

Supernova Forecast, Echoes of Supermassive Black Holes and Axion Star Explosions

with Dark Matter Experiments



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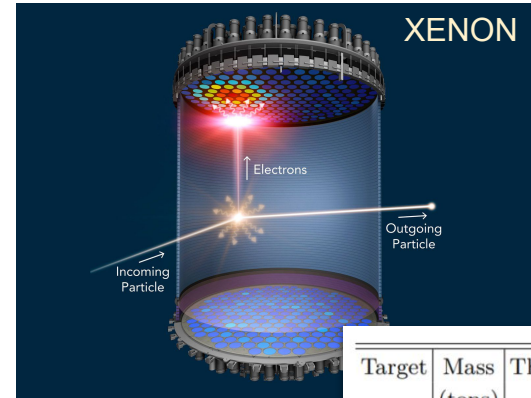
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Direct Dark Matter Detection

- Look for interactions of halo DM in detector → nuclear (and electron) recoils
- Typical setup:
 - heavy target ($A \sim 30-130$)
 - low threshold ($\sim \text{keV}$)
 - potentially scalable (Argon, Xenon...)
- Generations: ton-scale → multi-ton



Target	Mass (tons)	Threshold (keV)	Reference
Ar	300	0.6	ARGO
Xe	50	0.7	DARWIN
Pb	2.4	1.0	RES-NOVA

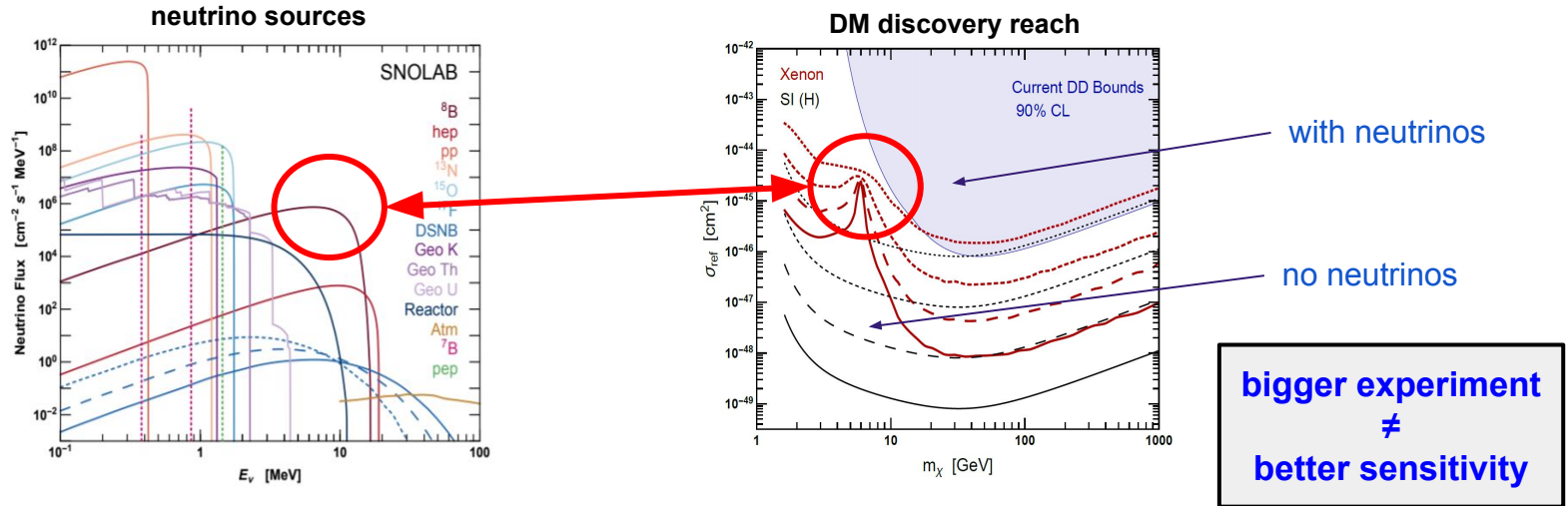
* *detection will allow inferring local DM halo properties*

→ *first application of “halo independent” analysis to electron-scattering*

[Chen, Gelmini, Takhistov, 2021]

Neutrino-DM Connection

- Probing deeper, big DM experiments will encounter contamination from neutrinos (Sun, supernovae, Earth, cosmic rays...) → “*neutrino floor*” via coherent scattering ($\text{CE}\nu\text{NS}$)



- Problem for DM searches, but opportunity for neutrinos

[Gelmini, Takhistov, Witte, 2018]

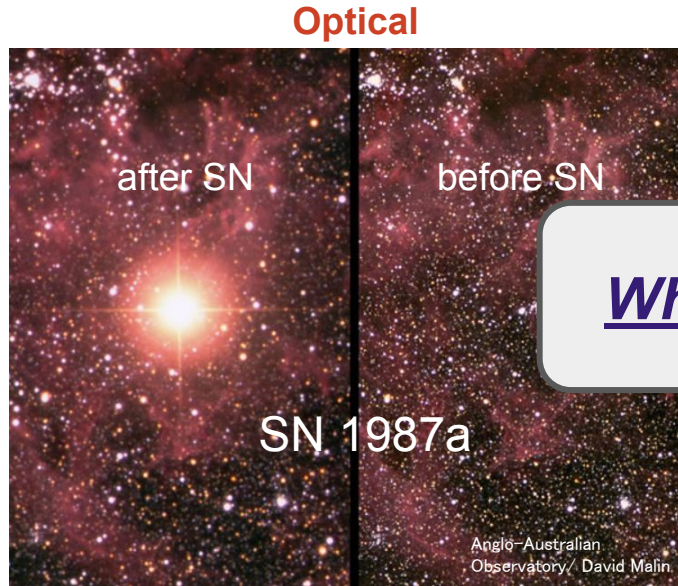
(also [Billard, Strigari, Figueroa-Feliciano, 2013; others])

New Window into Astronomy with Dark Matter Labs

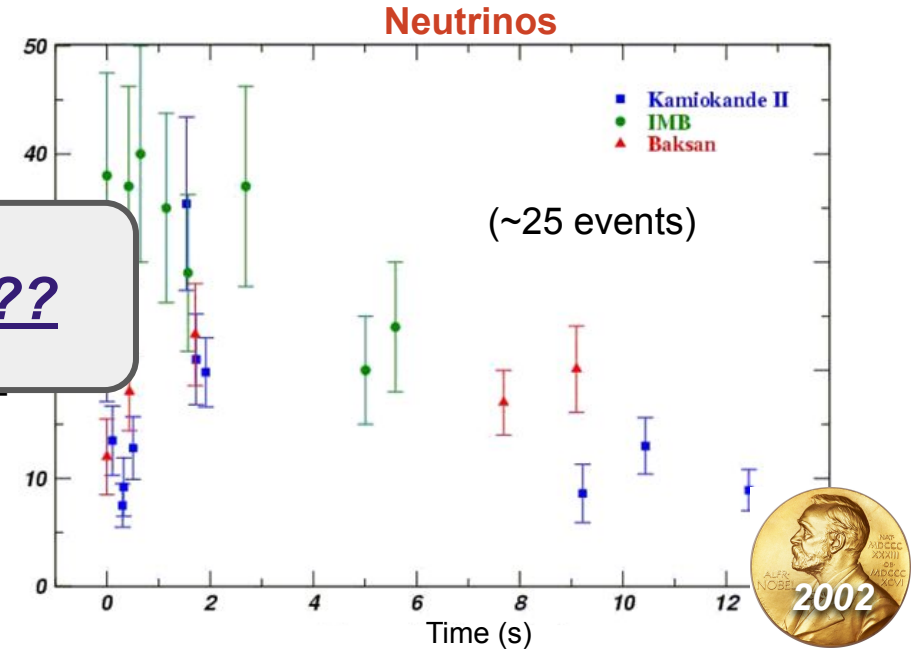
Supernova Forecast

Historic ν -Astronomy Breakthrough: SN 1987a

- Core-collapse SN: most energy released as neutrinos → mechanism confirmed by SN1987a



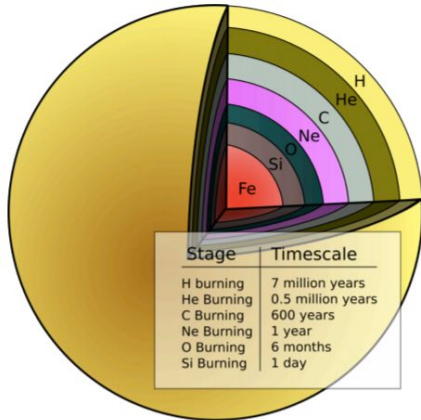
When ???



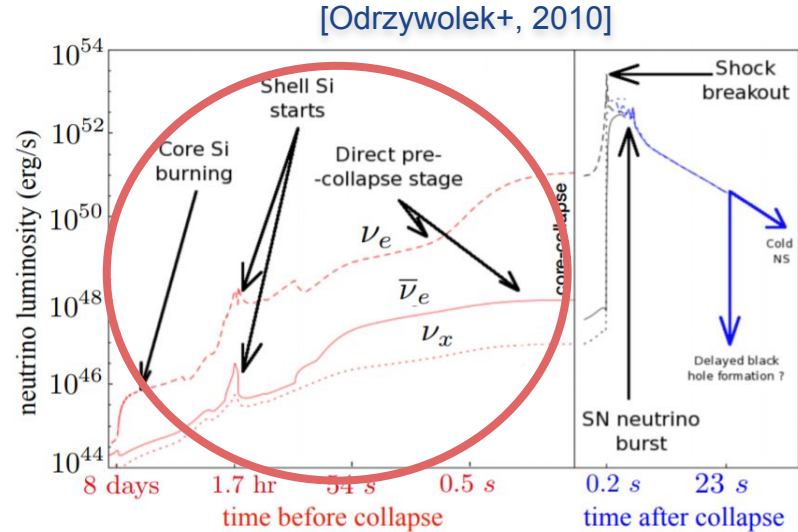
- Many unknowns → *hunt for ν 's from next Galactic SN (rate $\sim 1/30$ yrs) a major target*

Last Stages of Stellar Evolution

- Dying star rapidly changes composition → ↑ density/temperature → ↑ ν -emission



A. C. Phillips, *The Physics of Stars, 2nd Edition* (Wiley, 1999)



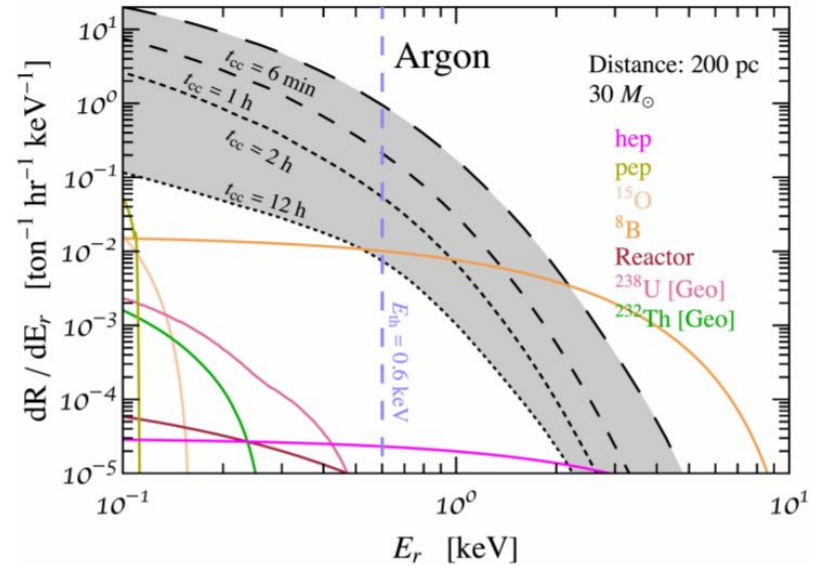
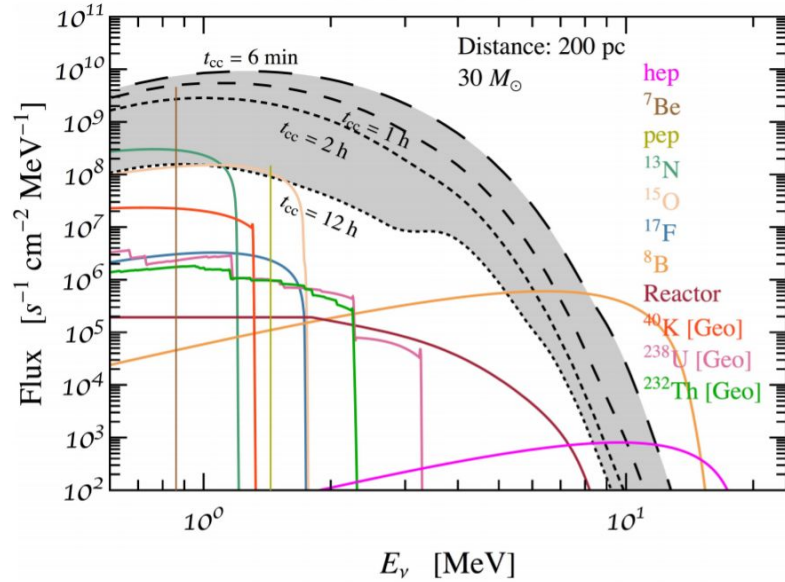
- Pre-SN neutrinos are low-energy (\sim few MeV) → **opportunity for dark matter labs**

* *Super-Kamiokande will see hundreds of events from Betelgeuse before SN*

Supernova Forecast with Pre-Supernova Neutrinos

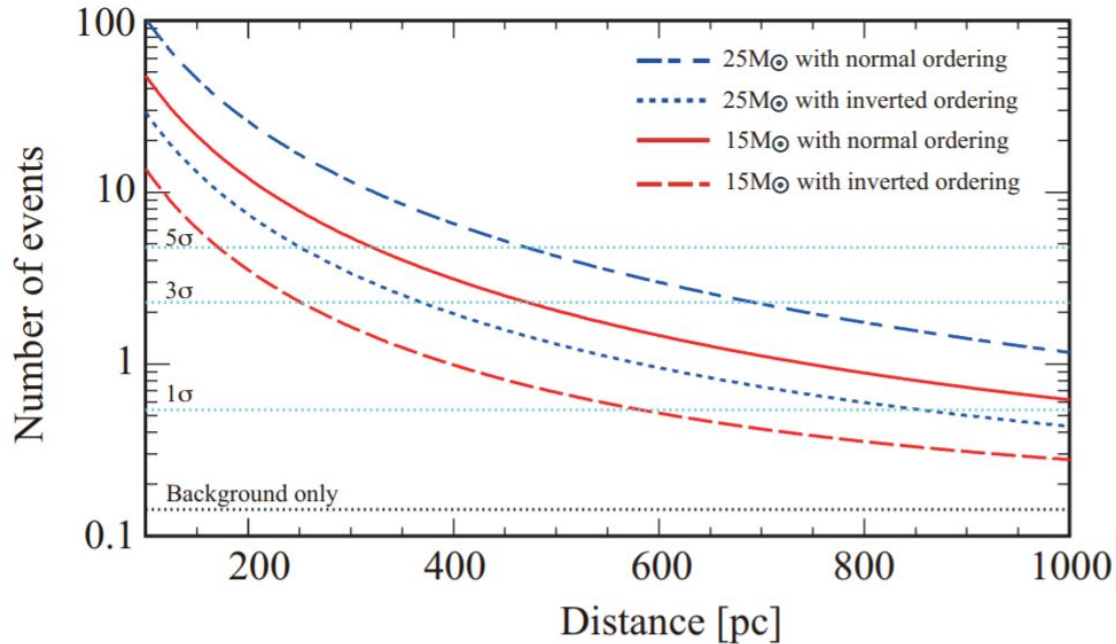
neutrino flux @ Betelgeuse (0.2 kpc)

nuclear recoils



[Raj, VT, Witte, 2019]

Do Not Suffer Oscillation Effects



[Asakura+ (KamLAND),
2015]

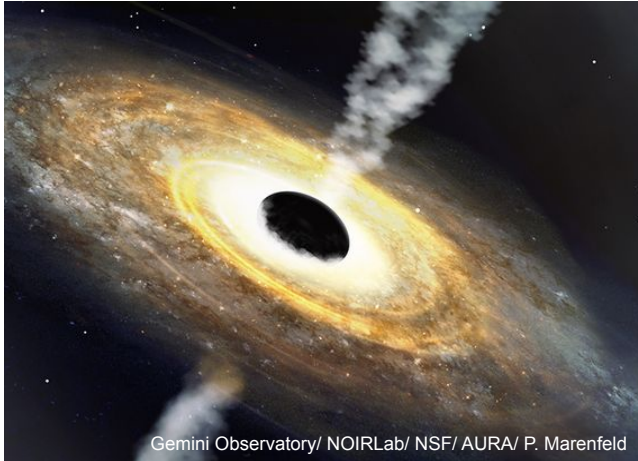
New Window into Astronomy with Dark Matter Labs

Unravelling the Origin of Supermassive Black Holes

Supermassive Black Holes

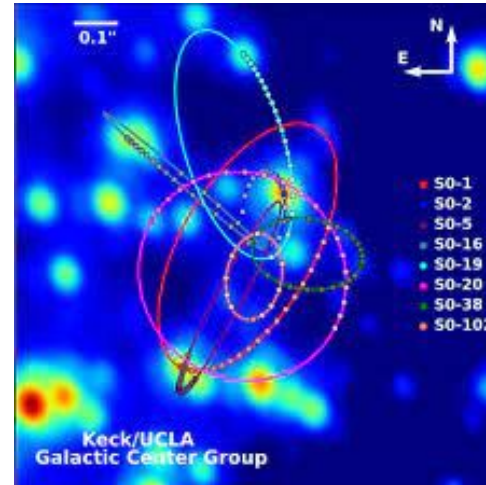
quasars

$$(M_{\text{BH}} \sim 10^9 M_{\odot})$$



galactic centers

$$(M_{\text{BH}} \sim 10^6 M_{\odot})$$



Milky Way

Where do huge BHs come from? → major problem

* *primordial BHs?* [Kusenko, Sasaki, Sugiyama, Takada, **Takhistov**, Vitagliano, *Phys.Rev.Lett.*, 2020]

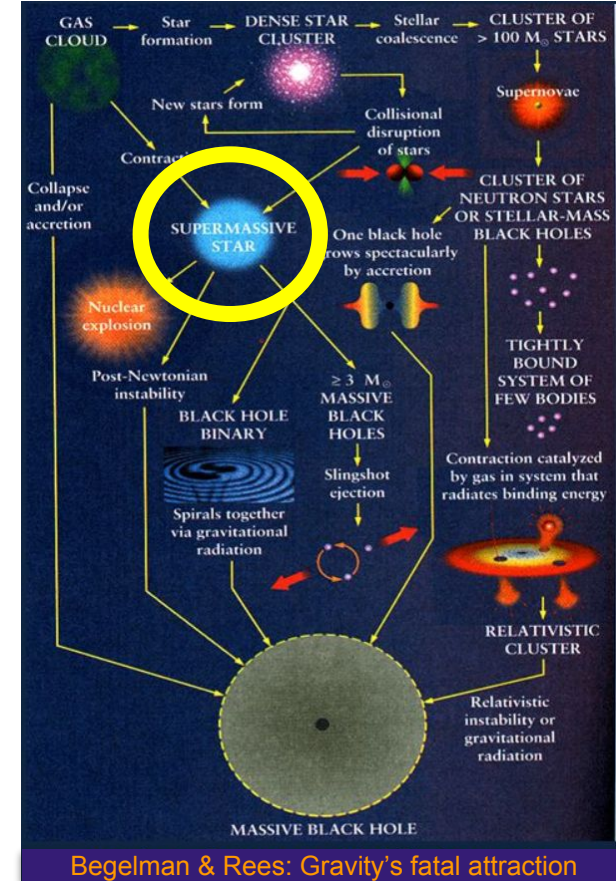
Supermassive Black Holes from Supermassive Stars (SMS)

- Even with vigorous feeding, hard to grow huge BH
→ easy if start with sizable “seed”



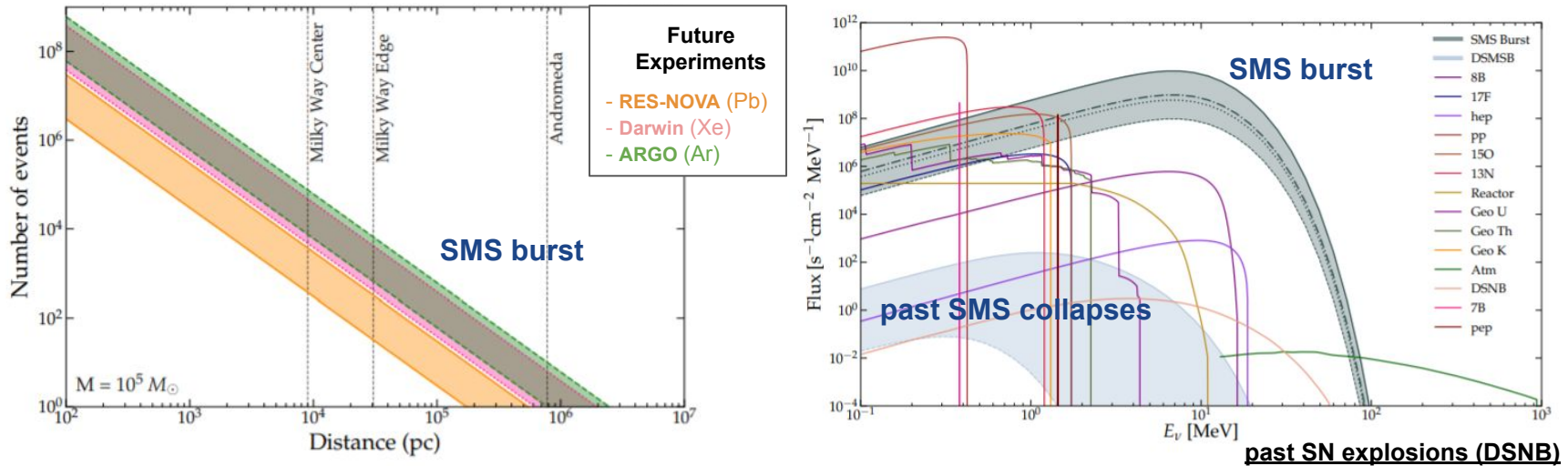
- Pathways predict ($\geq 10^4 M_{\odot}$) supermassive stars

How to test ?



SMS Neutrinos in Dark Quantum Labs

- SMS collapse releases enormous total energy in neutrinos ($\sim 10^4$ x supernova) [Shi, Fuller, 1998]
- Individual neutrinos low energy \rightarrow big DM experiments can easily detect



- Neutrinos from past SMS collapses pose novel contamination source for DM !

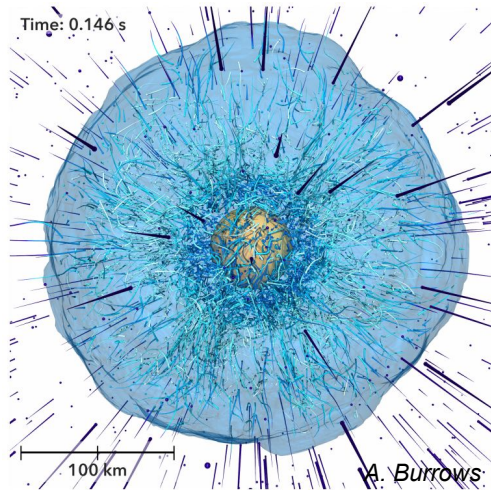
[Munoz, Takhistov, Witte, Fuller, 2021]

Neutrinos are Wonderful Teachers ...

Axion Astronomy, a Neutrino Analogy



- Axions well motivated (CP, fundamental theory...see [Ellis+, 2021] review)
→ usual experiments focus on non-relativistic axion DM
- **Novel direction: axion astronomy with transients**... *take lessons from neutrinos*



Neutrino Burst
(transients: SN, SMS,
mergers...)



Axion Burst
(transients: axion star explosion, BH
bombs, mergers...)

**Diffuse Neutrino
Background**
(historic transients)

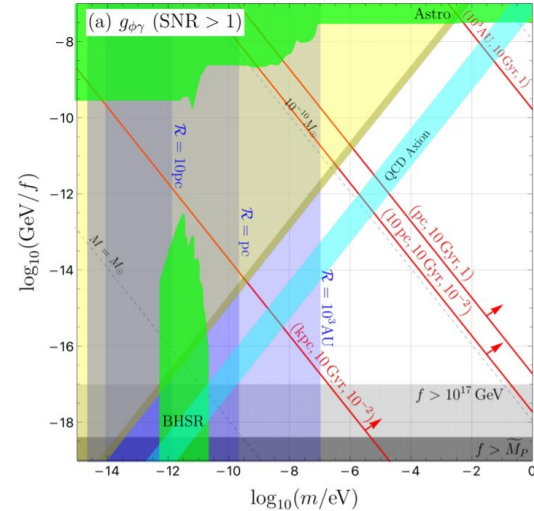
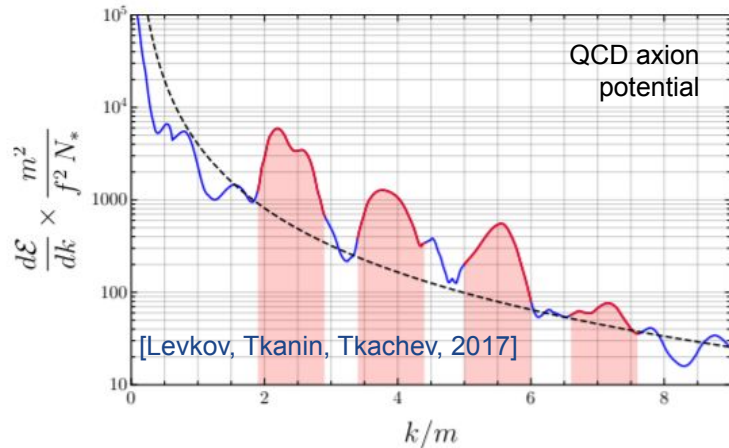


**Diffuse Axion
Background**
(historic transients)

[Eby, Shirai, Stadnik, Takhistov, 2021]

Axion Star Explosions

- Axion stars can arise within e.g. minicluster “clouds” [Kolb, Tkachev, 1993]
- With self-interactions, collapsing stars burst in relativistic axions (bosenova)



- **New signatures, could lead to insights into the fundamental axion potential !**

[Eby, Shirai, Stadnik, Takhistov, 2021] * axion star collapse also [Fairbairn+, 2019]
 ...microlensing signals [Fairbairn+, 2017]

Concluding Remarks

- Tackling **Big multi-faceted questions** requires broad perspective and exploiting connections between historically distinct fields (cosmology, particle physics, astrophysics...)
- Interplay and complementarity between dark matter and neutrino experiments

→ ***Neutrinos serve as a guiding ghostly light for discoveries in multi-messenger era !***

