sqrt{\frac{2}{3}} \frac{13.6(\text{MeV})}{\beta P} x \sqrt{\frac{x}{X_0}} \right)^2 + (\sqrt{6} \sigma_{\text{pos}})^2$$



Performance with position resolution of 0.4 μm , in 100 tungsten plates (MC)



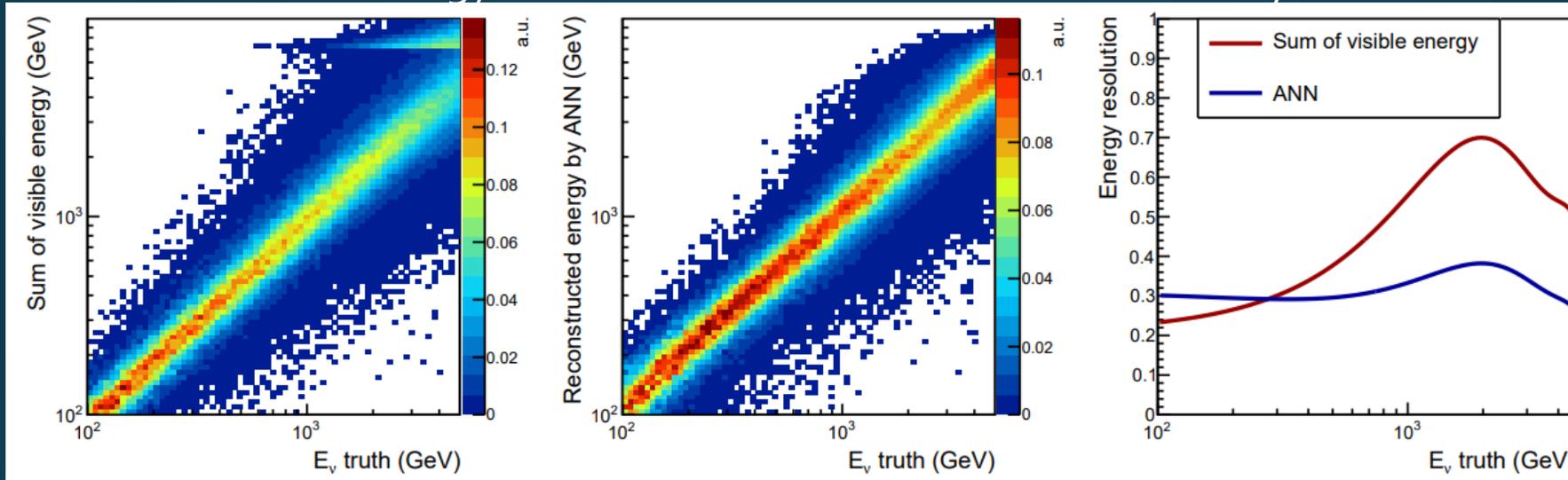
Measurable energy vs position resolution

Energy reconstruction (ν_μ CC)

Sum of visible energy

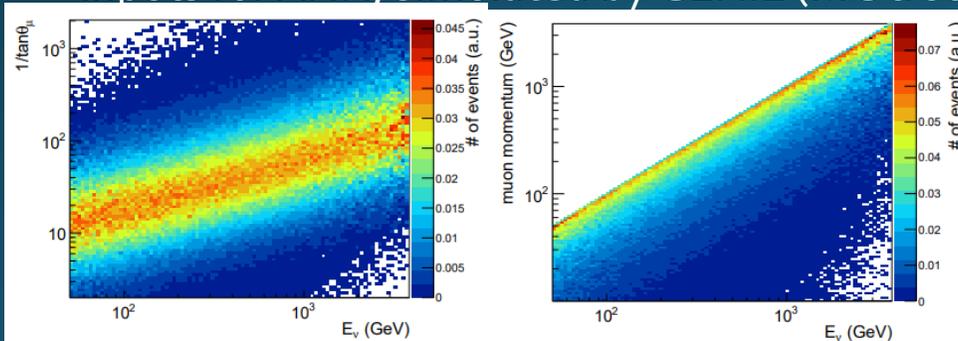
ANN method

$\Delta E/E$



(smeared)

inputs for ANN, simulated by GENIE (MC truth)



Angular info

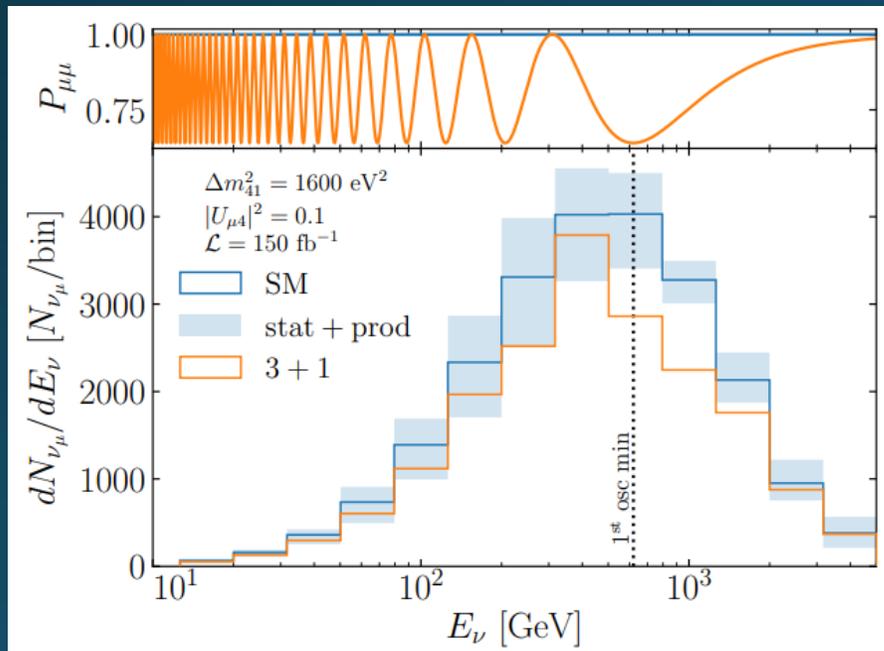
Momentum

...

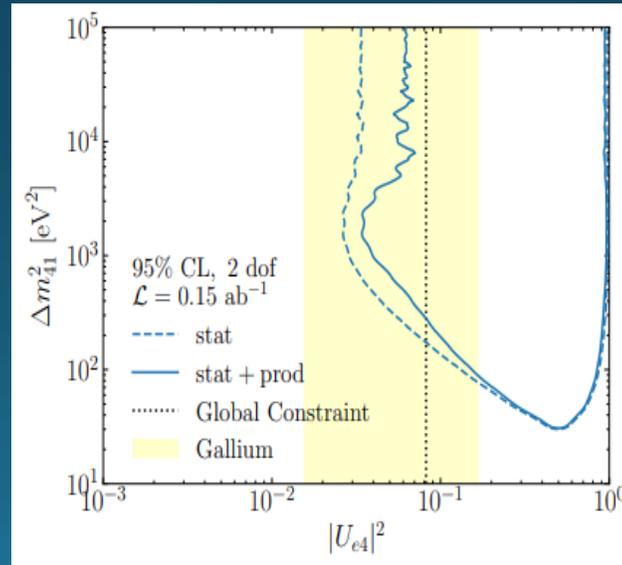
- Sum of visible energy (model independent) already gives a reasonable resolution
- ANN can solve problem at high energy and gives about 30% resolution at relevant energy range.

Sterile neutrino oscillation

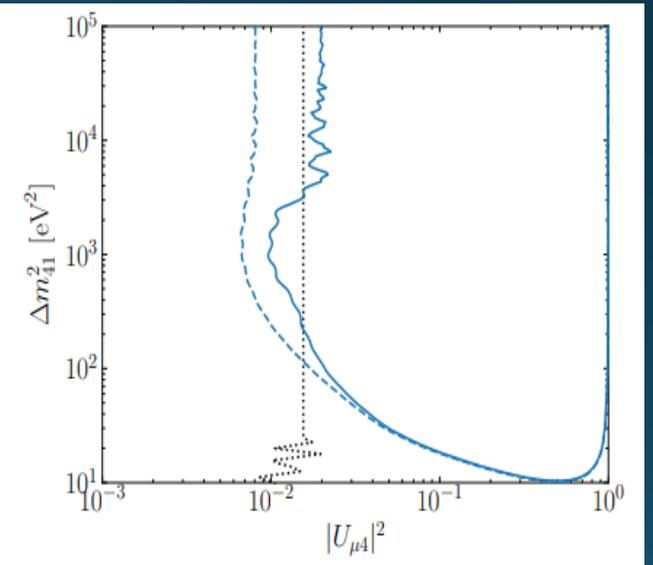
- Due to unique energy and baseline ($L/E \sim 10^{-3}$ m/MeV), FASER ν is sensitive to large $\Delta m^2 \sim 10^3$ eV 2 .
- Neutrino spectrum deformation
- Competitive in disappearance channels.



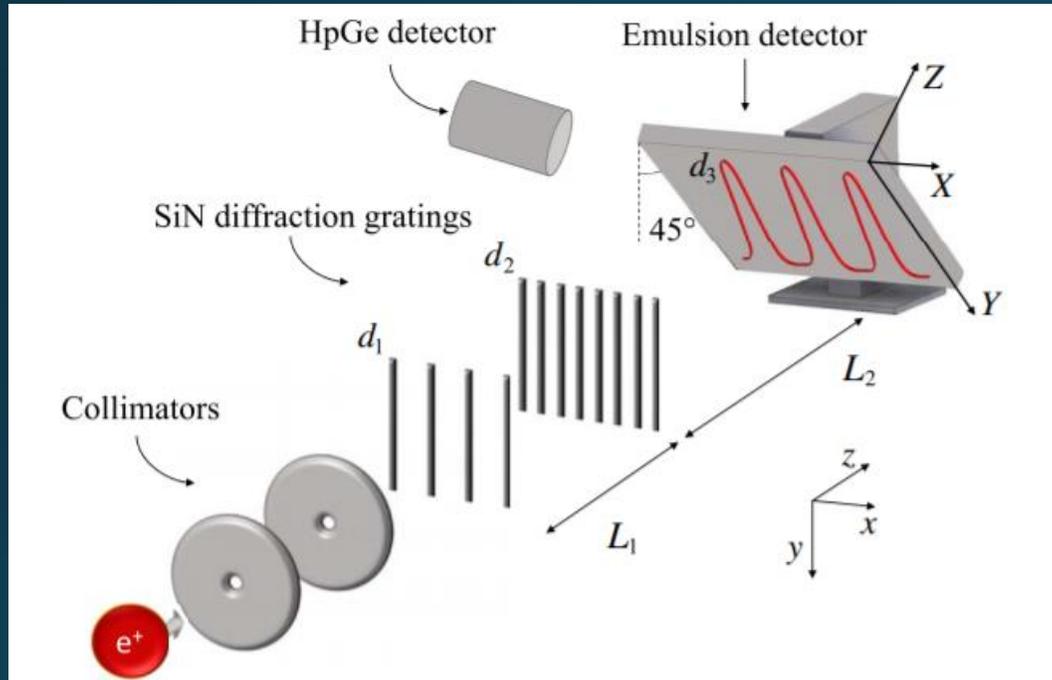
ν_e disappearance



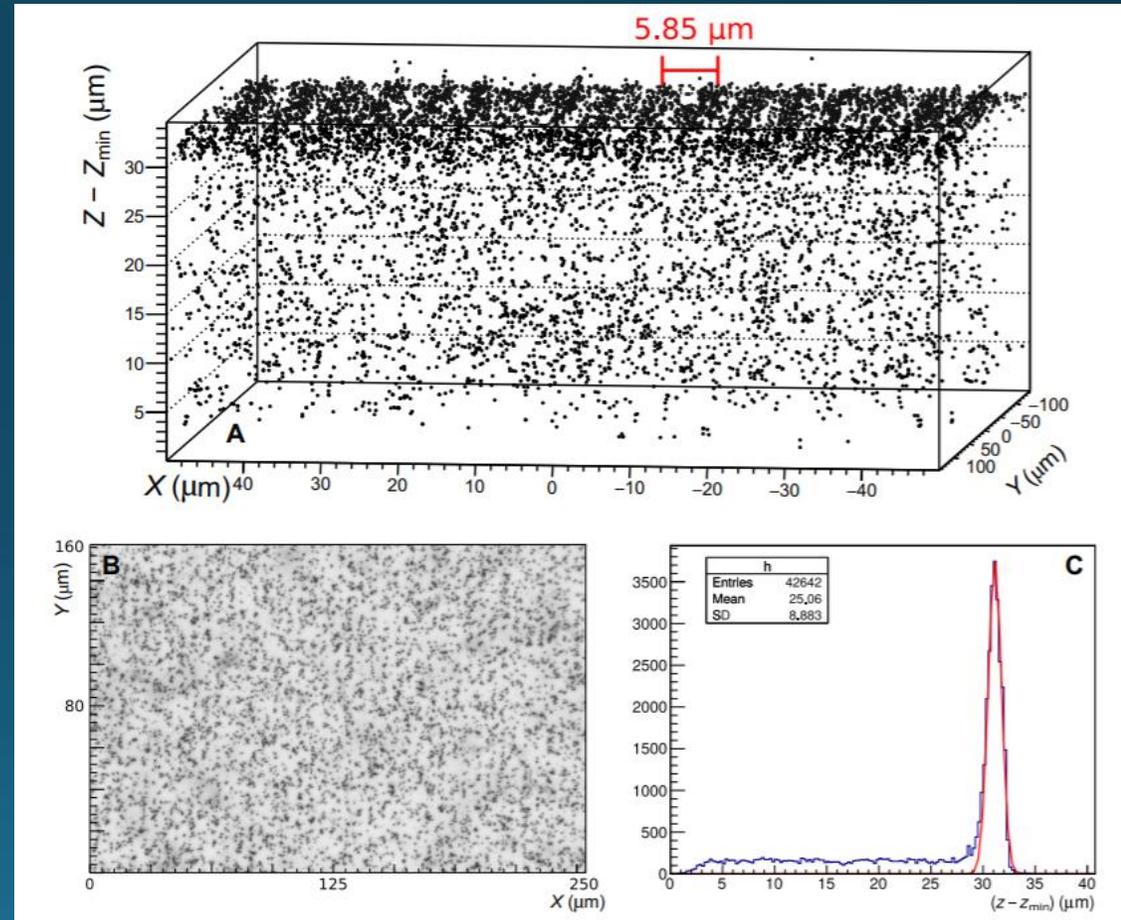
ν_μ disappearance



QUPLAS: First demonstration of antimatter wave interferometry



- 8-14 keV positron



Glacier bedrock radiography

- Muon radiography applied to Swiss alps
- Discovery of steep bedrock shape, need a new understanding of glacial erosion process.

• [Nature Scientific Reports](#)

• [s41598-019-43527-6](#)

