



Muon Modulation due to atmospheric temperature variation

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Contents



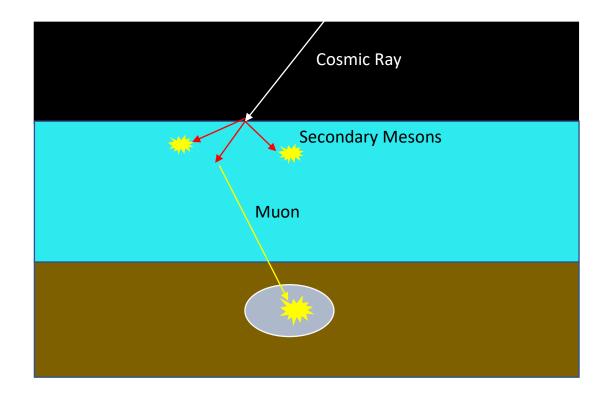
- Cosmic Rays an introduction
- IceCube and DMIce-17 the detectors
- Modulation in DMIce-17
- The α parameter
- IceCube Hard Local Coincidence rate
- Work for the future



Muons from Cosmic Rays



- Cosmic rays impact the atmosphere, producing mesons, which decay to muons or interact hadronically
- These muons can be detected at ground level and below in practically any detector

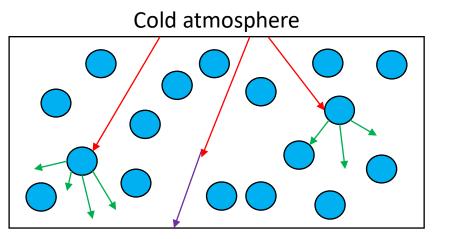


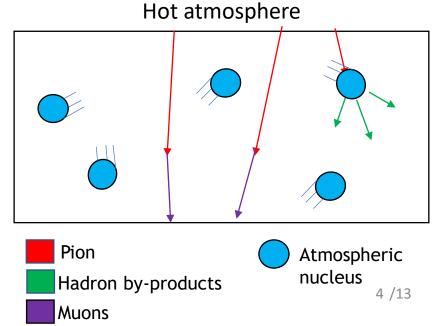


Seasonal Variation of Muons

- In local summer, atmosphere is warmer, less dense. This means cosmic ray daughter mesons are less likely to interact hadronically with atmosphere, and more likely to decay to muons.
- This modulation in rate should have ay early period, and is of a large enough fraction to be detected compared to the average flux.



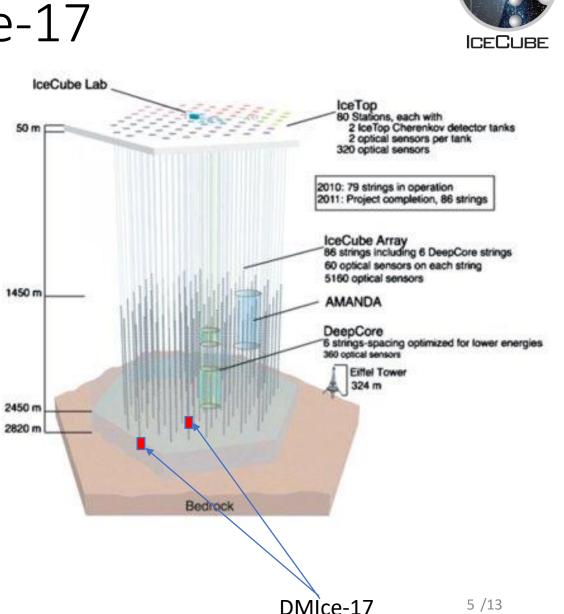






Muons in IceCube / DMIce-17

- IceCube is a neutrino observatory, situated within the Antarctic ice cap.
- DMIce-17 is a pair of modules, each with an 8.5kg scintillation crystal, located at the bottom of two IceCube strings.
- We aim to show the muon modulation using these two detectors

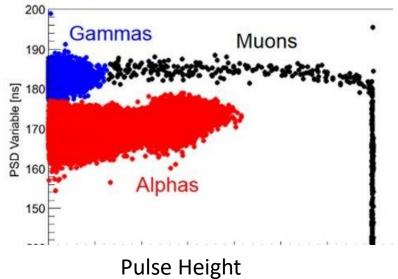




DMIce-17 – an improvised muon detector

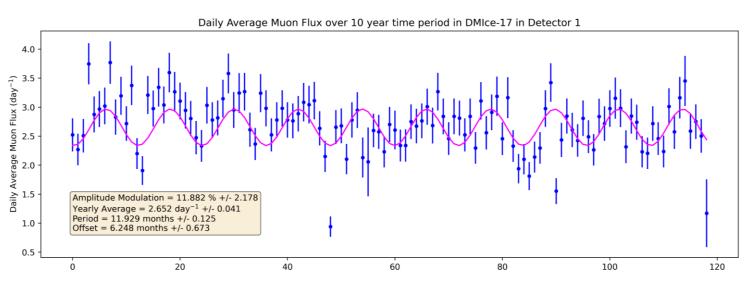


- Implemented to look for direct Dark Matter modulation signal (none seen), repurposed to look for muon signals
- Use phase space of pulse height pulse width to isolate muon population.
- Theoretical calcs show ~3 muons a day
 - Low statistics mean we take a monthly average
 - We have data from each detector covering 2011-2021

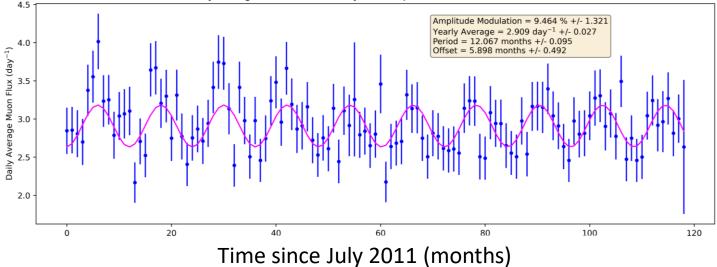




Modulation – Dets 1 and 2



Daily Average Muon Flux over 10 year time period in DMIce-17 in Detector 2





- Period of both consistent with 1 year
- Modulation peak in local Summer, minima in local winter, as expected
- See average of ~3 muons /day/ detector
- Modulation on both ~10%



Link directly to temperature



• Model the correlation between rate and temperature variations with simple linear dependence

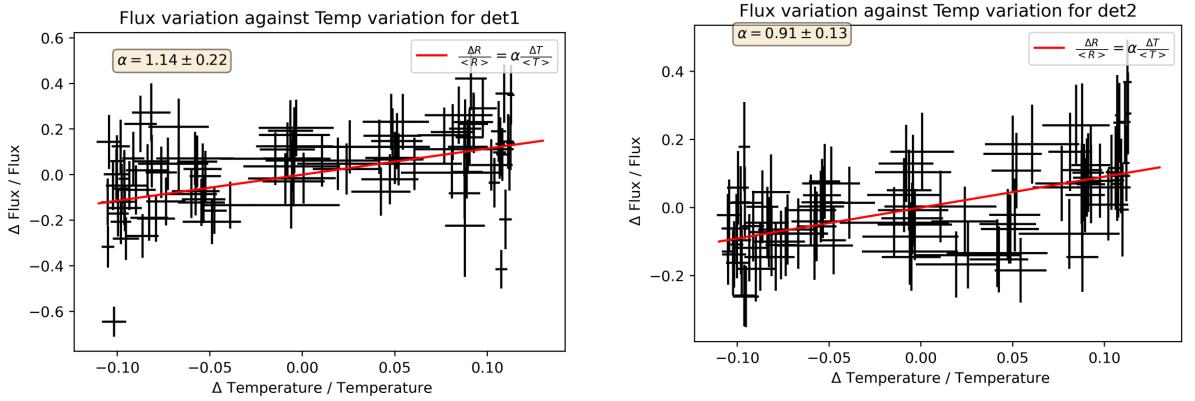
$$\frac{R(t) - \langle R \rangle}{\langle R \rangle} = \alpha \frac{T(t) - \langle T \rangle}{\langle T \rangle}$$

 In this case, T = T_{effective} taken as an effective temperature relevant to muon production throughout the atmosphere





Plotting this variation

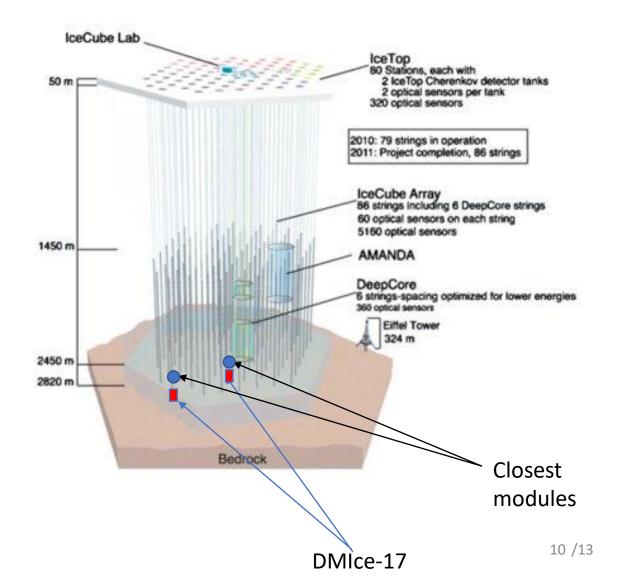


IceCube as a whole sees ~0.8, DMIce may be slightly higher due to preferentially seeing the higher energy muons which modulate more, as lower energy ones are screened by the 2km of ice above





- IceCube modules trigger if they have a local coincidence with a neighbour, or a neighbour of a neighbour.
- This should primarily be due to cosmic ray muons, so if we plot the rate in modules closest to DMIce, they should be similar

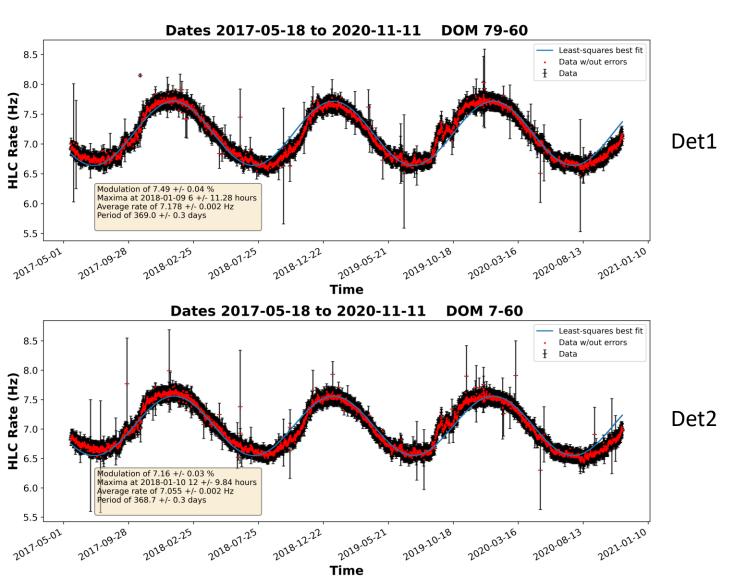






Modules closest to DMIce-17





- Both show lower modulation than the relevant DMIce-17 module
 - Potentially due to a constant background which hasn't been filtered out?
- Period of slightly >365 days?
 - Could be due to smaller range of data
- Phase matches



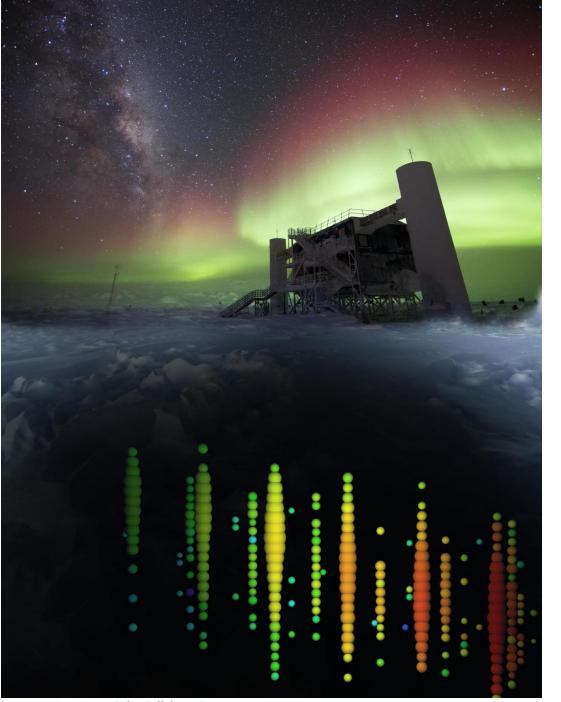
What happens next?



- Closer integration of DMIce-17 with IceCube
- Looking at modulation across the height of IceCube, and potential variation across the width at the lowest height
 - Check for variance in data
- Aim to improve reconstruction of muons found in both DMIce-17 and IceCube



Any questions?



Source: IceCube Collaboration

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