

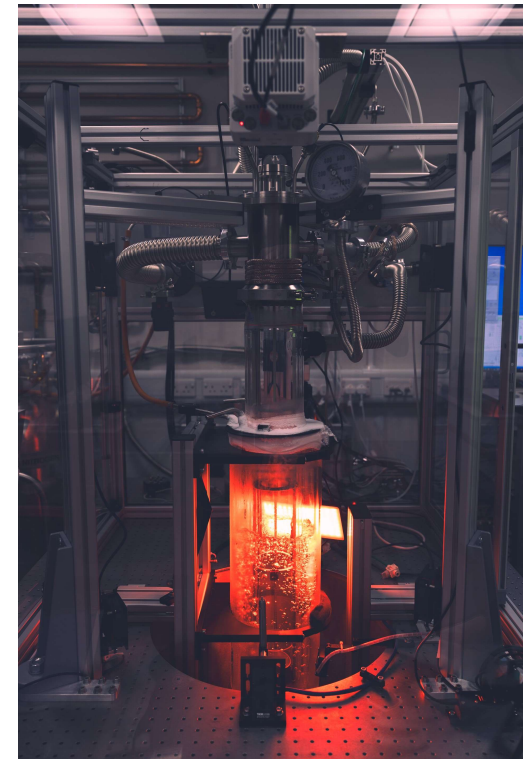
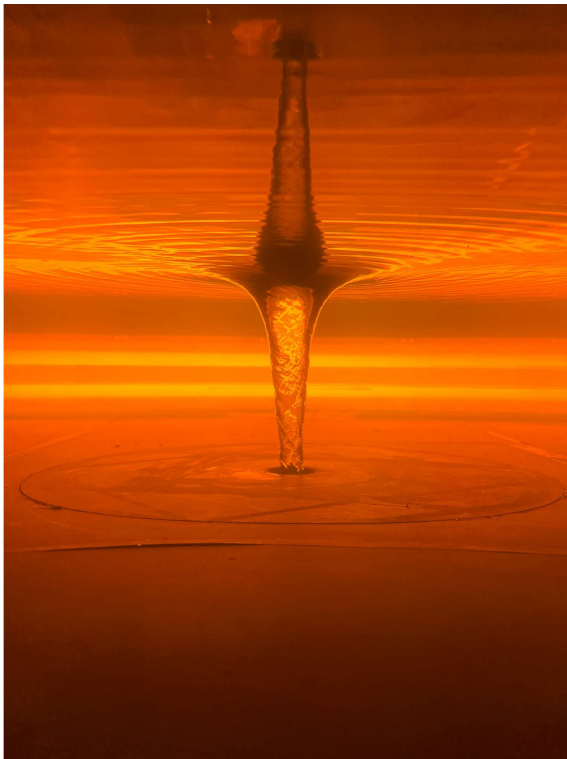
Rotating black holes in the lab

Sam Patrick



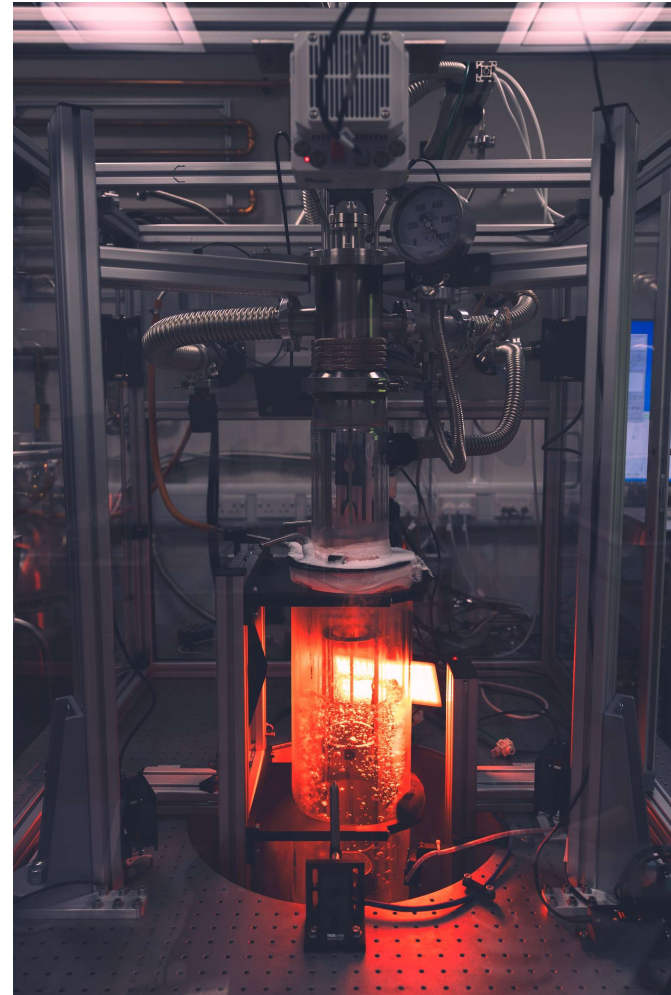
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 **QSimFP**



Overview

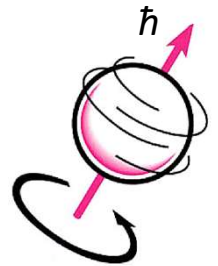
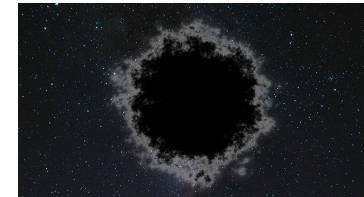
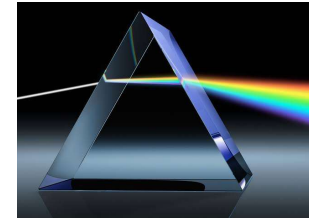
1. Analogue spacetimes
2. Black hole ringdown
3. Superradiance
4. Summary



Motivation

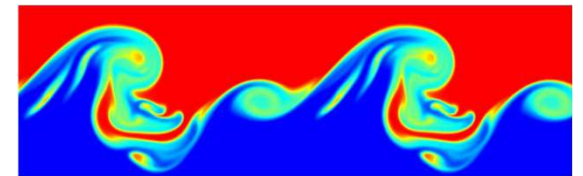
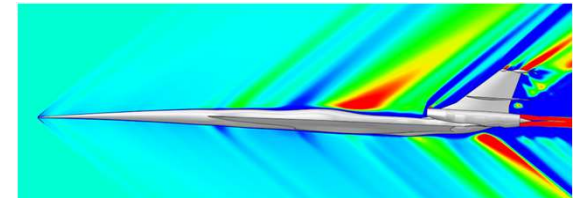
1. Study possible signatures of quantum gravity, for example...

- Lorentz violations (string theory, LQG, modified gravity)
- Modified boundary conditions (echoes, fuzzballs, compact objects)
- Quantised charges (black hole mass/area, angular momentum)



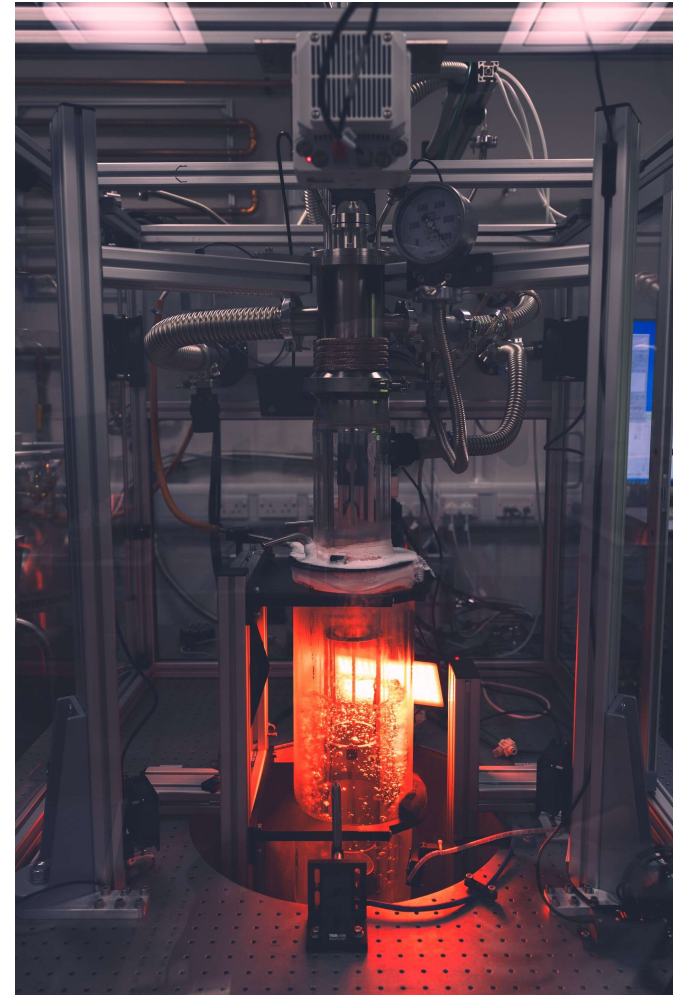
2. Use gravity analogue to study fluid behaviour, for example...

- Stability of supersonic flows (relation to horizon physics)
- Vortex dynamics (connected to superradiance)
- Import techniques (and insights) from gravity



Overview

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Analogue spacetimes

A massless scalar field ϕ in curved spacetime obeys:

$$\square\phi \equiv \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} g^{\mu\nu} \partial_\nu \phi) = 0$$

$\phi = \text{waves}$, $g_{\mu\nu} = \text{metric fixed by background}$

1. Surface waves in water



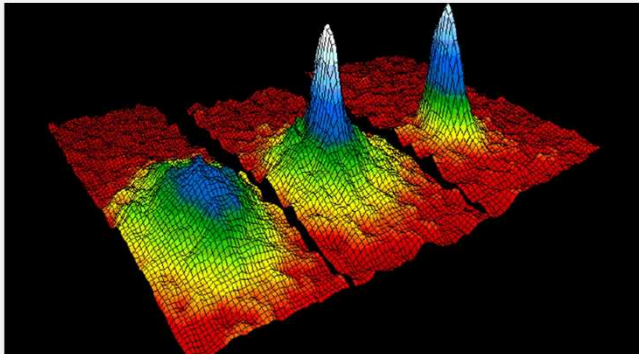
$$\delta v = \nabla \phi$$

$$(\partial_t + v \cdot \nabla)^2 \phi - c^2 \nabla^2 \phi = 0$$

$c = \sqrt{gh}$ wave speed

$$g_{\mu\nu} = \begin{bmatrix} -(c^2 - v^2) & -\vec{v} \\ -\vec{v} & \vec{1} \end{bmatrix}$$

2. Sound waves in BECs



$$\text{GPE) } i\hbar \partial_t \Psi = -\frac{\hbar^2}{2M} \nabla^2 \Psi + \lambda |\Psi|^2 \Psi$$

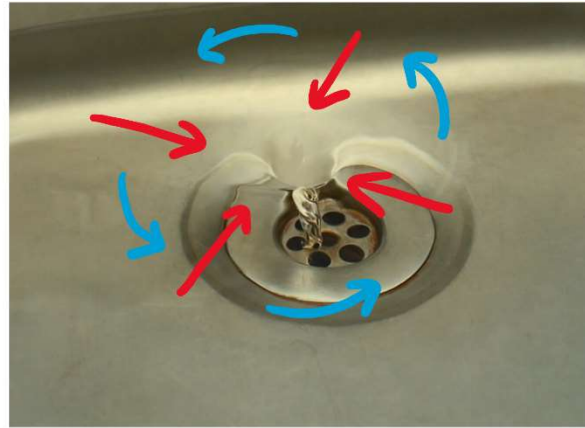
$$\text{Fluid) } \Psi = \sqrt{n} e^{iS - i\mu t/\hbar}, \quad \vec{v} = \frac{\hbar}{M} \nabla S$$

$$\partial_t v + v \cdot \nabla v + \frac{\nabla P}{\rho} = \nabla \left(\frac{\hbar^2}{2M} \frac{\nabla^2 \sqrt{n}}{\sqrt{n}} \right)$$

$$\partial_t n + \nabla \cdot (nv) = 0, \quad P = \frac{1}{2} n^2$$

$$\text{Sand) } n \rightarrow n_0 + \delta n, \quad S \rightarrow S_0 + \frac{M}{\hbar} \phi, \quad c = \sqrt{\frac{\lambda n_0}{M}}$$

Rotating black hole analogues



$$\vec{V} = v_r \hat{e}_r + v_\theta \hat{e}_\theta$$

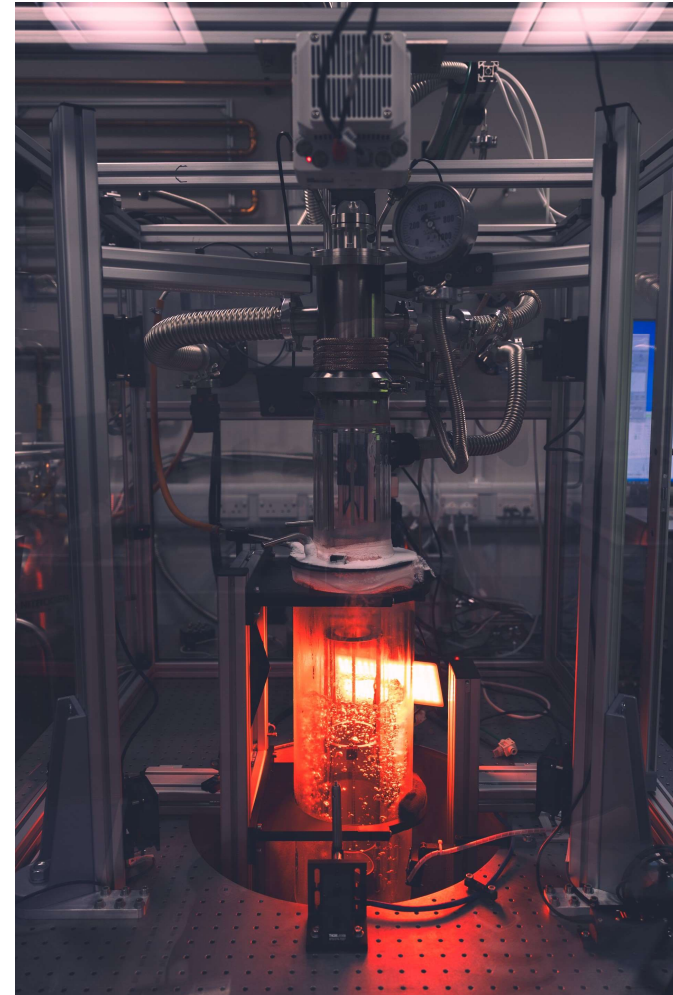
$$v_r = -D/r \quad \text{draining flow}$$

$$v_\theta = C/r \quad \text{circulating flow}$$

$$ds^2 = -c^2 dt^2 + \left(dr + \frac{D}{r} dt \right)^2 + \left(r d\theta - \frac{C}{r} dt \right)^2 \quad \text{horizon: } r_h = D/C \quad \text{ergosphere: } r_e = \frac{\sqrt{C^2 + D^2}}{c}$$

Overview

1. Analogue spacetimes
- 2. Black hole ringdown**
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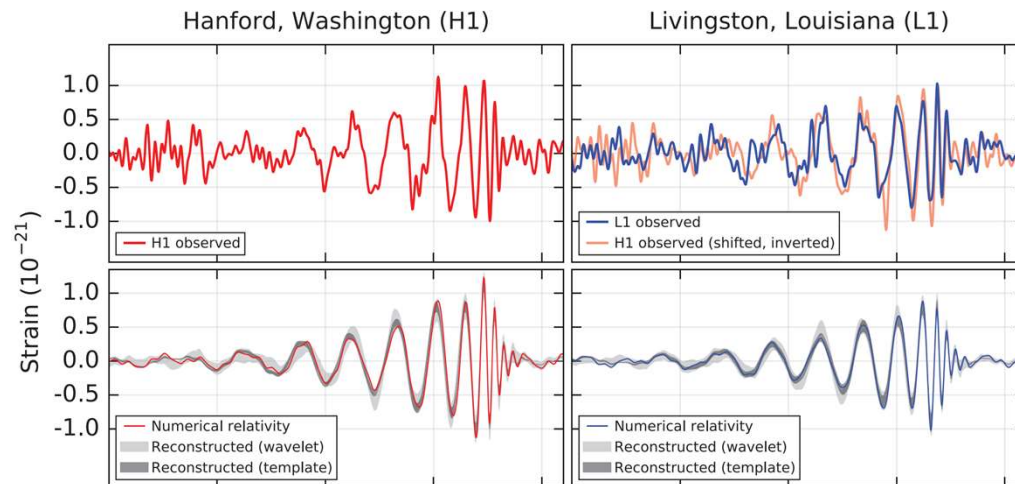
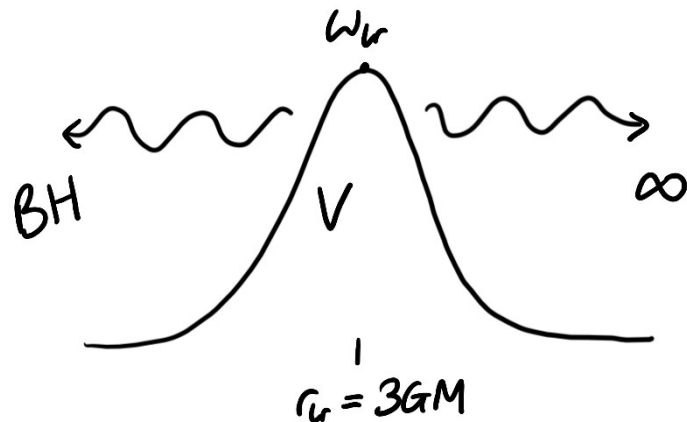


Black hole ringdown

- Effect in linear field theory which probes strong field regime.
- Measured by LIGO/VIRGO

e.g. for Schwarzschild

$$-\partial_x^2 \psi + (-\omega^2 + V)\psi = 0$$



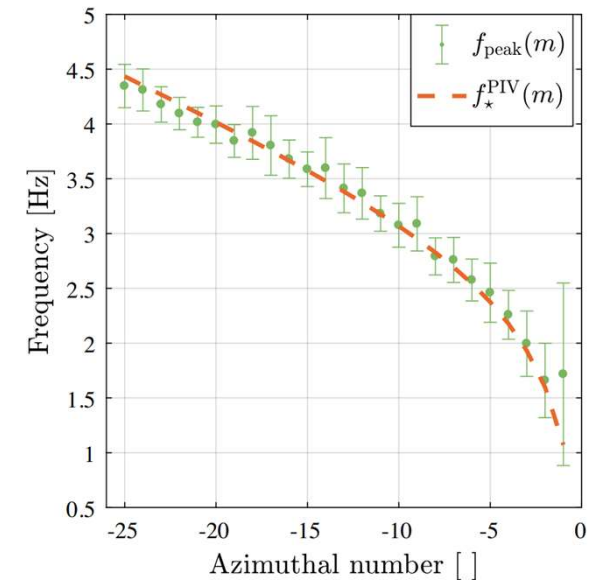
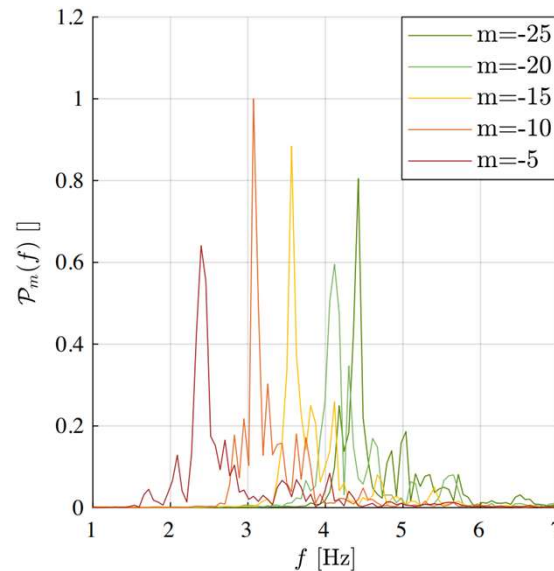
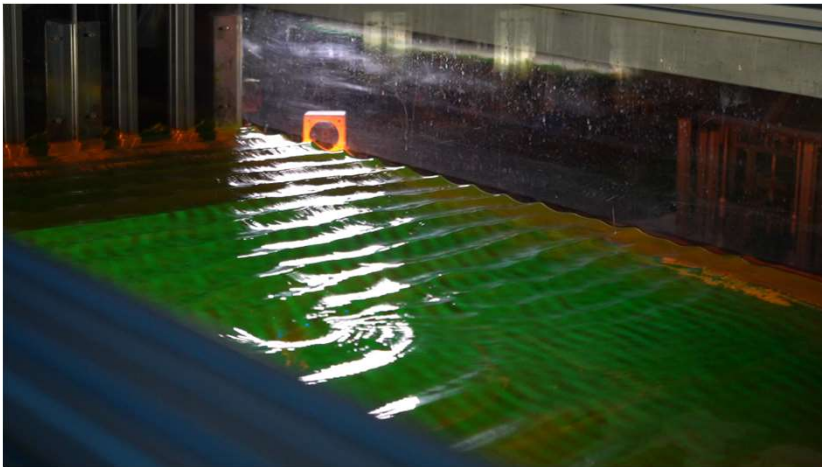
Phys. Rev. Lett. **116** 061102 (2016)

$$\omega_{QNM} \simeq \omega_r - i \frac{\sqrt{-2V_r''}}{2\omega_r} \left(n + \frac{1}{2}\right)$$

↑ overtones $n=0,1,\dots$

Ringdown experiments (2017)

Phys. Rev. Lett. **125** 011301 (2020)



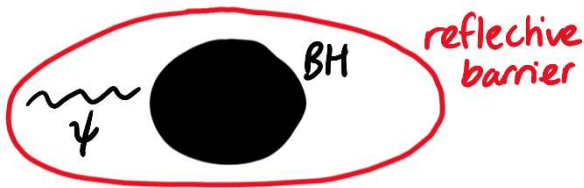
- Boundary reflections obscure ringdown signal

- Each m -mode has distinct spectral peak in the noise

- These peaks are located at the light-ring frequency

TAKEAWAY: suggests “*finite-size + noise = QNMs*” ?!

Finite-size effects



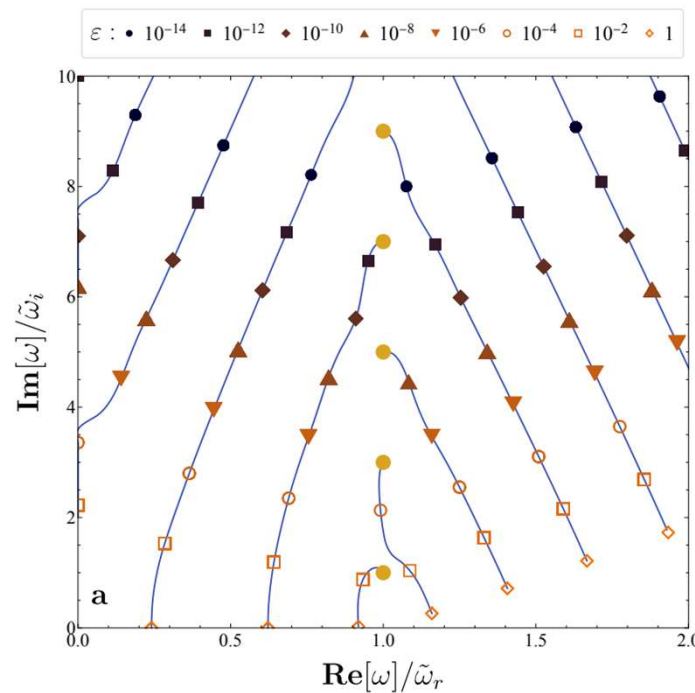
$$-\partial_x^2 \psi + (-\omega^2 + V)\psi = 0$$

Pöschl-Teller model

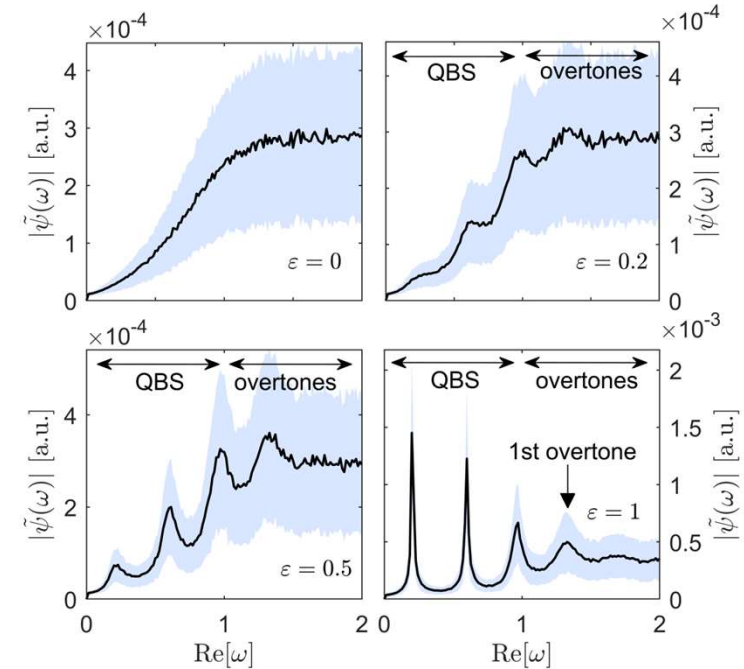
$$V = \text{sech}^2 \alpha x$$



- Varying ε determines a **spectral flow**



- For large ε , QNMs emerge as **peaks in the noise**

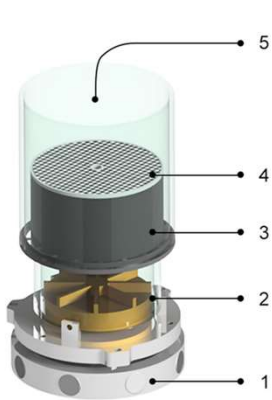


TAKEAWAY: finite size effects *increase overtone lifetime* and *enable QNM excitation by noise*

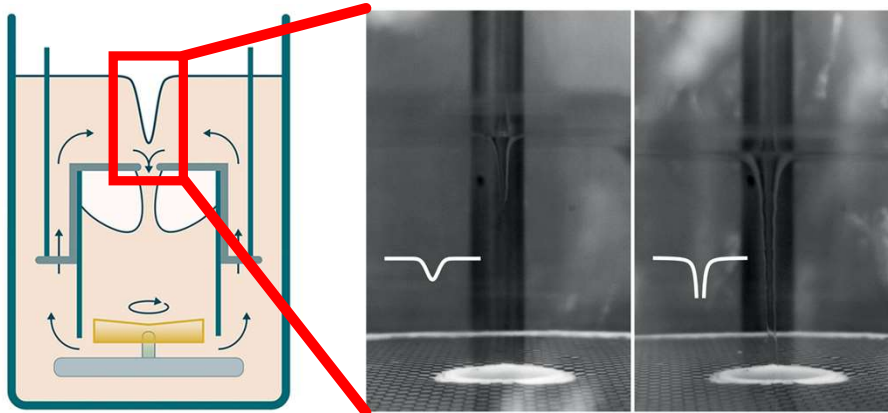
Based on work in
arXiv:2502.1109

Superfluid experiments (2025)

➤ Experimental schematic



➤ Side profile

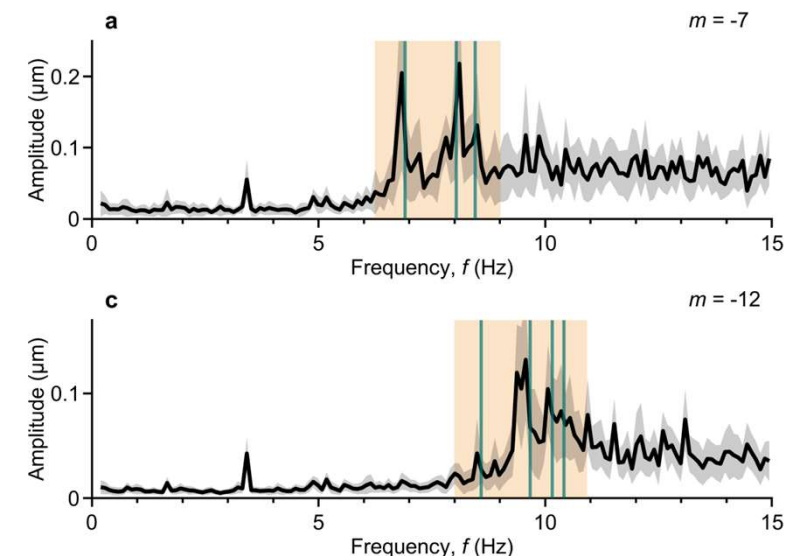


Nature **628** 66 (2024)

TAKEAWAY

- Proof of principle that noise in finite size analogues can be used to measure overtones
- Use to investigate (a) spectral stability, (b) overtones vs. nonlinearities

➤ Results for $m = -7, -12$

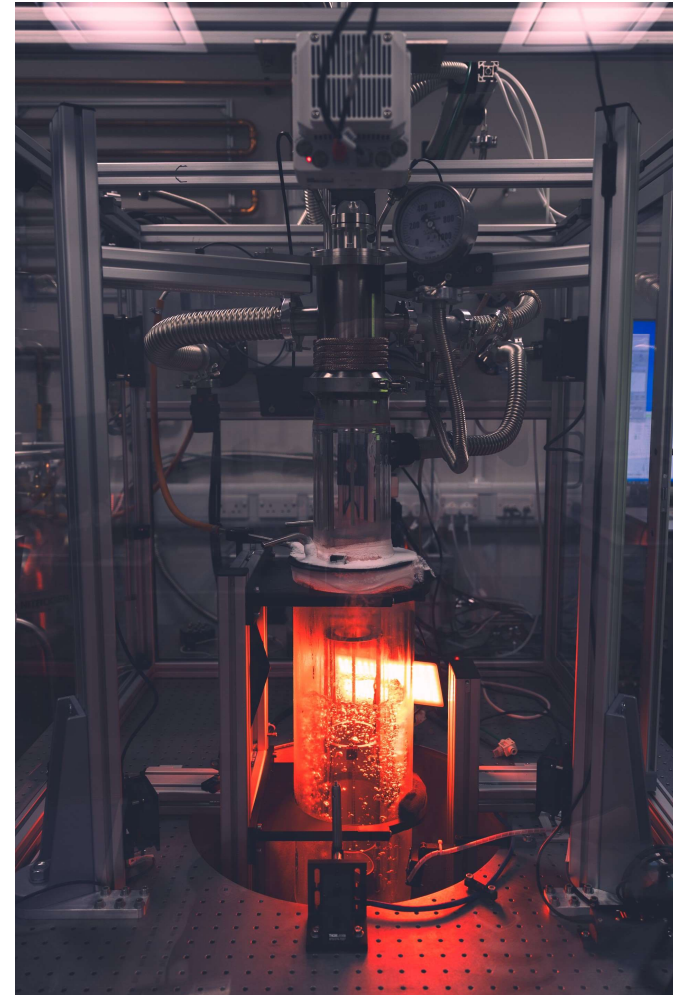


Consistent with...

- Reflecting BCs at the wall
- Absorbing BCs in the core

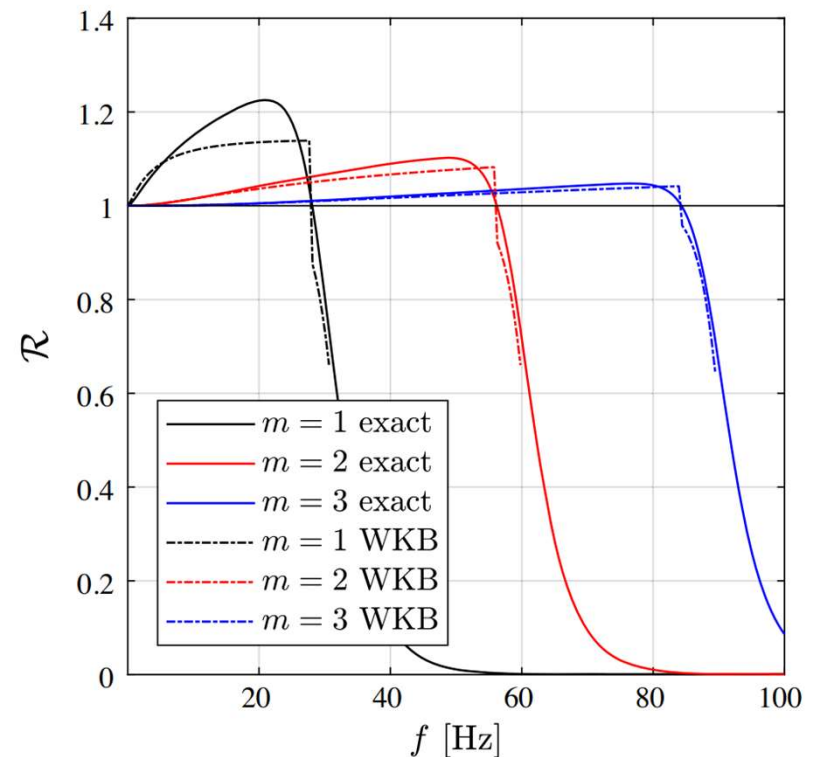
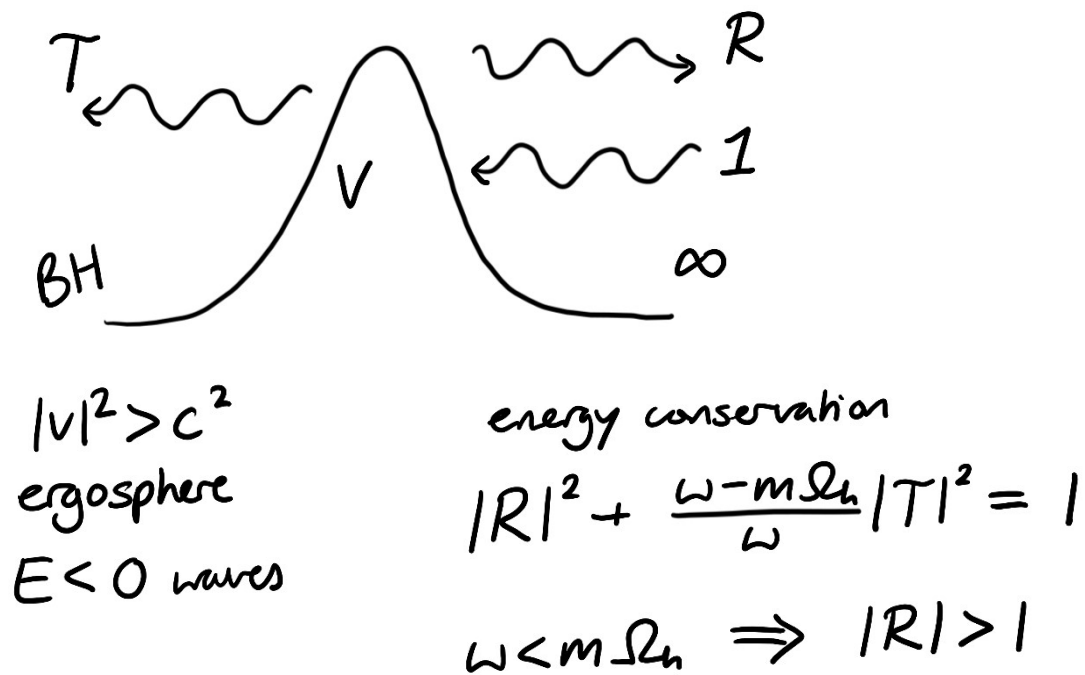
Overview

1. Analogue spacetimes
2. Black hole ringdown
- 3. Superradiance**
4. Conclusion

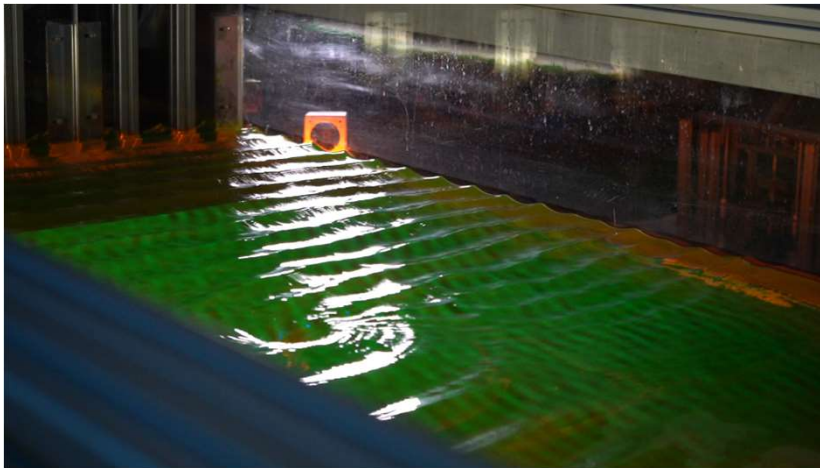


Rotational superradiance

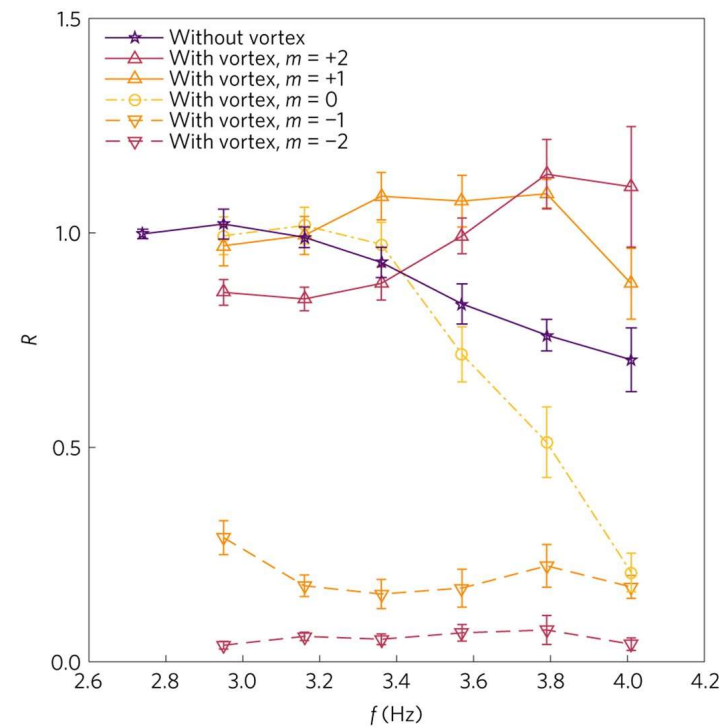
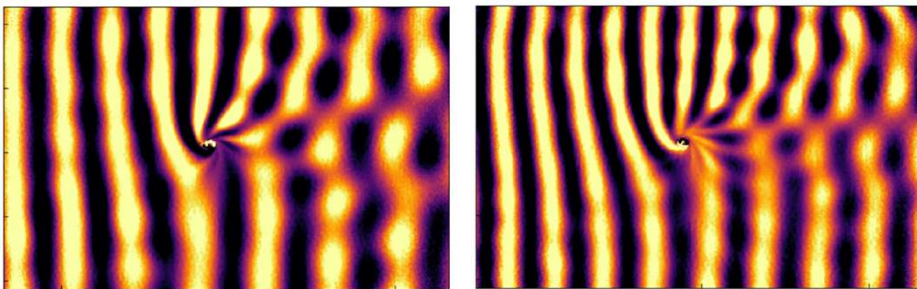
- Amplified wave extracts rotational energy from system
- Allows rotating BHs to shed their angular momentum



Superradiance experiments (2016)



➤ Monochromatic wave of different frequencies



➤ Amplification of corotating waves

Superradiance with LV

$$(\omega - v \cdot k)^2 = F(k)$$

$$F_{\text{surf}}(k) = (gk + \frac{\sigma}{\rho} k^3) \tanh(hk)$$

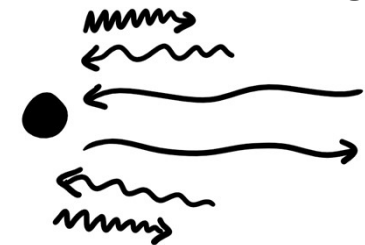
$$F_{\text{BEC}}(k) = c^2(k^2 + \xi^2 k^4)$$

➤ Relativistic at low k



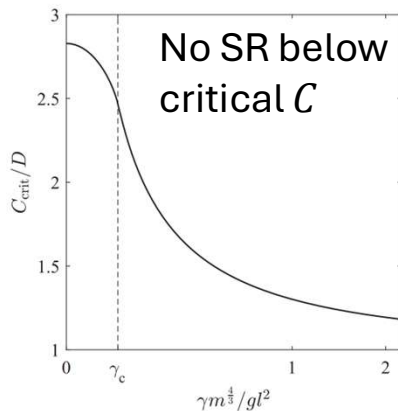
- How do extra modes scatter?
- When does core absorb negative energy?

➤ More modes at high k

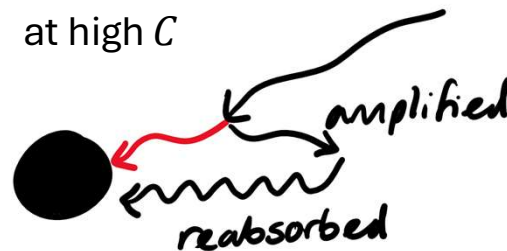


Surface waves

$$0 < \omega < \left(\frac{g^2}{D}\right)^{\frac{1}{3}} \frac{m^{1/3}}{2^{2/3} 3} \left(\frac{C}{D} - 2^{3/2}\right)$$

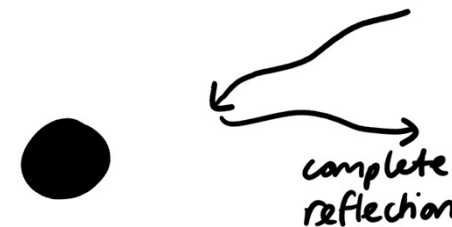


SR modes
reabsorbed
at high C



Sound waves (BEC)

$$0 < \omega < m\Omega_h/1 + m\xi^2/r_h^2$$



No SR for
high m
and low C

TAKEAWAYS

- LV alters SR threshold
- LV can stop SR for certain modes

Black hole bomb

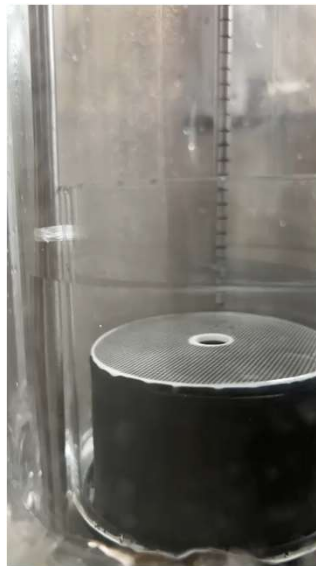
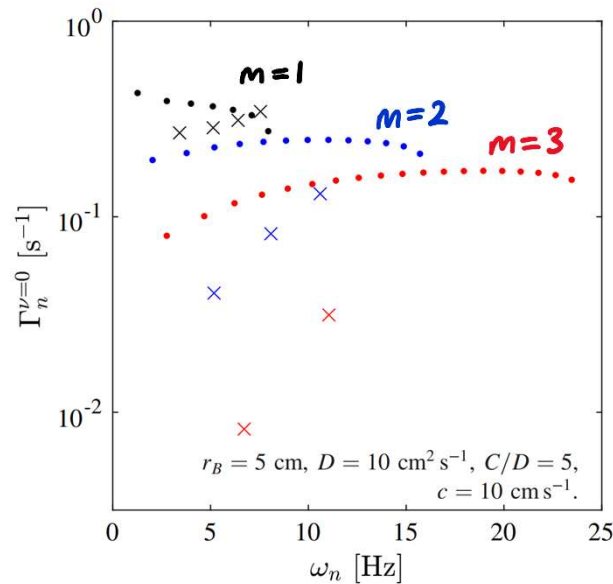


- Black hole bomb (Press & Teukolsky)
- Occurs for ultra-light scalars (e.g. axions)

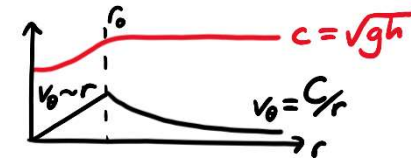
$$S(\omega_n) = \pi(n + \frac{1}{4})$$

$$\Gamma_n = \frac{\log|R|}{T} - \langle \nu k^2 \rangle_T$$

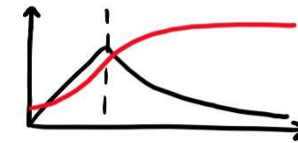
Phys. Rev. D **110** 124068 (2024)



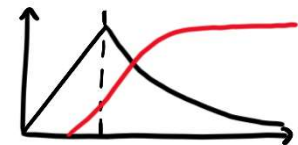
Vorticity leads to extra degrees of freedom!



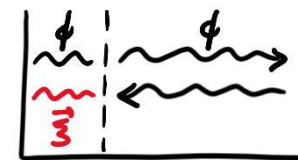
➤ Core with $\nabla \times v \neq 0$



➤ Ergoregion intermediate C



➤ No rotational core for large C



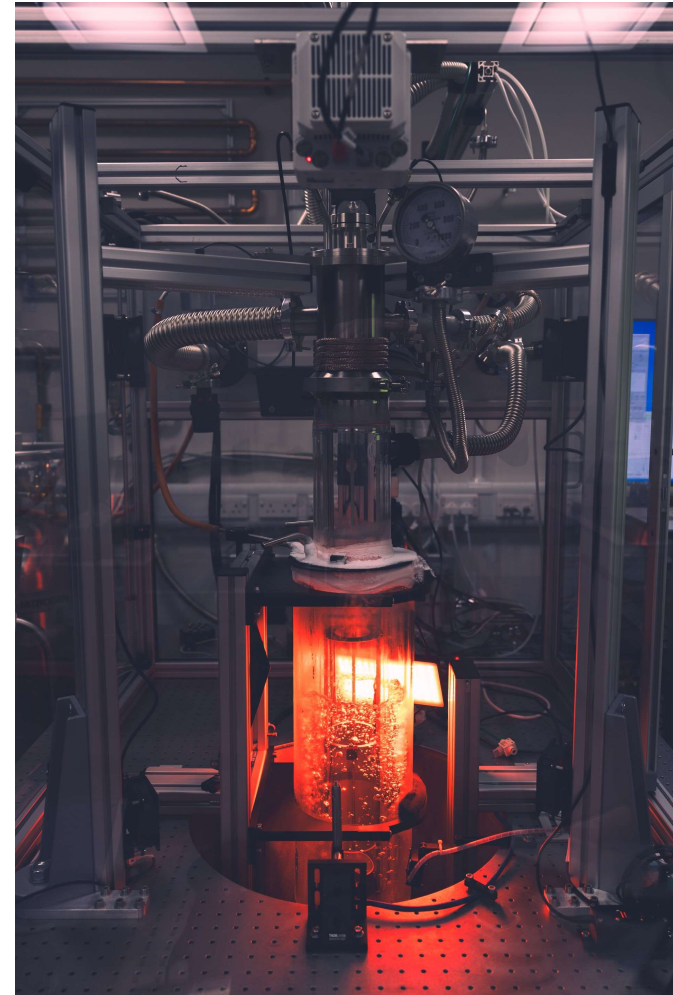
➤ ϕ has $E < 0$ in ergoregion
➤ ξ has $E < 0$ with no ergoregion

TAKEAWAYS:

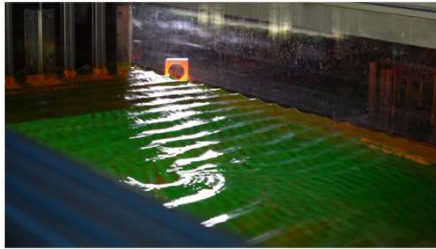
- Amplification + confinement = instability!
- BHB dominates a rapid rotations

Overview

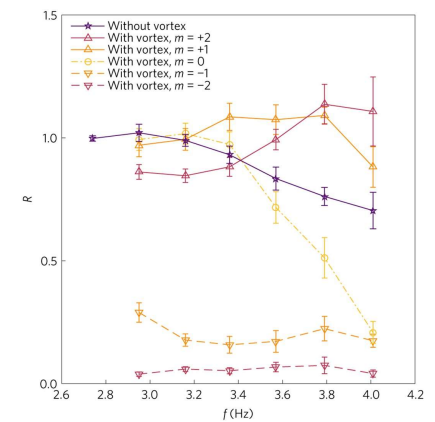
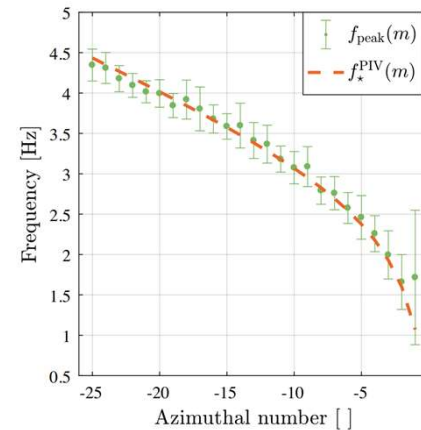
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Summary



- Waves see draining vortex as a rotating BH
- Measured QNMs and superradiance in water



- Finite size + noise $\rightarrow$ QNM overtones
- Finite size + superradiance $\rightarrow$ BH bomb instability
- LV effects lead to spectral modifications of QNMs and SR (can suppress both)
- Next: study effect of quantised angular momentum.

