

#### Neutrino interactions with nuclei and star matter

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#### Current Status of low-energy nuclear physics



### Benchmark of ab-initio methods for oxygen isotopic chain

First success of chital-EFT interactions on oxygen isotopes....



#### Spectroscopy via knock out reactions-basic idea

Use a probe (ANY probe) to eject the particle we are interested to:



### Concept of correlations



Understood for a few stable closed shells:

URRF

UNIVERSITY OF [CB and W. H. Dickhoff, Prog. Part. Nucl. Phys 52, 377 (2004)]

#### One-nucleon spectral function from SCGF



W. Dickhoff, CB, Prog. Part. Nucl. Phys. 53, 377 (2004); CB, M.Hjorth-Jensen, Pys. Rev. C79, 064313 (2009)

Electron and neutrino

scattering off nuclei

N. Rocco, CB, Phys. Rev. C98, 025501 (2018)

N. Rocco, CB, O. Benhar, A. De Pace, A. Lovato, Phys. Rev. C99, 025502 (2019)



### Lepton-nucleon cross section

$$\left(\frac{d\sigma}{dT'd\cos\theta'}\right)_{\nu/\bar{\nu}} = \frac{G^2}{2\pi} \frac{k'}{2E_{\nu}} \left[ \hat{L}_{CC}R_{CC} + 2\hat{L}_{CL}R_{CL} + \hat{L}_{LL}R_{LL} + \hat{L}_TR_T \pm 2\hat{L}_{T'}R_{T'} \right] \,,$$

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## Nuclear structure is in the hadronic tensor:

$$\begin{aligned} R_{CC} &= W^{00} \\ R_{CL} &= -\frac{1}{2} (W^{03} + W^{30}) \\ R_{LL} &= W^{33} \\ R_T &= W^{11} + W^{22} \\ R_{T'} &= -\frac{i}{2} (W^{12} - W^{21}) \,, \end{aligned}$$

$$W^{\mu\nu} = \sum_{f} \langle 0|j^{\mu\dagger}|f\rangle \langle f|j^{\nu}|0\rangle \delta(E_0 + \omega - E_f)$$



### Lepton-nucleon cross section

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Nuclear structure is in the hadronic tensor:

Two models of the Spectral function

$$W^{\mu\nu}(\mathbf{q},\omega) = \int \frac{d^3k}{(2\pi)^3} dE P_h(\mathbf{k},E) \frac{m^2}{e(\mathbf{k})e(\mathbf{k}+\mathbf{q})}$$
$$\times \sum_i \langle k|j_i^{\mu\dagger}|k+q\rangle \langle k+q|j_i^{\nu}|k\rangle$$
$$\times \delta(\omega+E-e(\mathbf{k}+\mathbf{q})),$$

$$\begin{split} P_{h}(\mathbf{k}, E) &= \frac{1}{\pi} \sum_{\alpha \beta} \tilde{\Phi}_{\beta}^{*}(\mathbf{k}) \tilde{\Phi}_{\alpha}(\mathbf{k}) \\ &\times \operatorname{Im} \langle \psi_{0}^{A} | a_{\beta}^{\dagger} \frac{1}{E + (H - E_{0}^{A}) - i\epsilon} a_{\alpha} | \psi_{0}^{A} \rangle \end{split}$$

SCGF/ADC(3) using chiral NNLOsat



- SURREY - N. Rocco, CB, Phys. Rev. C98, 025501 (2018).

## <sup>16</sup>O-e<sup>-</sup> cross sections from the SCGF Spect. Fnct.



### Lepton-nucleon cross section

$$\left(\frac{d\sigma}{dT'd\cos\theta'}\right)_{\nu/\bar{\nu}} = \frac{G^2}{2\pi} \frac{k'}{2E_{\nu}} \left[ \hat{L}_{CC}R_{CC} + 2\hat{L}_{CL}R_{CL} + \hat{L}_{LL}R_{LL} + \hat{L}_TR_T \pm 2\hat{L}_{T'}R_{T'} \right],$$

Π.

Nuclear structure is in the hadronic tensor:

$$W^{\mu\nu}(\mathbf{q},\omega) = \int \frac{d^3k}{(2\pi)^3} dE P_h(\mathbf{k},E) \frac{m^2}{e(\mathbf{k})e(\mathbf{k}+\mathbf{q})}$$
$$\times \sum_i \langle k | j_i^{\mu\dagger} | k+q \rangle \langle k+q | j_i^{\nu} | k \rangle$$
$$\times \delta(\omega+E-e(\mathbf{k}+\mathbf{q})),$$

$$W_{2\mathrm{b}}^{\mu\nu}(\mathbf{q},\omega) = \frac{V}{2} \int d\tilde{E} \frac{d^3k}{(2\pi)^3} d\tilde{E}' \frac{d^3k'}{(2\pi)^3} \frac{d^3p}{(2\pi)^3} \\ \times \frac{m^4}{e(\mathbf{k})e(\mathbf{k}')e(\mathbf{p})e(\mathbf{p}')} P_h^{\mathrm{NM}}(\mathbf{k},\tilde{E}) P_h^{\mathrm{NM}}(\mathbf{k}',\tilde{E}') \\ \times \sum_{ij} \langle k \, k' | j_{ij}^{\mu \dagger} | p \, p' \rangle \langle p \, p' | j_{ij}^{\nu} | k \, k' \rangle \\ \times \, \delta(\omega + \tilde{E} + \tilde{E}' - e(\mathbf{p}) - e(\mathbf{p}')) \,.$$
(41)

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• Two-body diagrams contributing to the axial and vector responses



N. Rocco, CB, O. Benhar, A. De Pace, A. Lovato, Phys. Rev. C99, 025502 (2019)

## **Charged-current** reaction for 1 GeV neutrinos



One-body current describe quasi elastic peak

Difference between CBF(AV18) and SCGF(NNLOsat) from 1-b terms

Two-body currents fiull up dip region

Missing Delta and meson emission contributions

X-sec. droppin with scattering angle

N. Rocco, CB, O. Benhar, A. De Pace, A. Lovato, Phys. Rev. C99, 025502 (2019)

# Role of two-body (meson exchange) currents in v-A





• Two-body diagrams contributing to the axial and vector responses



• Preliminary implementation discards 1b-2b interference:

$$W^{\mu\nu}_{2p2h} = W^{\mu\nu}_{ISC} + W^{\mu\nu}_{MEC} + W^{\mu\nu}_{MeC}$$

N. Rocco, CB, O. Benhar, A. De Pace , A. Lovato, arXiv:1810.0wxyz (in preparation)

# Neutrino Oscillations - next generation experiments

DUNE experiment will measure long base line neutrino oscillations (mass hierarchy, CP violation in weak interactions, physics beyond SM)



Liquid Argon projection chamber is being used. It will require one order of magnitude ( $20\% \rightarrow 2\%$ ) improvement in theoretical prediction for v-<sup>40</sup>Ar cross sections to achieve proper event reconstruction.

- → Need good knowledge of <sup>40</sup>Ar spectral functions.
- Jlab experiment E12-14-012 (Hall A) Phys. Rev. C 98, 014617 (2018); arXiv:1810.10575





### Spectral function for <sup>40</sup>Ar



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- Experimental datat now available from Jlab:
  H. Dai et al., arXiv:1803.01910/ 1810.10575
- Ab initio simulations based on the ADC(2) truncation of the N2LO-sat Hamiltoninan

→ Want validation of initial state correlation <u>before</u> they are implementer in neutrino-<sup>40</sup>Ar simulations



## Electron and v scattering on <sup>40</sup>Ar and Ti

Jlab experiment E12-14-012 (Hall A)

[Phys. Rev. C 98, 014617 (2018)]

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## Electron and v scattering on <sup>40</sup>Ar and Ti

#### Jlab experiment E12-14-012 (Hall A)

[Phys. Rev. C 98, 014617 (2018)]



('mix') is nearly identical to neutrons in  ${}^{40}Ar$ .

<sup>40</sup>Ar(e,e'p) and Ti(e,e'p) data being analyzed UNIVERSITY OF

#### Neutrino transport in dense matter supernovae and proton-neutron stars

A. Rios, S. Reddy, CB, work in progress...



## Neutrino propagation in dense matter

#### Kinematical constraint

$$E_n(|\vec{k}|) - E_p(|\vec{k} + \vec{q}|) = -\omega$$

Non-interacting

$$E_{\tau}(|\vec{k}|) = M_{\tau} + \frac{k^2}{2M_{\tau}}$$

#### Interacting quasi-particle

$$E_{\tau}(|\vec{k}|) = M_{\tau} + \frac{k^2}{2M_{\tau}} + \Sigma_{\tau}(k)$$
$$\Sigma_n(k) - \Sigma_p(k) \simeq 30$$

Fragmented strength?









# Neutrino propagation in dense matter

#### Effective masses:

(dependence of the modelling of the nuclear force...)

Response funtions:

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<u>Very</u> preliminary computations, by A. Rios

#### Summary

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Thank you

#### Ab initio computations for nuclear structure

- → Strong link to underlying QCD thorugh EFT-inspiered Hamiltonians
- → Ground state properties up to masses A~100 now with high accuracy; spectral distributions becoming available...

![](_page_20_Figure_4.jpeg)

![](_page_20_Figure_5.jpeg)

Applications to electron and neutrino scattering:

→ Good reproduction of charge/momentum distribution and electron scattering.

→ Inclusion of electroweak currents (1b and 2b) underway (by N. Rocco).

 $\rightarrow$  Validation of nuclear structure input from <sup>40</sup>Ar and Ti data from Jlab.

Neutrino propagation in dense matter

![](_page_20_Picture_11.jpeg)

![](_page_20_Picture_12.jpeg)

![](_page_20_Picture_13.jpeg)

O. Benhar